





PACS Unchopped Spectroscopy

Dario Fadda













Why not chopping?

It is impossible to chop beyond 6 arcmin.

So, in case of extended sources or crowded fields, the only alternative is using the unchopped mode. Remember also that with a large chopper-throw only the central pixel is usable in pointed observations because of distortions.

Unchopping is more time efficient than chop-nodding.

However, it is very difficult to correct the signal for the varying response of the detector. So, only the strength of the line and not the underlying continuum can be precisely estimated. Also, the imperfect knowledge of the RSRF affects the reduction of unchopped data.

Users are discouraged to use unchopped modes whenever it is possible to use chop-nod modes.











NEW MODE!!!

With this call (AO2), a new unchopped mode is available. This new mode, called "bright line mode", allows one to save up to 30% of the time by cutting the wavelength range by 30%.

This makes sense only with bright lines since in other cases it is difficult to estimate the underlying continuum to evaluate the strength of the line. The user is discouraged to adopt this mode when the line is not detected at least at a 10-sigma level.







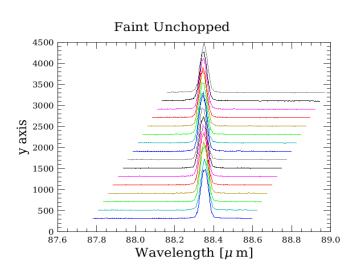


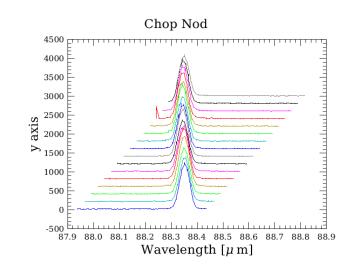


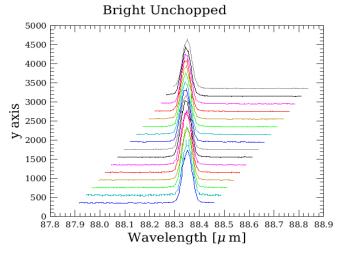


Wavelength coverage

Standard unchopped has the largest coverage. The bright mode is similar to chop-nod.













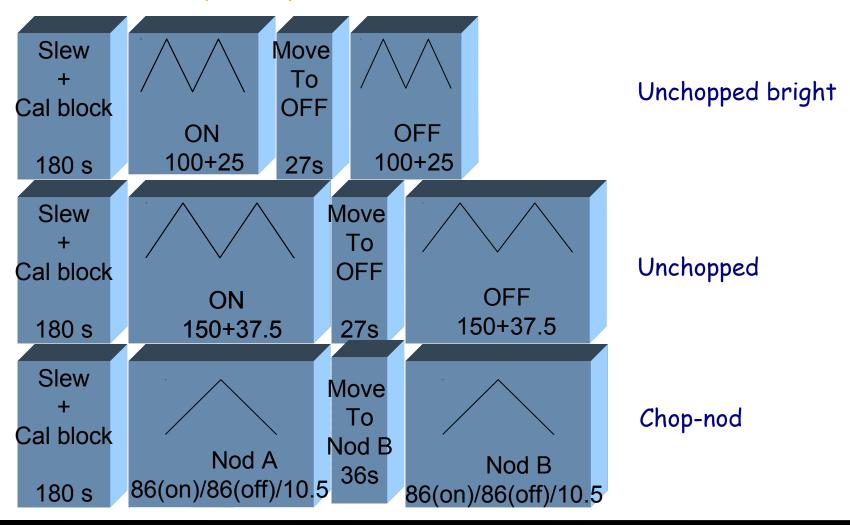








Time efficiency comparison





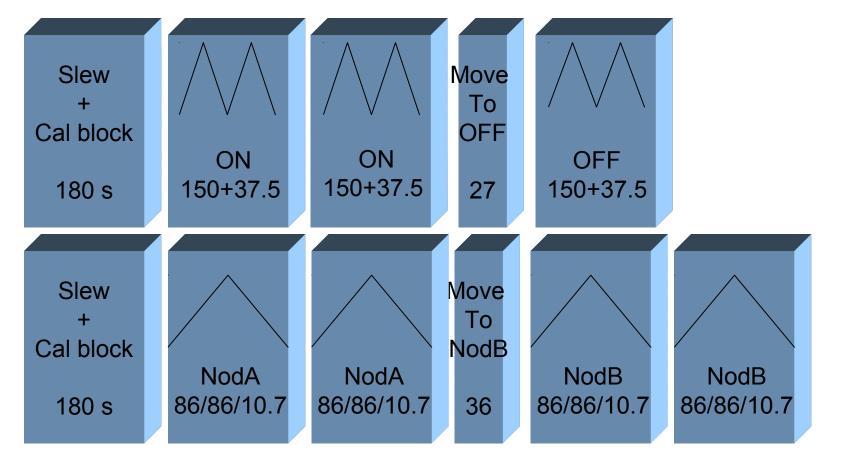








When repeating a scan, the off can be executed only once.















