

# Studying exoplanets with SOFIA

Claudia Dreyer

Institute for Planetary Research, German Aerospace Center

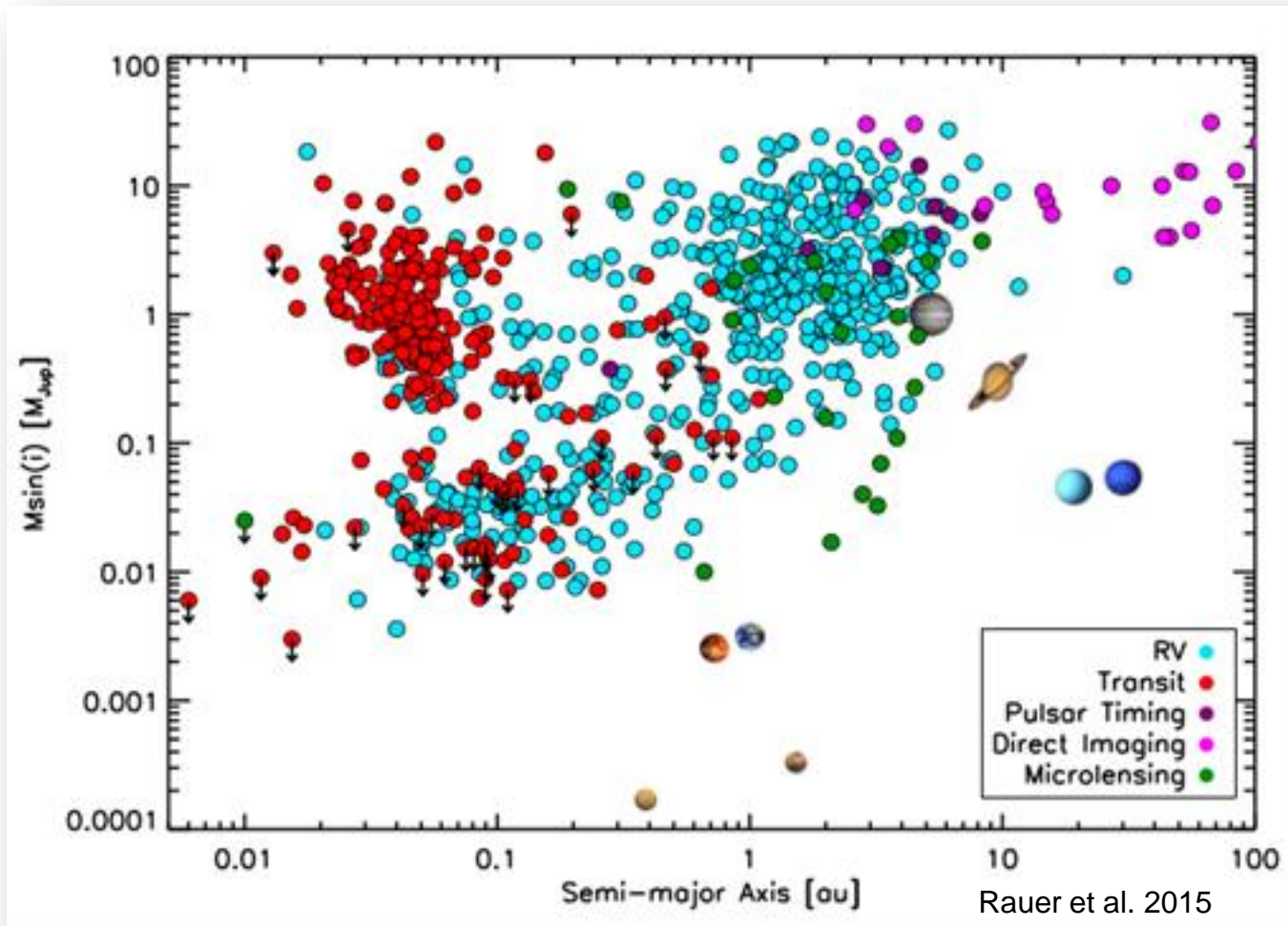
14.10.2015

Wissen für Morgen



# Known Exoplanets

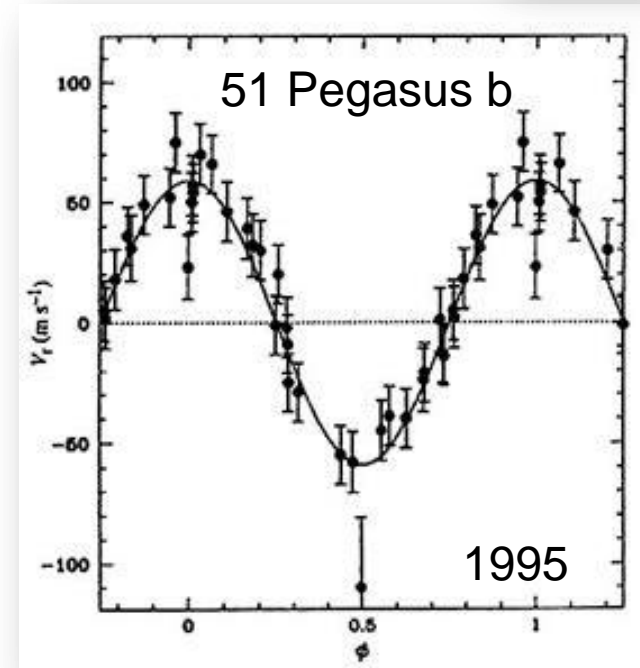
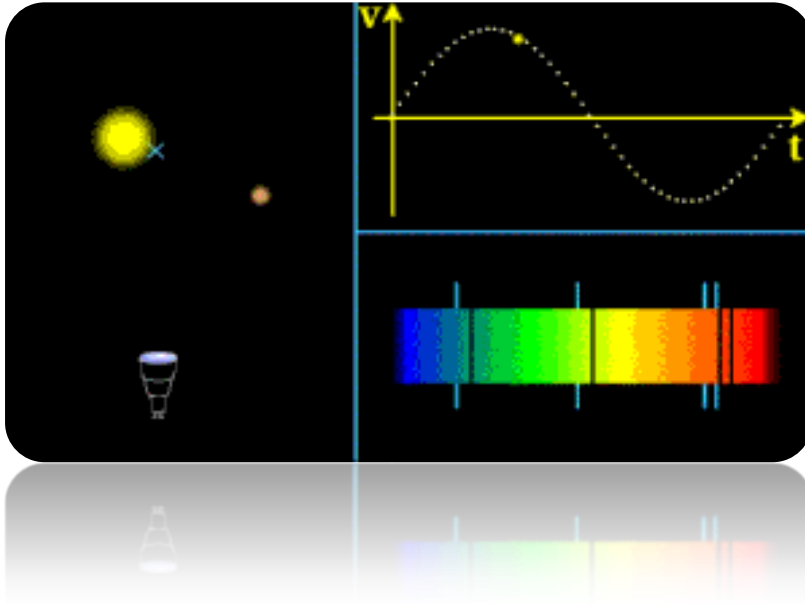
Current status: ~1600 confirmed exoplanets and ~3800 planet candidates  
(October 2015; exoplanets.org)



# Radial velocity method



Michel Mayor  
and  
Didier Queloz



→ orbital period and semi-major axis

→ minimum planet mass ( $m \sin i$ )

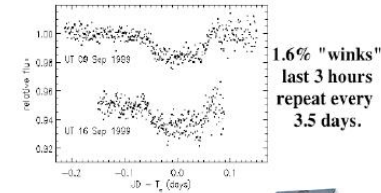
- Period: ~4 days
- Semi-major axis: 0.05 AU
- Mass: ~ ½ M<sub>J</sub>



# Transit Methode



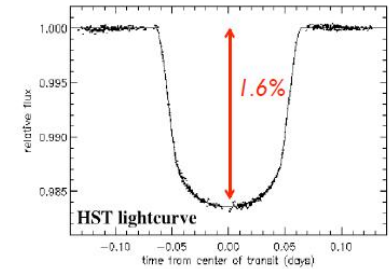
## 1999 First Planet Transits



Charbonneau and Brown  
STARE 10 cm telescope



## HD 209458 The First Transiting Planet



Diameter = 1.3 x Jupiter  
Mass = 0.6 x Jupiter  
Temp = 2000 Centigrade

**This "Hot Jupiter"  
is a Gas Giant.**

→ orbital Period and semi-major axis

→ Radius

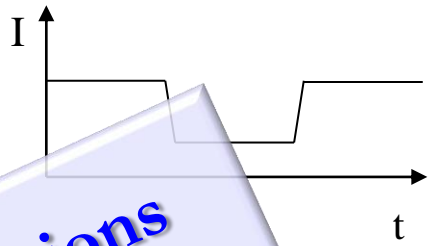
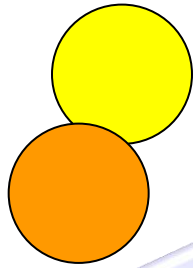
→ Inclination  $i$

- Period:  $\sim 3\frac{1}{2}$  days
- Semi-major axis: 0.05 AU
- Radius:  $\sim 1\frac{1}{2} R_J$

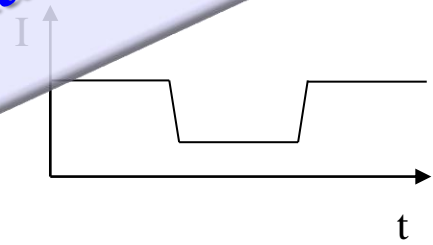


# Transit false alarm possibilities

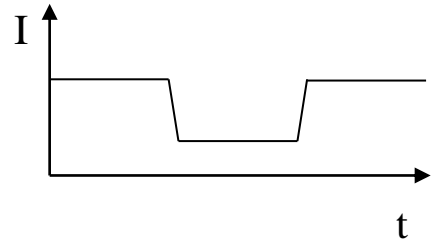
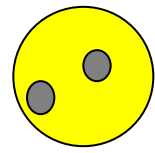
\* Grazing eclipse of binaries



