



## Treating HIFI Artifacts

Flags and Spurs



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- Artifacts that require attention
- Review of flag types relevant to data processing
- Visualization of flags within HIPE
- Practical example:
  - Spectral scan deconvolution failure
  - Manipulation of flags via scripts
  - Successful deconvolution of cleaned data









- **Spurs** (above) are features which are generally narrow (<50MHz) and Gaussian in shape and. They can look like astronomical lines, and have the potential to 'fool' users.
- Saturations and underpumped mixers (causing ringing) are more obvious (see right)





- Strong spurs, such as the (now fixed) spur in 1a, can knock out an entire dataset.
- Weak spurs can impact only a very narrow region (4b for example)



- Since spurs can appear as lines, we only look for them in the loads, and then propagate the flags to the science data.
- They can move in IF over time. If an observation goes on a long time without taking a load observation, the spur can 'migrate' out of the flagged region and complicate data reduction (most notably for spectral scans).
- Spurs can be missed by the algorithm if the load is not as 'smooth' as expected.
- Spurs generally have a predictable position in the IF as a function of the LO frequency.





## Spur Reporting



## **Historical Spur Table** 000 Frequency Editor **Frequency Editor** LO Frequency 952.225 Lower Sideband Upper Sideband HRS 1 IF 6.00 HRS 2 IF 6.00 WBS ON 946.0GHz 948.0GHz 950.0GHz 952.0GHz 954.0GHz 956.0GHz 958.0GHz 960.0C **Pipeline generated Spur Table** $\bigcirc$ 866 876 886 916 926 936 946 896 906 Local Oscillator Frequency (GHz) 000 obs Redshift Reset 💁 obs 🗙 Redshift -0.000055 Reset all frequencies Spur Parameters Summary **Frequency Selection** Object: LDN1157-B1 Instrument: HIFI Type Transition ... Line Up... RA: 20h 39m 9.74s DEC: 68° 2' 36.45" ÷ 🗸 WBS -No Lines- -No Lines-Observation ID: 1342181161 **Operational Day: 80** $\checkmark$ + \$ ☑ HRS 1 -No Lines--No Lines-Observation Mode: HifiSScanModeDBS 1 $\checkmark$ HRS 2 + -No Lines--No Lines-Meta Data 🔻 Data 🗁 obs obs.refs["trendAnalysis"].product.refs["SpurTable"].product.refs["WBS-H"].product["spur"] 🕂 🎥 History LO [MHz] IF [MHz] Index obsid hcid apid Band subband Pixel auxiliary Warning messages 0 1342181... 131 1030 582025.5 4 7374.0 7752.933 1b calibration Strong spur in WBS sub-band: 1 1342181... 135 1030 1b 582990.75 4 7199.0 7652.746 1 🖻 level0 Strong spur in WBS sub-band: 1,2,3 1342181... 139 1030 584082.0 4 6980.0 7526.807 2 1b level0\_5 3 1342181... 143 1030 1b 585029.25 4 6736.0 7386.814 level1 4 1342181... 147 1030 1b 586061.9...4 6523.0 7263.763 ievel2 彦 logObsContext 🔊 quality 🗁 trendAnalysis - 🃂 CombTrend 🤌 FpuTrendProduct LoTrendProduct - 🗁 SpurTable 🔶 🧁 WBS-H spur 🗄 🎾 History 🖶 🥟 WBS-V 🔊 TMpage 🗄 🥭 Tsvs Waterloo NASA Herschel Science - page 5





- The pipeline and post-pipeline tools rely mainly on two types of flags
- ROWFLAGS apply to entire datasets
- CHANNEL FLAGS (aka pixel flags) apply to individual data points
- The pipeline sets both types of flags, and while it generally does a good job, it is sometimes necessary to manually intervene



## Flag presentation in contexts



















 For the SUSPECT\_LO flag, there is additional piece of information in the 'badLo' array which identifies which subband is affected.





- flagTool: built into HIPE, and useful for setting channel flags.
- Spectrum Explorer. Can operate on datasets but only when detached from the context
- Command line tools based on scripts. Some are packaged with the following:
- HifiFlagVisualizer: available as a plugin https://nhscsci.ipac.caltech.edu/sc/index.php/Hifi/HIPI