Pointing and raster modes

Two different strategies are possible:

- a) pointed observation: on-source is followed by an offsource observation. The user can repeat the on more times than the off.
- b) mapping observation: the user can choose the frequency of the off-source observation. It is recommended to have an OFF at least every 30 minutes.











OFF position

In principle, if interested only in the strength of the line, the OFF position is superfluos. However, there are at least two reasons for which it is recommended to have OFF positions:

a) correction of RSRF inaccuracies

Scaling the off to the on signal and subtracting it corrects for the RSRF imperfections.

b) spatial flat field

In case of a mapping observation, a good flat-field can be obtained using the OFF observation.















Typical errors

Default values are dangerous!

a) forget to define the OFF offset

The default is OFF by offset and it is set to 0, which means at the center of the observation

b) forget to define the frequency of OFF in mapping

The default is an OFF every 2 raster position, which is usually too much and inflates unnecessarily the length of an observation. In case of small maps (e.g. 3×3) one OFF at the end is enough. Otherwise, for large maps, a solution can be one OFF at the end of each row.













Find the most efficient way to observe a C+ line (157.7µm) in the nucleus of an extended object at 10:10:45 + 11:10:20, redshift = 0.01, peakFlux = 3 Jy, width = 250 km/s, continuum = 1 Jy.

Check if the line is unresolved ...

Remember that: $R=\lambda/FWHM$, $v[km/s]=c*FWHM/\lambda$

For flux: $\Delta v = c * FWHM/\underline{\lambda}^2$

 $F[W/m^2] \sim fluxPeak[Jy] \Delta v[Hz] 10^{-26}$

In the following, you can see how to prepare an observation and to check execution time and SNR.









Where to find more

Hspot manual

http://herschel.esac.esa.int/Docs/HSPOT/html/hspot-help.html

PACS observer's manual

http://herschel.esac.esa.int/Docs/PACS/html/pacs_om.html

Release notes

http://herschel.esac.esa.int/AOTsReleaseStatus.shtml#ReleaseNotes http://herschel.esac.esa.int/twiki/bin/view/Public/PacsAotReleaseNotes









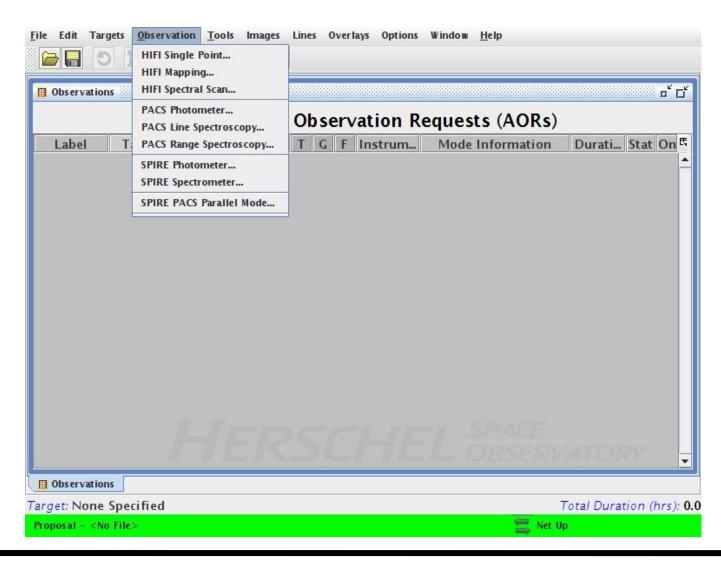






How to do it

Choose the observation mode.

















How to do it

Set the AOR label and click on "new target" to select your target.

==		PACS I	ine Spectros	co	ру		
Unique	AOR Label: PSp	ecL-0000					
			get: None Spec	ifie	33555		
	New Targ	et	Modify Target		Target List		
	Number	of visibl	e stars for the tar	get:	None Specified		
Se	lection of wave		elength Set	tin	ıgs		
Wa	velength ranges	[70-220] microns (2nd +	1st	orders)	_	
		РΔ	CS Line Edi	to	r		
Line Id Wavelen	gthRedshifted		ux Line Flux Co	-	20	Line WidtLir	ne Repeti
Add Lin	e Manually		e From Database		Modify Line	Delete Line	
	Redshift s	election					
	Unit Redsh	ift (z)	▼ Valu	ie C	0.000000		
	chopping and u	nchopp	ed scan N To con adjusti	ng, uml tro ing	tings grating scan or ber of cycles 1 I the absolute s the number of i	ensitivity cor	nsider
	ок		Cancel		He	lp	















How to do it

Choose the target name. If it is a known object (by NED of Simbad) can be resolved. Otherwise, insert the coordinates.