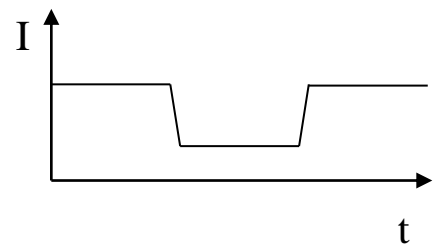
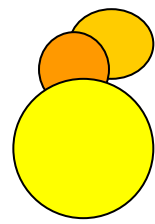
\* Transit of brown dwarf



\* Stellar spots



\* Binary system in background

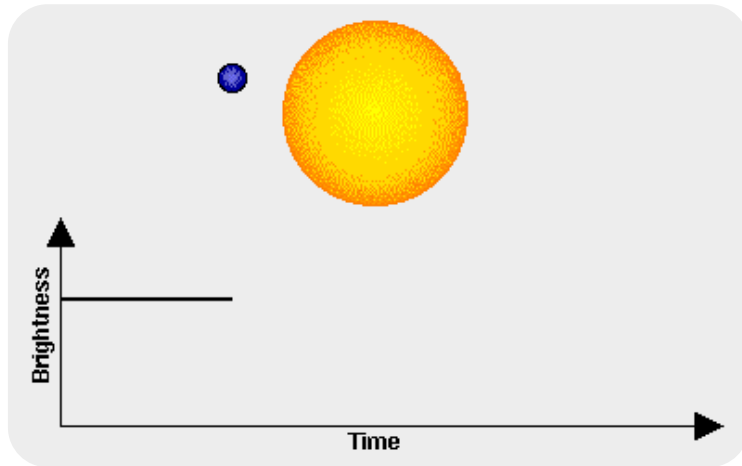


**→ Need for follow-up confirmations**



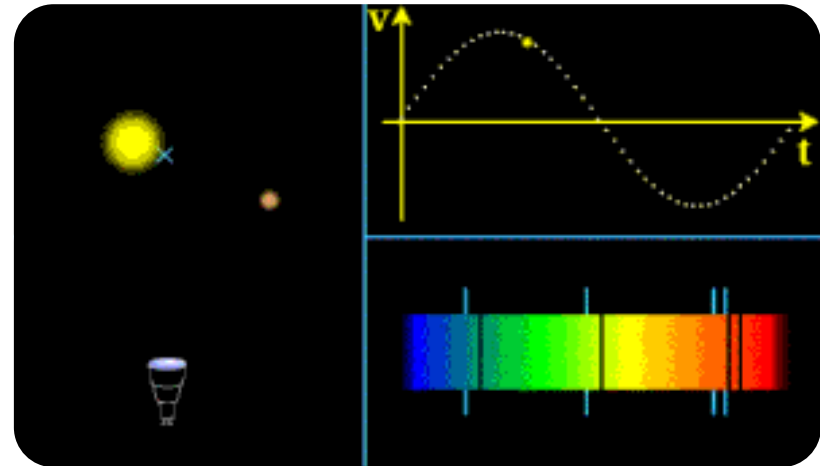
# Exoplanet Detection methods

## Transit Method



- Orbit parameters
- Orbital inclination,  $i$
- Planet radius

## Radial velocity method



- Orbital parameters
- Minimum planet mass,  $m \sin i$

**True planet mass and mean density**



# Our Solar System

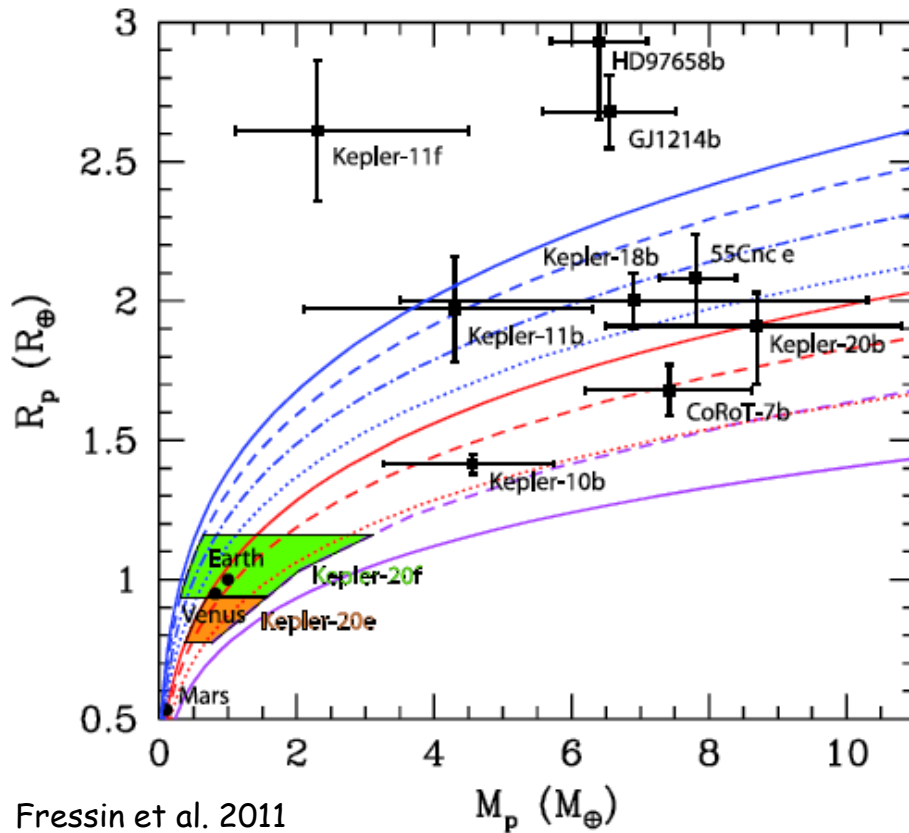


and its planets are **NOT** the general rule

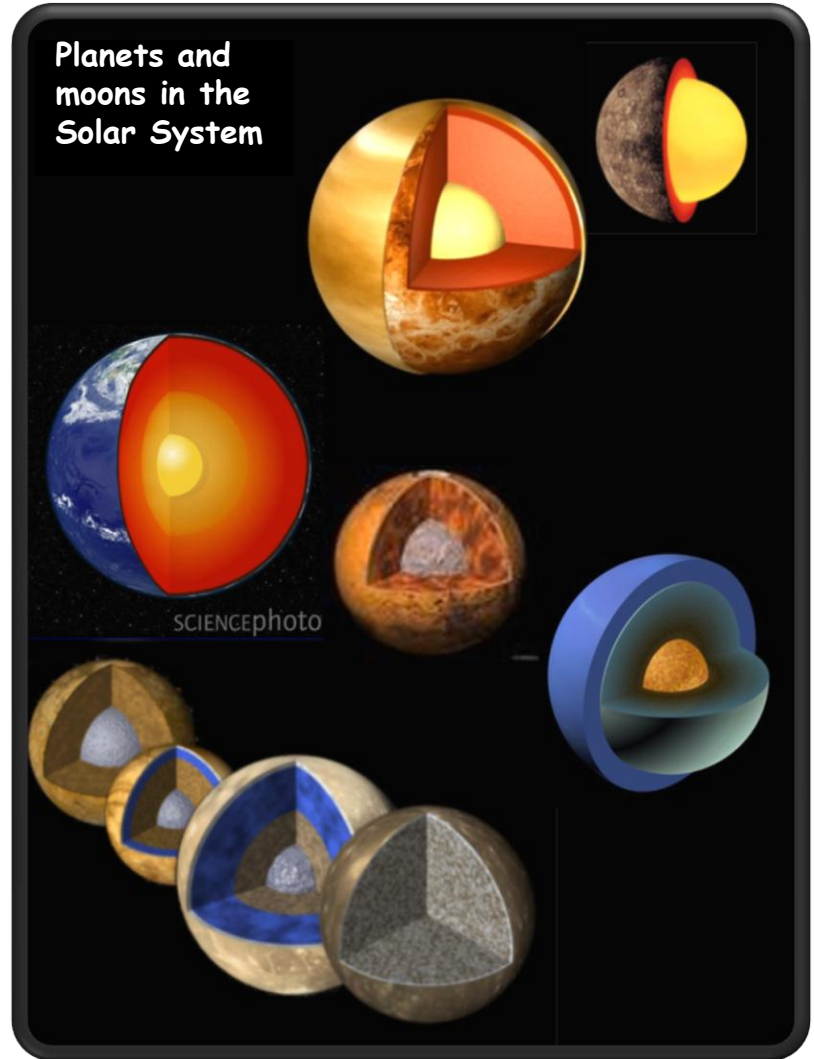
- Small planets  $\neq$  rocky, large planets  $\neq$  gaseous
- Small planets  $\neq$  close-in, large planets  $\neq$  far out
- New class of planets not existent in our solar system



# Planet diversity



Fressin et al. 2011



Degeneracy between the mass and bulk composition and a possible atmosphere in theoretical models



