

Introduction to HIPE & HIPE Help

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HIFI+Herschel Data Processing Editorial Board







Welcome to HIPE!



What is HIPE?

Herschel Interactive Processing Environment





- page 3



What is HIPE?

- ٥ HIPE 9.1.0 File Edit Run Pipelines Scripts Window Tools Help 🏠 🥥 🔍 🖆 🦏 Welcome _ 0 Welcome to Herschel Interactive Processing Environment! HIPE Version: 9.1.0 RC2 Hover your mouse over one of the images below for more information. Workbench Access Data Documentation Preferences Updates External Tools Tip: To get back to this page, click the 鄰 icon in the toolbar. Welcome 271 of 4444 MB 6
- HIPE is the software used by
 engineers, calibration scientists and
 astronomers to reduce, visualize and
 analyze Herschel data of the PACS,
 SPIRE and HIFI instruments:
 - Interactively
 - Automatically: jython scripting
 - In the background, without user interface: "jylaunch"
- For a high level overview of HIPE development, see the presentation by Stephan Ott this morning.





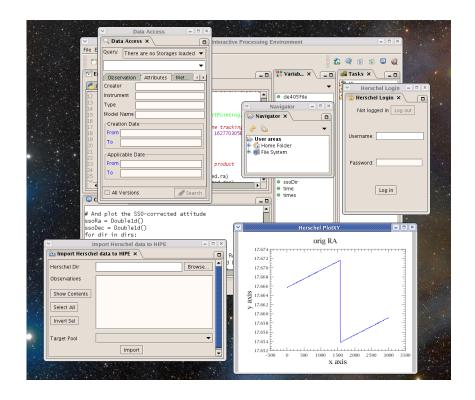






An integrated graphical environment

- HIPE brings several applications together under a common, consistent interface. From data retrieval to plotting, from image analysis to scripting, powerful utilities are one click away.
- From raw data fresh off the Herschel Archive to publicationready plots, all you need to get science out of your observations.









The power of Java and Jython

- HIPE is based on Java, one of the most popular programming languages. The multi-platform nature of Java allows HIPE to work flawlessly under Window, Mac OS and many Linux and UNIX flavours.
- Jython is the Java-based version of Python, used worldwide for quick development of complex applications.