	Target		بـــــــــــــــــــــــــــــــــــــ
me (required):	NED -	Resolve the Name	
		Target Visibility	Background
ving			
Equatorial 12000	Proper Mot	ion	
		Use Proper Motion	
	PM RA (arcs	ec/year): 0.000	
2000.00	PM Dec (arc	sec/year): 0.000	
	ving Equatorial J2000	Equatorial J2000 Proper Mot	Equatorial J2000 Proper Motion Use Proper Motion PM RA (arcsec/year): 0.000















How to do it

Now, name the AOR and select the line(s) to observe. Lines can defined manually or selected from a list.

:	PACS Line Spectroscopy
Uı	nique AOR Label: M82-pointed
	Target: M82 Type: Fixed Single Position: 9h55m52.73s,+69d40m45.8s
	New Target Modify Target Target List_
	Number of visible stars for the target: 10 Star tracker target: Ra: 328.97 degrees Dec:-69.679 degrees
	Wavelength Settings Selection of wavelength ranges
	Wavelength ranges [70-220] microns (2nd + 1st orders) ▼
	PACS Line Editor
Line Id Wa	velength Redshifted Line Flux Line Flux Continuum Line Width Line Widt Line Repeti
Add	d Line Manually Add Line From Database Modify Line Delete Line
	Unit Redshift (z) Value 0.000000
Source ty	Observing Mode Settings Nodding, grating scan or mapping cycles Number of cycles 1 To control the absolute sensitivity consider adjusting the number of integration cycles. Observation Est_ Add Comments_ AOR Visibility
	OK Cancel Help

















How to do it

When using the line database, select a line by checking the box.

Origin	Name	Transition	Wavelength	Line Width	Selected	
DEFAULT	ОН	2Π1/2J=3/	163.122	1	1111	-
DEFAULT	SH	2П3/2Ј=5/	216.784	1		
DEFAULT	Hα	Η15α	169.412	1		
DEFAULT	Hα	Η16α	204.412	1		
DEFAULT	CII	C+	157.741	1	V	
DEFAULT	HD	1-0	112.07	1		
	Manager faci	s and modify sility: HSpot Me	enu -> Lines -			Lin













How to do it

The selected line appears on the list. It is now time to assign the redshift. To change the number of repetitions, double click on the line.

==			PACS Line	Spectro	scopy			□×
	Unique A0	R Label: M8	2-pointed]
			Target: M82 tion: 9h55m		69d40m			
			visible stars fo target: Ra: 32			679 degrees		
	Selecti	ion of wave	Waveler		ttings			
	Wavele	ngth ranges	[70-220] mic	rons (2nd	+ 1st order	s)		
			PACS	Line Ed	litor			
Line Id	Wavelength	Redshifted .			7777	Line Width	Line Widt	Line Repeti
CII C+	157.741	157.85	100	10^-18		1.00	km/s	1
	Add Line Ma	nually	Add Line Fron	m Database	Mod	ify Line	Delete Li	ne
		Redshift						
		Unit Redsh		▼ Va	lue 0.0006	77		
Source		oping and u		Node can To co	ling, grati Number of ontrol the sting the r	ng scan or	ensitivity	consider
		ОК		Cancel		Hel	р	*

















How to do it

In this pop-up window it is possible to set the flux of the line and other features. It is also the place to define the number of line repetitions.

==	Update a	line	□×	
	Spectral line parameters			
	Line ID	CII C+		
	Wavelength (µm)	157.741		
	Line flux unit	10^-18 W/m^2	-	
	Line flux	0.00		
	Continuum flux density (mJy)	0.00		
	Line width unit	km/s	_	
	Line width (FWHM)	1.00		
Lii	ne repetition factor Line repetition	2		
	e relative line strength (fractie) can be set by the line rep Note: the sum of line repeti on-source time per i	etition factor for (ition factors affec	each line.	
	ок	Cancel		

















How to do it

It's time to select the observing mode. Clicking on Pointed, you can select between the two unchopped modes. Then, select the off either by offset or by specifying the position.

Observing Modes	· • • • • • • • • • • • • • • • • • • •	
	erving Mode Settings oose one of the modes below.	
None selected Pointed Pointed w	with dither Mapping	
Obser	rving mode selection	
Chop	pping/nodding	
○ Chop	oping/nodding (bright lines)	
○ Unch	opped grating scan	
○ Unch	opped grating scan (bright lines)	
Chopper throw Small Medium Large Off position Type RA offset (arch Dec offset (arch RA (degrees) Dec (degrees)		
ОК	Cancel	















How to do it

If the Mapping mode is selected, one has to choose also the steps and the frequency of the off position.

The default is on OFF after two raster positions, usually too much!