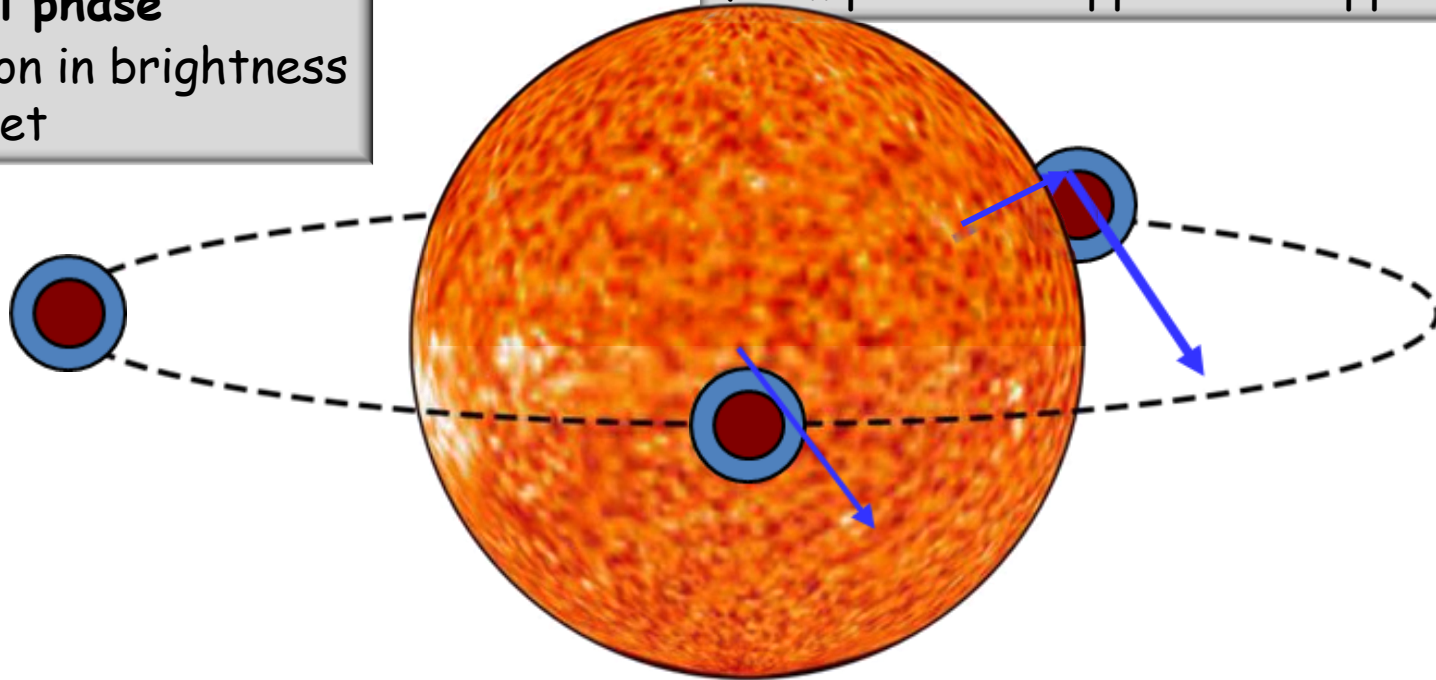


# Exoplanet atmosphere characterization

**Orbital phase**  
variation in brightness  
of planet

## Secondary Eclipse (Occultation)

Thermal radiation and reflected light  
from planet disappear and appear again

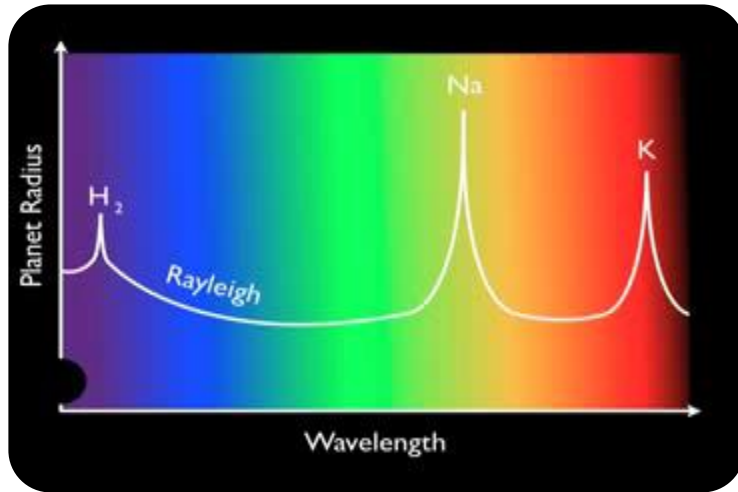


## Primary Eclipse (Transit)

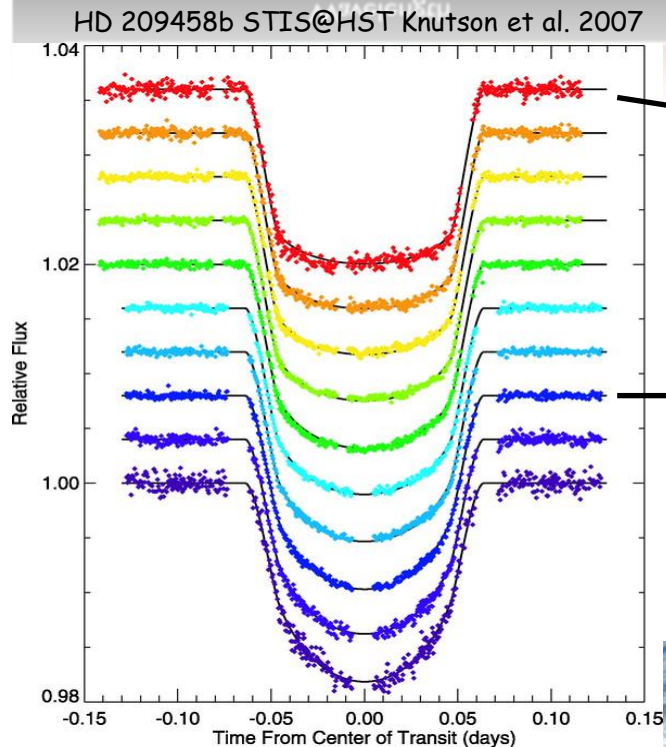
Radiation from star transmitted  
through the planets atmosphere



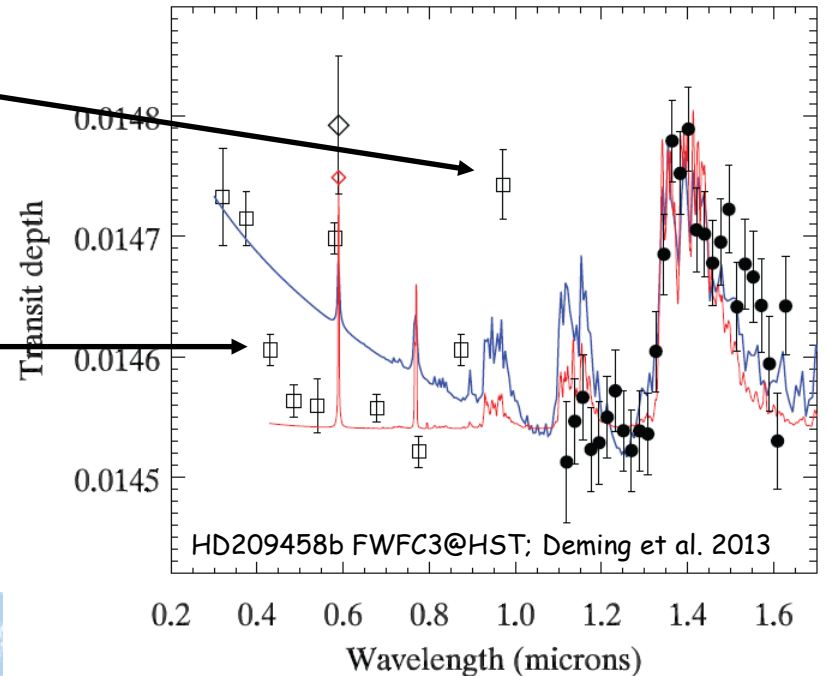
# Characterizing Exoplanet Atmospheres with Transit-Spectrophotometry



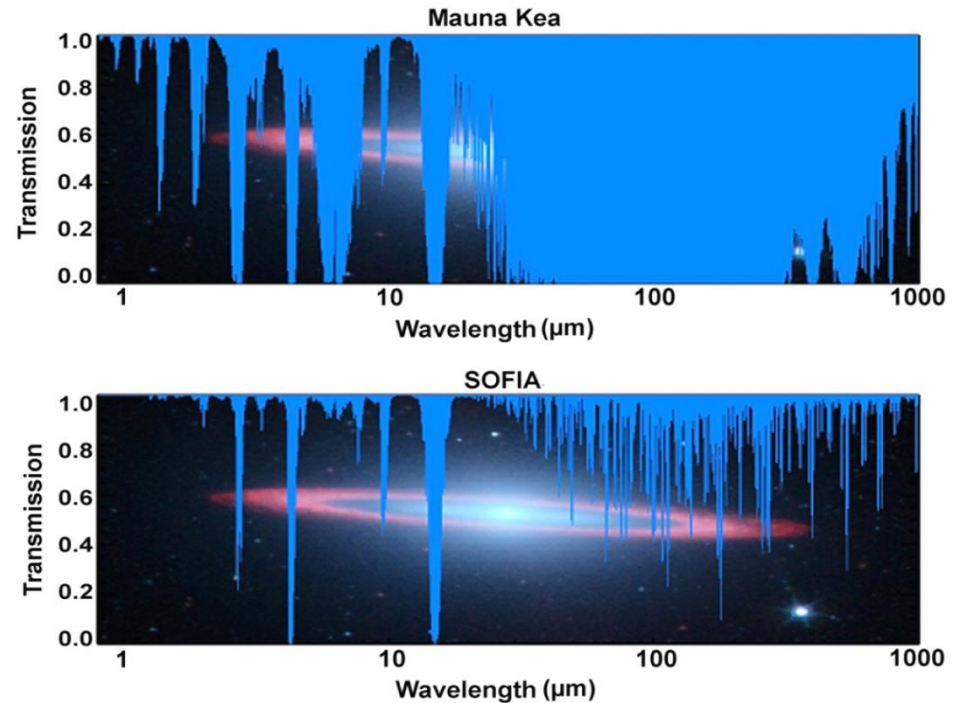
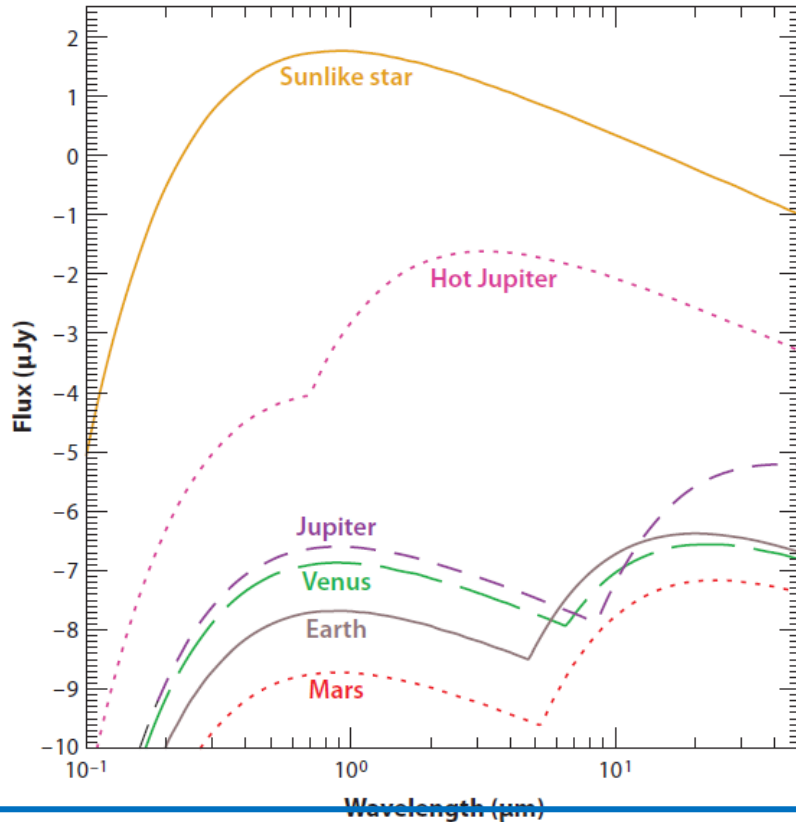
- Atmospheric composition
- Atmospheric structure (T/P profile, day/night side heat exchange)
- Formation history