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r = mpsDir + "horizons"
merides = Ephemerides(orbitFile, de405File)
d = 2000018 # Melpomene
2ons = Horizons(ssoDir, ephemerides)
= Sso.getDirections(pp, fineTimes, horizons, name
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```
d plot the SSO-corrected attitude
a = Doubleld()
ec = Doubleld()
dir in dirs:
oRa.append(dir.raDegrees)
```







A *very* powerful tool with l**ots** of documentation to help you: •Help for new HIPE and HIFI users:

- Help-->Help Contents: Herschel Data Analysis Guide
- Help-->Help Contents: HIFI Data Reduction Guide
 - Contains launch pad for new users
- Help-->Video tutorials
- twitter.com/learnhipe
- •Help on more advanced topics:
 - Help-->Help Contents: HIFI Pipeline Specification
 - Help-->Help Contents: Herschel and HIFI Reference Manuals

But...





HIPE Help



Nobody wants to read manuals...





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Manual-free help



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"Show methods"



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	Jython Interpreter 102 of	2040 MB







HIPE 5.1.0 – /Applications/hipe_v5.1.0/scripts/hifi/pipeline/generic/Level1PipelineAlgo.py				
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PACS HrsPipelineAlgo				
🗹 Editor 🗙 SPIRE 🕹 Level 1 Pipeline Algo				
27 obs=task. OtfMappingPipelineAlgo				
28 apidName = PipelineUtils.getApidName(task.apid)				
<pre>29 if task.obs.level.get("level0_5").getProduct(apidName)== None: 30 print "No data available for obsid: " + str(obs.obsid) + " at level 0.5; apid: " + apidName</pre>				
31 return obs				
<pre>32 htp = obs.level.get("level0_5").getProduct(apidName).copy()</pre>				
33 ####################################				
<pre>obs=tast. * OtMappingPineIneXigs apiNName = TipeTineVtIS.getApidName(task.apid) if task.obs.level.get('level0_5').getProduct(apidName)== None: print "No data available for obsid: " + str(obs.obsid) + " at level 0.5; apid: " + apidName return obs htp = obs.level.get('level0_5').getProduct(apidName).copy() #################################</pre>				
<pre>36 params = PipelineConfiguration.getConfig(htp) 37 # some sanity checks on the data</pre>				
37 <i>* Some Safity checks on the data</i> 38 htp=checkDataStructure(htp=htp)				
# analysis the LO groups				
40 freqGrouping = checkFreqGrid(htp=htp, params=params)				
task.setCalibrationOutput(freqGrouping, 'FrequencyGroups', obs)				
# analysis the patterns found in the data, calibration is used for the chopper position values				
<pre>43 phases = checkPhases(htp=htp, calibration=obs.calibration)</pre>				
44 task.fillQuality(phases, 'PhaseChecks', obs)				
45 <i># prepares the bandpass and tsys, calibration is used for coupling coeff, mixer current tolerances, chopper positions</i> 46 hc = mkFluxHotCold(htp = htp, params=params, calibration=obs.calibration)				
<pre>task.setCalibrationOutput(hc, 'Tsys', obs)</pre>				
task.fillTsysTrendProduct(hc, obs)				
# compute that channel dependent weights, typically (default) using the radiometric formula				
doChannelWeights(htp=htp, cal=hc, params=params)				
# reference subtraction, calibration is used for chopper positions, mixer current tolerances				
doRefSubtract(htp=htp, params=params, calibration=obs.calibration)				
<pre># compute a baseline from the OFF positions - if applicable baseline = mkOffSmooth(htm = htm parame = parame)</pre>				
<pre>baseline = mkOffSmooth(htp = htp, params = params) task.setCalibrationOutput(baseline, 'Baseline', obs)</pre>				
# do the off subtraction, calibration is used for the mixer current tolerances				
doOffSubtract(htp=htp, cal=baseline, params=params, calibration=obs.calibration)				
# do the intensity calibration, calibration is used for the mixer current tolerances				
doFluxHotCold(htp=htp, cal=hc, params=params, calibration=obs.calibration)				
<i># do the velocity correction</i>				
61 doVelocityCorrection(htp=htp, aux=obs.auxiliary)				
62 <i># put the result into the observation context - remove level 0.5 product</i> 63 task.setSpectrumOutput(htp. obs)				



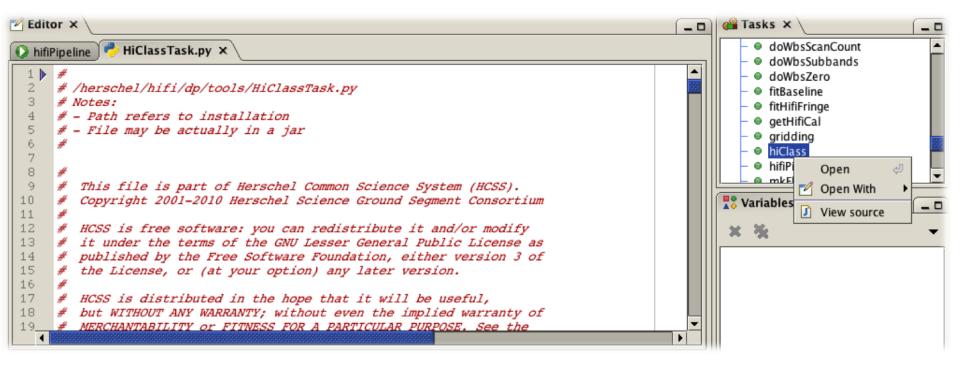






View source



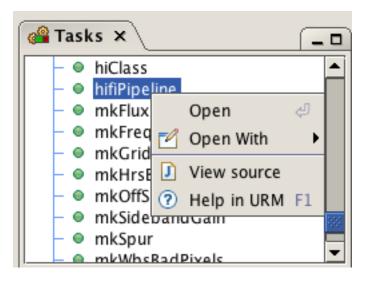






Help in URM





URM = Users Reference Manual

Contains information about the main classes and tasks you can use in your scripts