Observing Modes		•			
	ring Mode Settings e one of the modes below.				
Observ	ing mode selection				
Choppin	•				
Choppin	g/nodding (bright lines)				
Wavelength s witching					
 Unchopped grating scan 					
 Unchopped grating scan (bright lines) 					
○ Small A	hopper avoidance angle ngle from (degrees) 0.00 ngle to (degrees) 0.00				
Raster Map	Off position				
Map reference frame	Type	By offset			
● Instrument ○ Sky	RA offset (arcmins) 0.00				
Raster point step (arcseconds) 2.0	Dec offset (arcmins) 0.00				
Raster line step (arcseconds) 2.0	RA (degrees) 0.00				
Orientation angle (degrees) 0.0	Dec (degrees) 0.00				
Number of raster points per line 2		e Position			
Number of raster lines 1	Repeat off-position Number of raster pos	after nth raster position ition 2			
ОК	Cance	el			













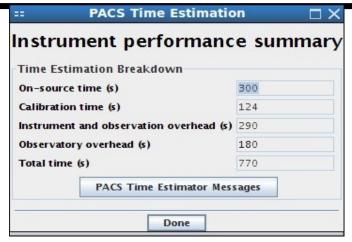




How to do it

Once the planning is completed, it is possible to make some checks, like computing the total time of the observation.

Note that the on-time source in chop-nod includes the time spent in the off position. In unchopped mode, the off position time is included in the overhead.



Unchopped scan

e summary
688
124
81
180
949
ages

Chop-nod

















How to do it

It is possible to repeat an observation without retaking the calibration block. In the AOR window, just Change the number of cycles.

In this case, with two cycles and repeating the line twice, It allows to have two OFF positions observed in unchopped scan. Obviously the overhead will be bigger than repeating the line four times.

== PACS Time Estimation	n □×
Instrument performanc	e summary
Time Estimation Breakdown	
On-source time (s)	600
Calibration time (s)	124
Instrument and observation overhead (s)	576
Observatory overhead (s)	180
Total time (s)	1356
PACS Time Estimator Mess	sages
Done	

Unchopped

Instrument performanc	e summary
Time Estimation Breakdown	
On-source time (s)	1376
Calibration time (s)	124
Instrument and observation overhead (s)	158
Observatory overhead (s)	180
Total time (s)	1714
PACS Time Estimator Mess	sages

Chop-nod















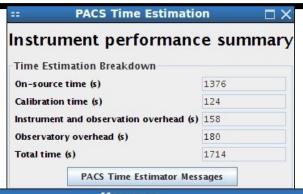


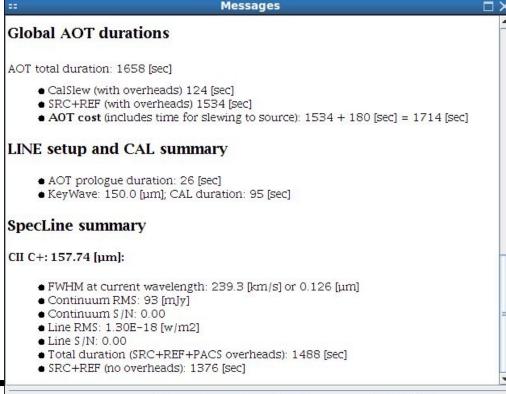
How to do it

Additional information is available by pressing the "PACS Time Estimator Messages"

If the line flux entered is correct, the SNR of the line will be displayed.

This can help in comparing different techniques of observation.









Save messages



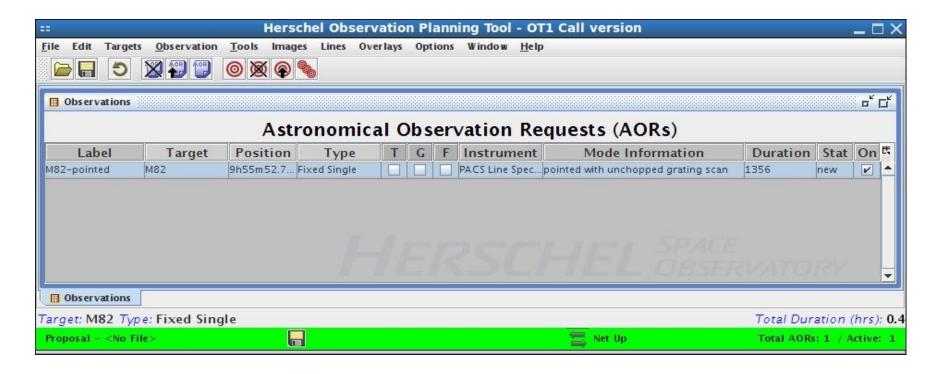
Cancel







How to do it



Once satisfied with the AOR, click OK in the AOR window and a line with the new AOR will appear in the main window. To save, go under "File" and "Save AOR(s) and Target(s).