$$\sim \left( \frac{R_p}{R_s} \right)^2$$



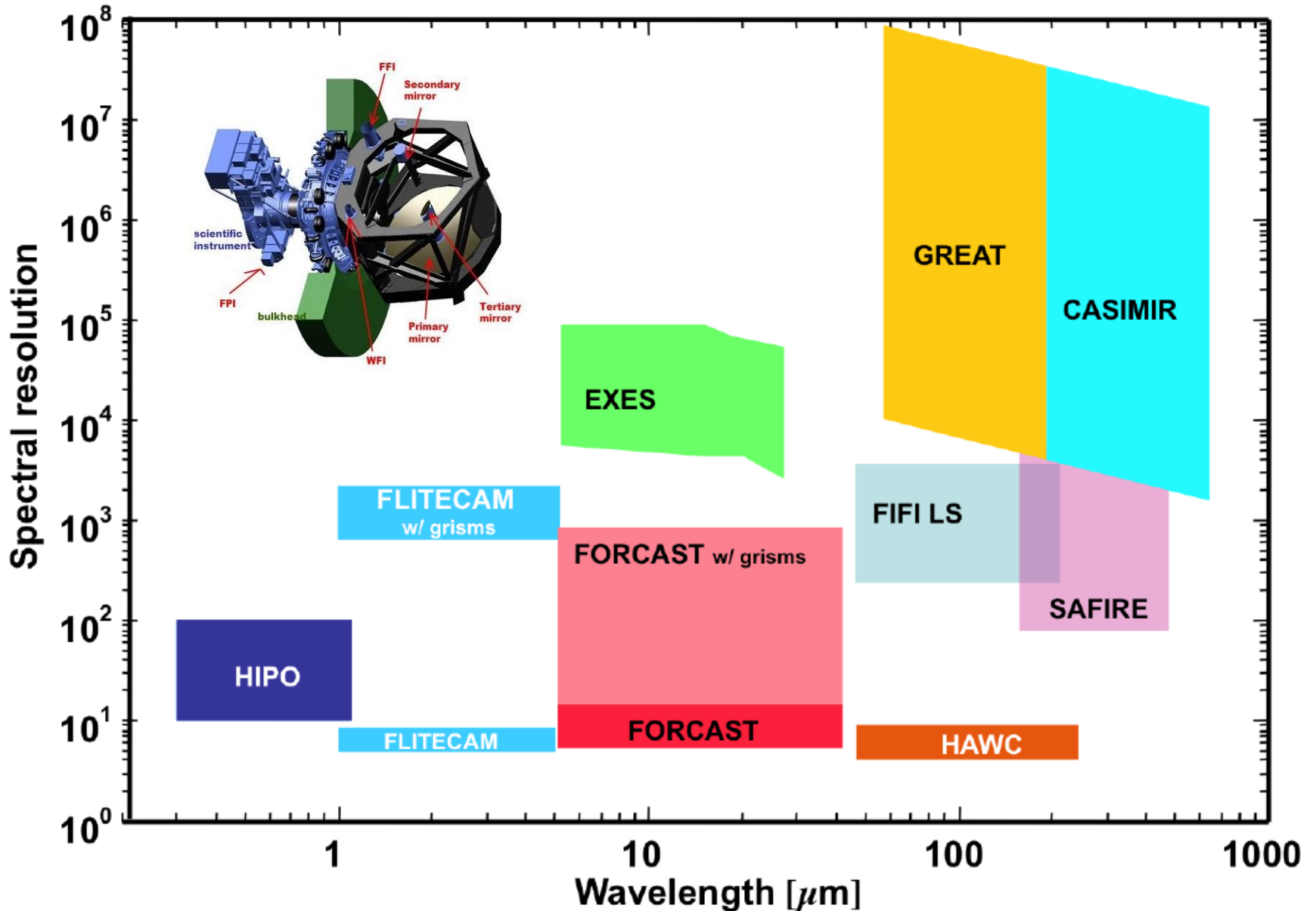
# SOFIA & Exoplanets



## SOFIA

- operates at wavelengths where the planet's black-body temperature peaks and improved contrast ratios star/planet
- operates at low temperatures (240K); therefore the contribution from the thermal background are significantly reduced.
- can observe simultaneously at infrared and optical wavelengths, 'FLIPO' mode, thereby obtaining light curves for a single transit event over a wide range in wavelength

# SOFIA Science Instruments



# The best-known Hot Jupiter HD 189733 b

## Planet HD 189733b

<b>Discovery:</b>	2005 (Bouchy et al)
<b>Mass (<math>m</math>):</b>	$1.144 \pm 0.056 M_J$
<b>Radius (<math>r</math>):</b>	$1.138 \pm 0.027 R_J$
<b>Dichte (<math>\rho</math>):</b>	$0.963^{+0.088}_{-0.079} \text{ g cm}^{-3}$
<b>Period (<math>p</math>):</b>	$2.21857312 \pm 7.6e-07 \text{ d}$
<b>Transit duration (T):</b>	$0.0760 \pm 0.0017 \text{ d}$
<b>Semi major axis (a):</b>	$0.03142 \pm 0.00052 \text{ au}$

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Period (p):  $2.21857312 \pm 7.6e-07 \text{ d}$

## Host star 189733

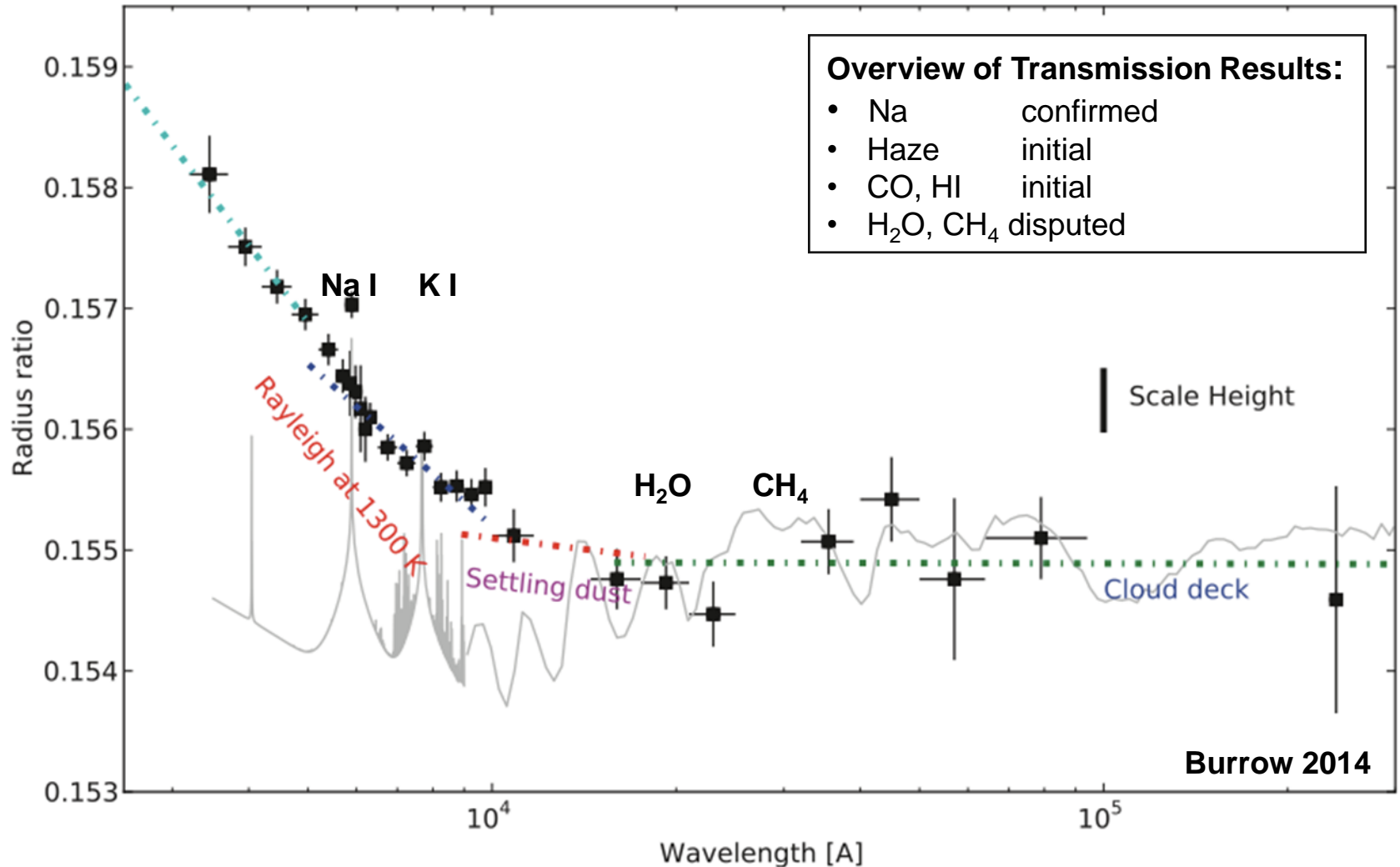
<b>Right ascension (<math>\alpha</math>):</b>	$20^{\text{h}} 00^{\text{m}} 43.0^{\text{s}}$
<b>Declination (<math>\delta</math>):</b>	$+ 22^{\circ} 42' 39''$
<b>Magnitude (<math>m_V</math>):</b>	7,67
<b>Distance (d):</b>	$19.3 \pm 0.2 \text{ pc}$
<b>Spectral type:</b>	K1-K2
<b>Mass (<math>m</math>):</b>	$0.8 \pm 0.4 M_{\odot}$
<b>Radius (<math>r</math>):</b>	$0.805 \pm 0.016 R_{\odot}$
<b>Temperature (T):</b>	$4875 \pm 43 \text{ K}$
<b>Age:</b>	0,6 Gyr

Age: 0,6 Gyr

Temperature (T):  $4875 \pm 43 \text{ K}$

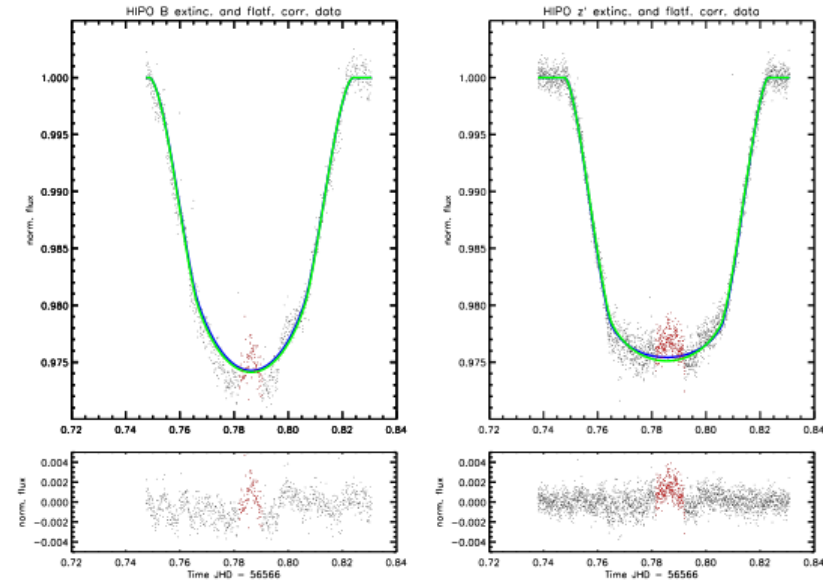
# Hot Jupiter HD 189733 b

Transmission spectrum of HD 189733b, showing the core of the sodium Na and potassium K lines, and the signature of aerosols



# First exoplanet transit observation: HJ HD189733b

- Examine the presence of a Rayleigh slope in the optical (using HIPO's B & z' filters)
- Examine the existence of H<sub>2</sub>O absorption in the NIR (using FLITECAMs 1.85  $\mu\text{m}$  filter)



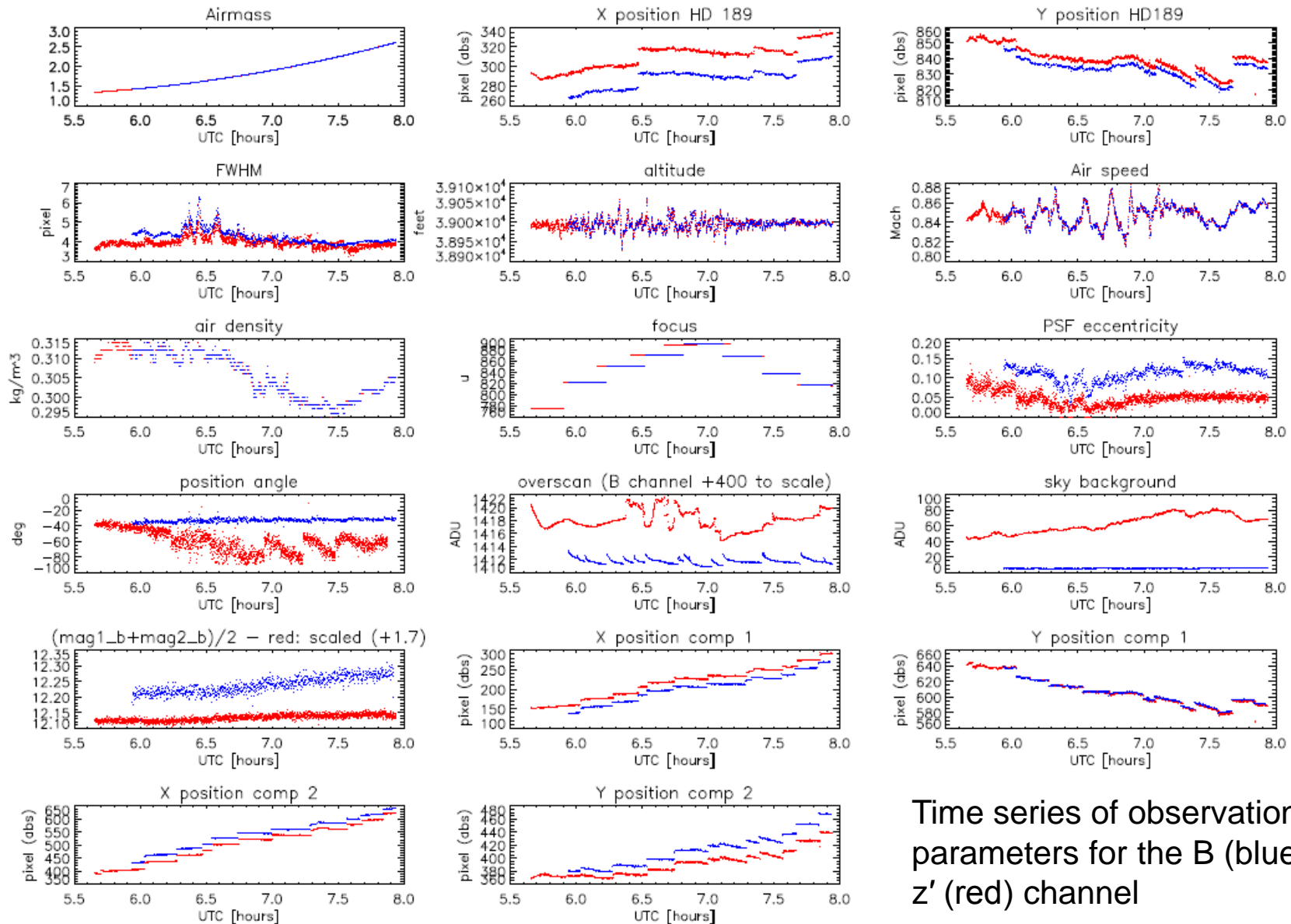
Extinction & flat field corrected absolute photometry of HIPO's B (left) and z' (right) channels before removal of correlated noise

Youtube: Exoplanet observations with SOFIA

Angerhausen et al. 2015



# First exoplanet transit observation: HJ HD189733b

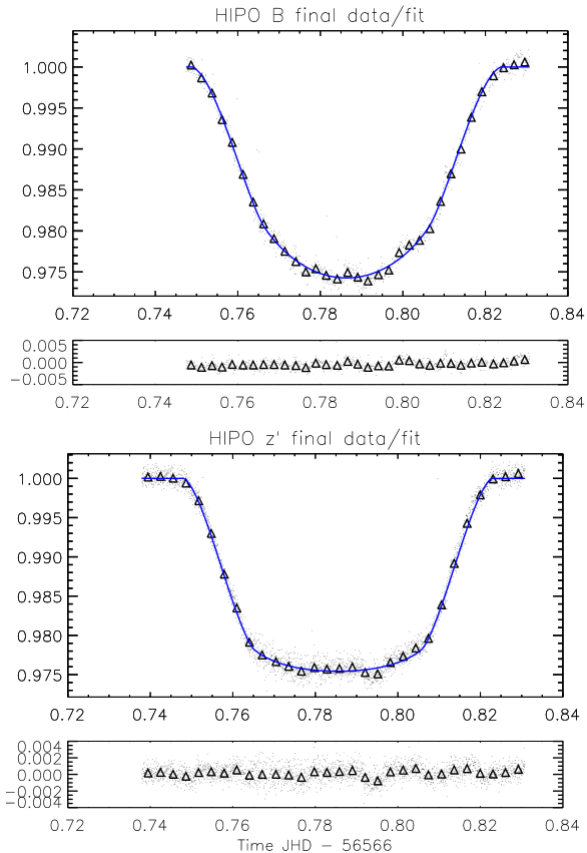


Angerhausen et al. 2015

Time series of observational parameters for the B (blue) and z' (red) channel



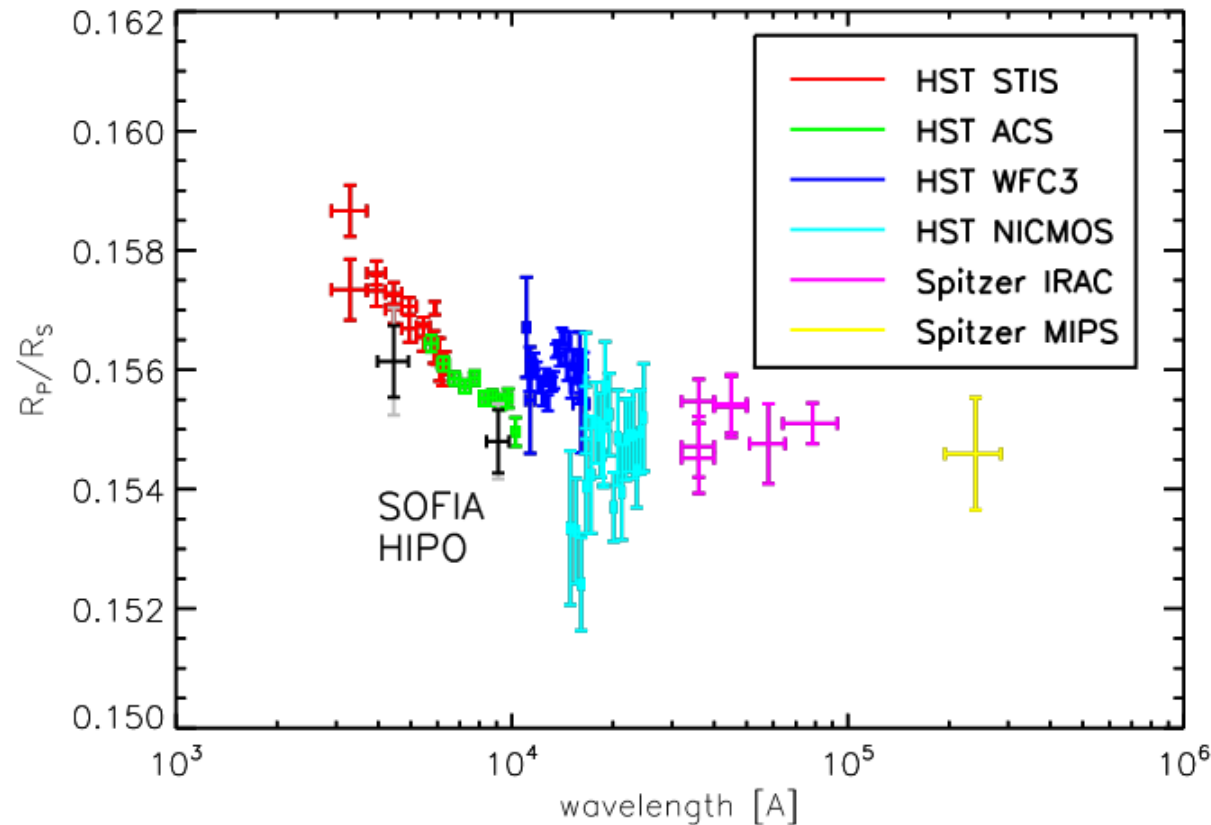
# First exoplanet transit observation: HJ HD189733b



## Final light curve in HIPO B & z'

Absolute photometry of a transit of HD 189733b observed by SOFIA/HIPO in the B (445 nm) and z' (905 nm) bands, corrected for extinction, flat field, and correlated noise

## Transmissionspectrum

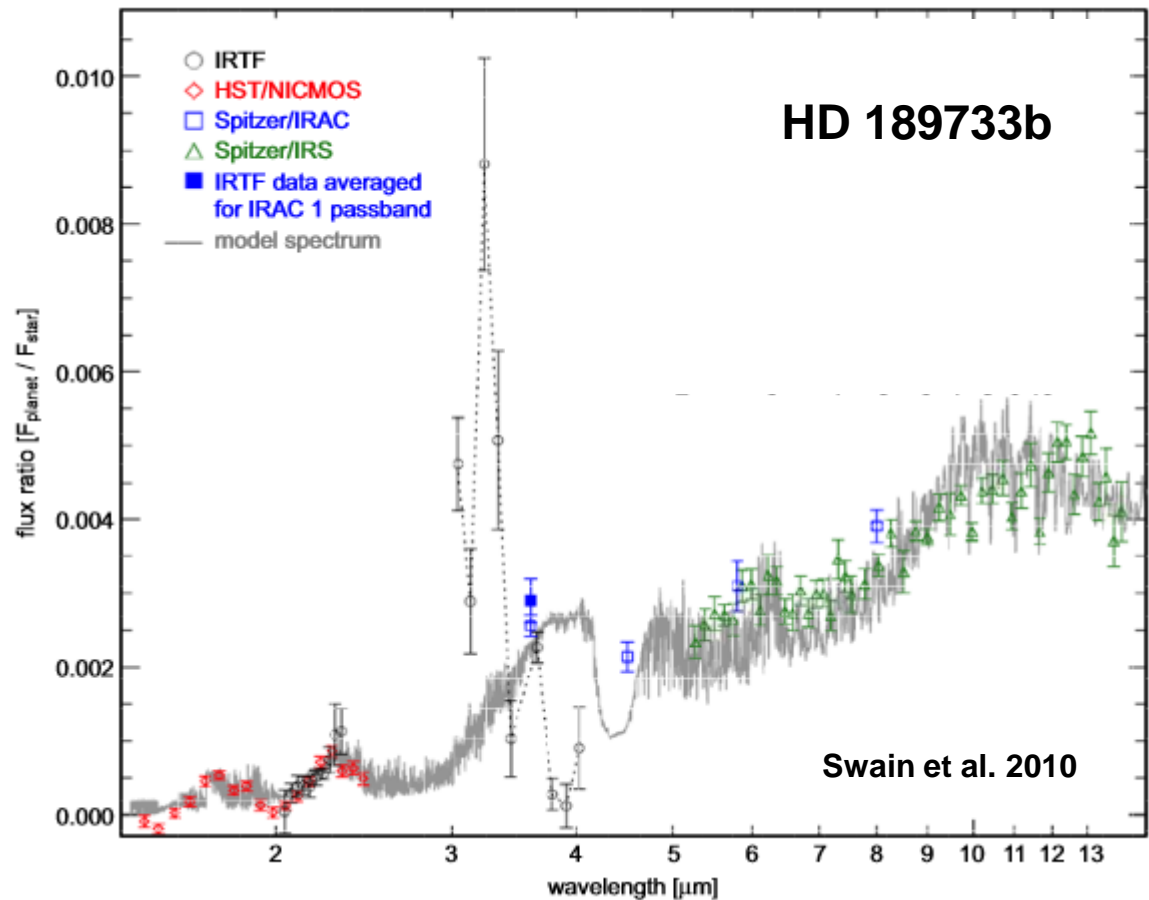


Angerhausen et al. 2015

# The Origin of non-LTE Emission on Dayside of a hot-Jupiter Exoplanet

- ground-based observations of unexpected strong emission around  $3.25 \mu\text{m}$
- Non LTE methane  $\text{CH}_4$  emission ?
- Take off date Oct. 1, 2015 (PDT)  
PI: Mark Swain, JPL

## Emission spectrum



# GJ 1214 - Super Earth or mini-Neptune ?

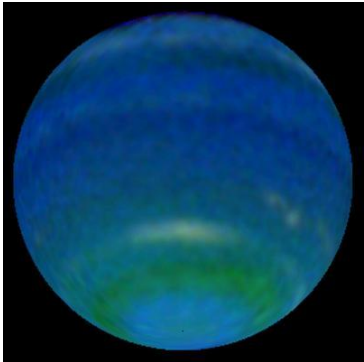
## Planet GJ 1214 b

<b>Discovery:</b>	2009 (Charbonneau et al)
<b>Mass (<math>m</math>):</b>	$6.55 \pm 0.98 M_{\oplus}$
<b>Radius (<math>r</math>):</b>	$2.678 \pm 0.13 R_{\oplus}$
<b>Dichte (<math>\rho</math>):</b>	$1.870 \pm 0.400 \text{ g cm}^{-3}$
<b>Period (<math>p</math>):</b>	$1.58040456 \pm 1.6 \times 10^{-7} \text{ d}$
<b>Transit duration (<math>T</math>):</b>	$\sim 52 \text{ min}$
<b>Semi major axis (<math>a</math>):</b>	0.01411 au

## Host star GJ 1214

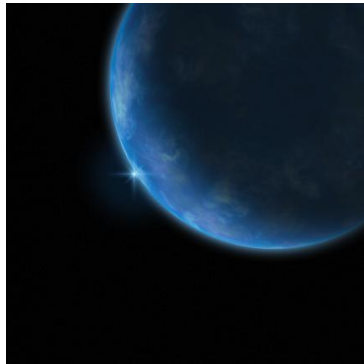
<b>Right ascension (<math>\alpha</math>):</b>	$17^{\text{h}} 15^{\text{m}} 18.942^{\text{s}}$
<b>Declination (<math>\delta</math>):</b>	$+04^{\circ} 57' 49.69''$
<b>Magnitude (<math>m_V</math>):</b>	$14.71 \pm 0.03$
<b>Distance (<math>d</math>):</b>	$13.0 \pm 0.9 \text{ pc}$
<b>Spectral type:</b>	dM 4.5
<b>Mass (<math>m</math>):</b>	$0.157 M_{\odot}$
<b>Radius (<math>r</math>):</b>	$0.2064^{+0.0086}_{-0.0096} R_{\odot}$
<b>Temperature (<math>T</math>):</b>	$3,026 \pm 130 \text{ K}$
<b>Age:</b>	6 Gyr

# Possible Compositions of GJ 1214b



## a. “Mini-Neptune” Scenario:

- Rock / ice interior + hydrogen-dominated atmosphere
- (mostly  $H_2$  + trace  $H_2O$ ,  $CH_4$ , etc.)



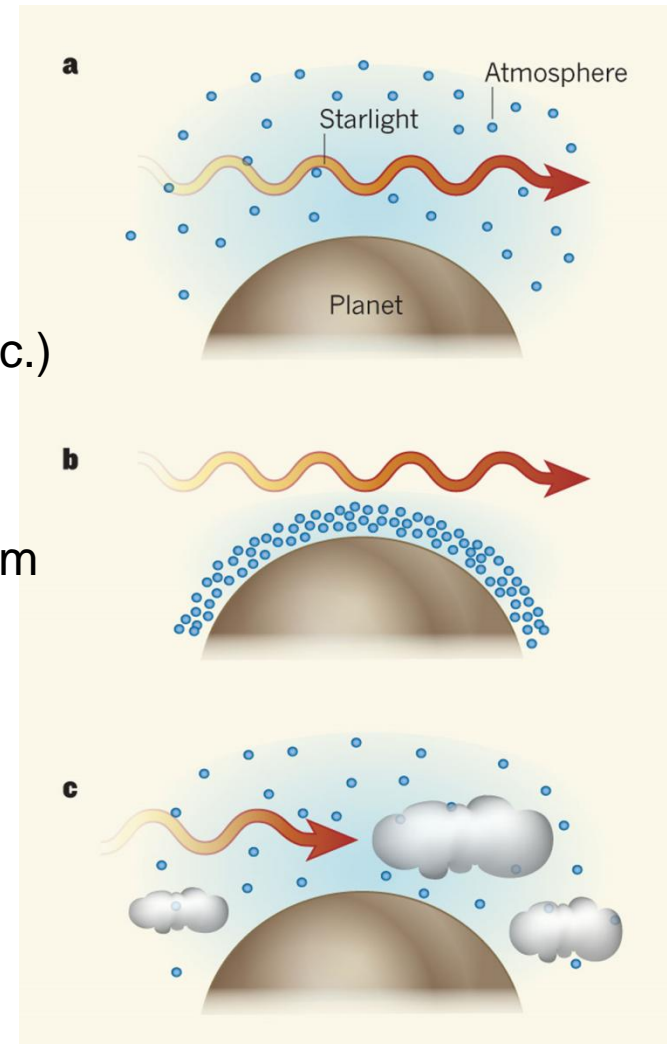
## b. Water World Scenario:

- Mostly  $H_2O$  - ice interior + steam atmosphere



## c. Clouds Scenario

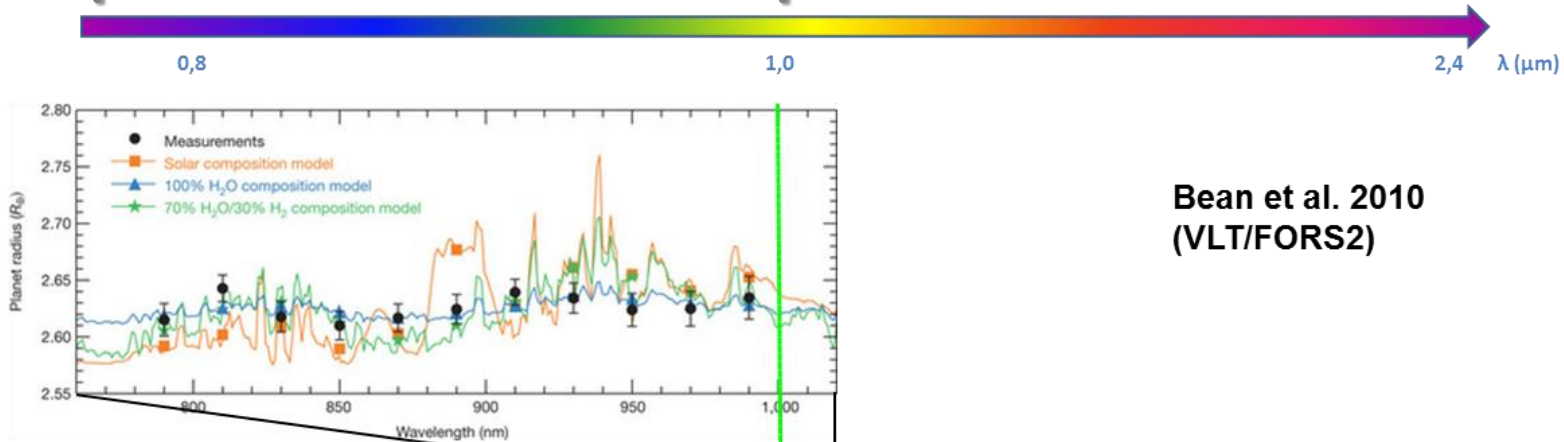
- Rock / ice interior + hydrogen-dominated atmosphere
- Clouds



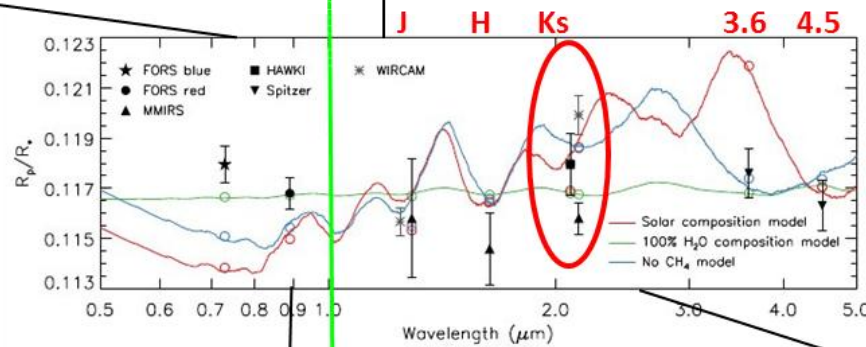
(Rogers & Seager, 2010; Nettelmann et al. 2011)



# Determining the composition of GJ1214b's atmosphere: Super-Earth or Mini-Neptune

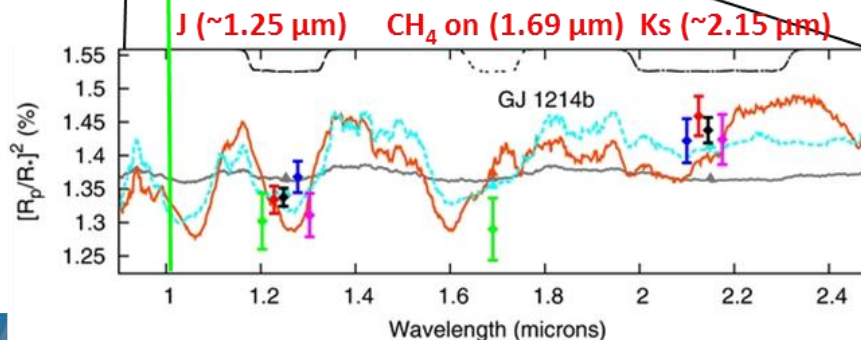


Bean et al. 2010  
(VLT/FORS2)



Bean et al. 2011  
(VLT/FORS)  
(MMIRS)

Miller-Ricci Kempton et al. 2011,  
Crossfield et al. 2011 (KECK)



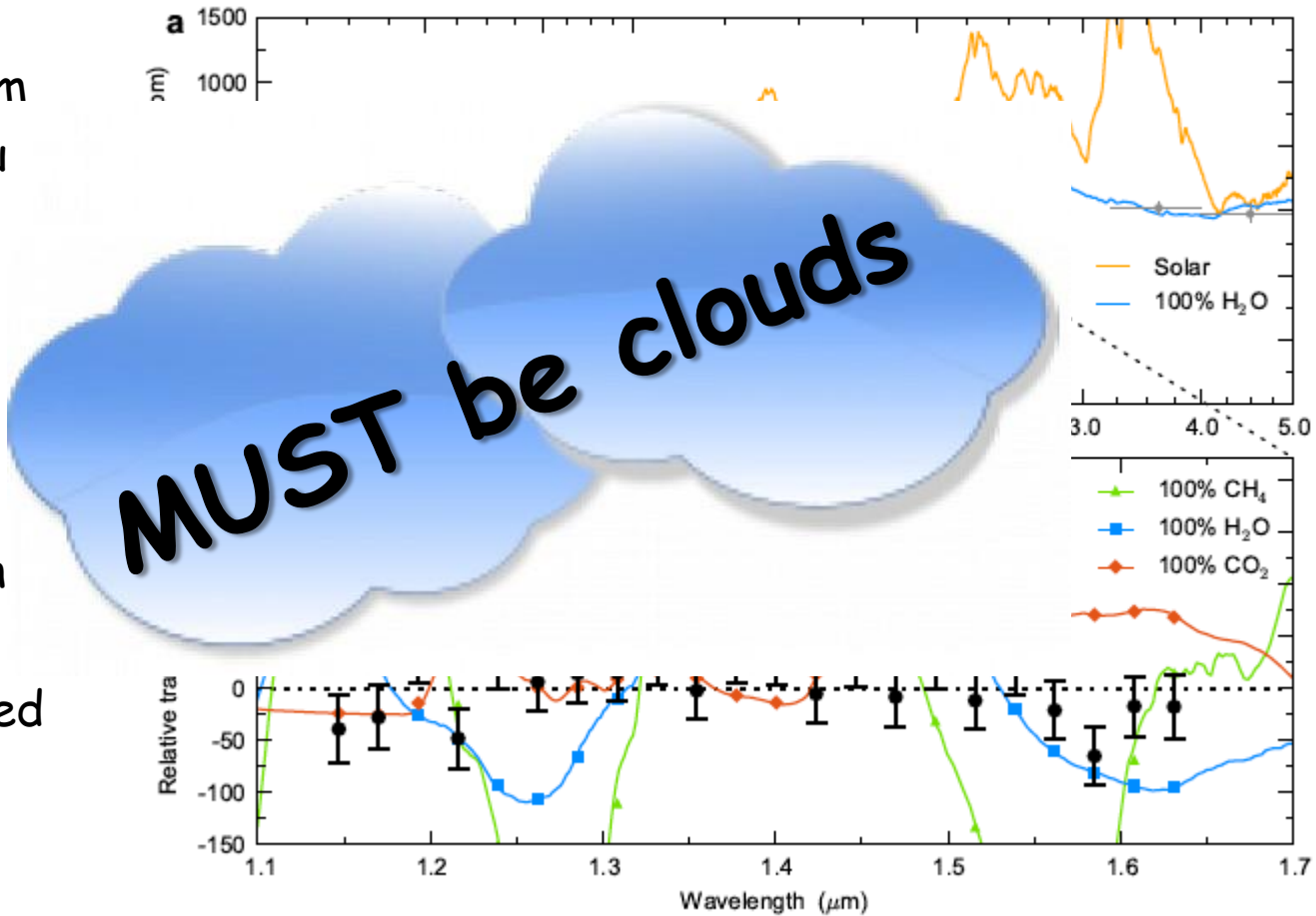
Croll et al. 2011  
(CFHT/ WIRCAM)



# Determining the composition of GJ1214b's atmosphere: Super-Earth or Mini-Neptune

- Featureless spectrum  
→ high-mean molecular weight atmosphere clouds

- Different high-mean molecular weight atmospheres are ruled out



Kreidberg et al. 2014, WFC3@HST



# GJ 1214b - Super-Earth or Mini-Neptune?

Optical & NIR spectro-  
photometry with FLIPO

## HIPO

Blue: 0.3-0.6  $\mu\text{m}$

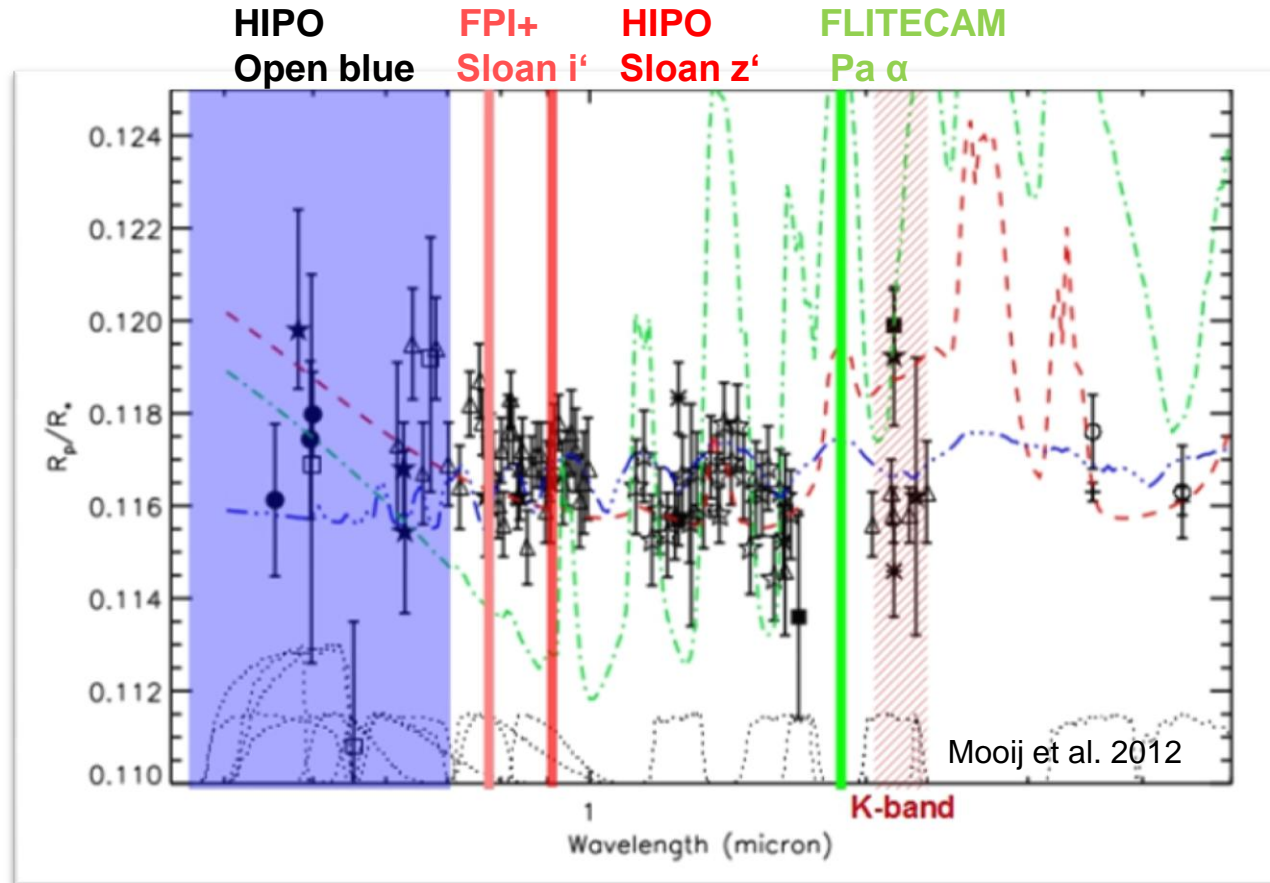
z': 0.9  $\mu\text{m}$

## FPIplus

i': 0.8  $\mu\text{m}$

## FLITECAM

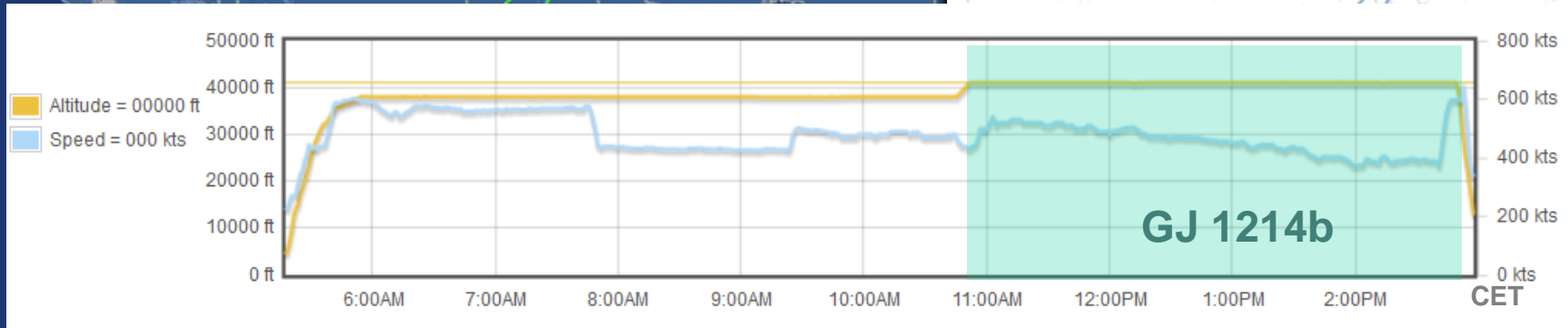
Pa  $\alpha$  cont: 1.9  $\mu\text{m}$



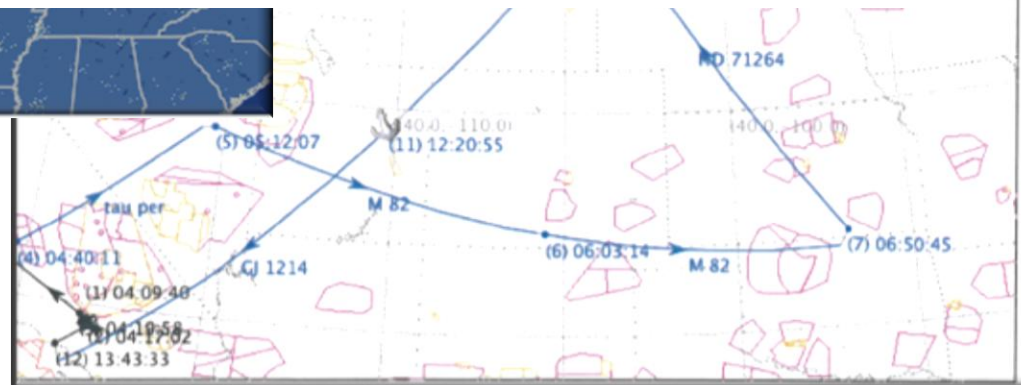
**SOFIA:** crucial wavelength region not observable  
from the ground: FLITECAM *Paschen a cont* 1.90  $\mu\text{m}$   
close to Ks band where contradictory data was obtained

# GJ 1214b - Super-Earth or Mini-Neptune?

FLT 149 (2014/26/2)



Flight Plan Name: File: 201402\_FP\_06\_WX12.fp  
 Flight ID: 2014/02/27  
 Est. Takeoff Time: 2014-Feb-27 04:09 UTC  
 Est. Landing Time: 2014-Feb-27 14:08 UTC  
 Flight Duration: 09:59  
 Weather Forecast : 1200 Wed Feb 26 2014 - 0000 Sat Mar 01 2014 UTC  
 Saved: 2014-Feb-26 15:15 UTC User: bwarrington





# FLT 149 (2/26/14) - People@work

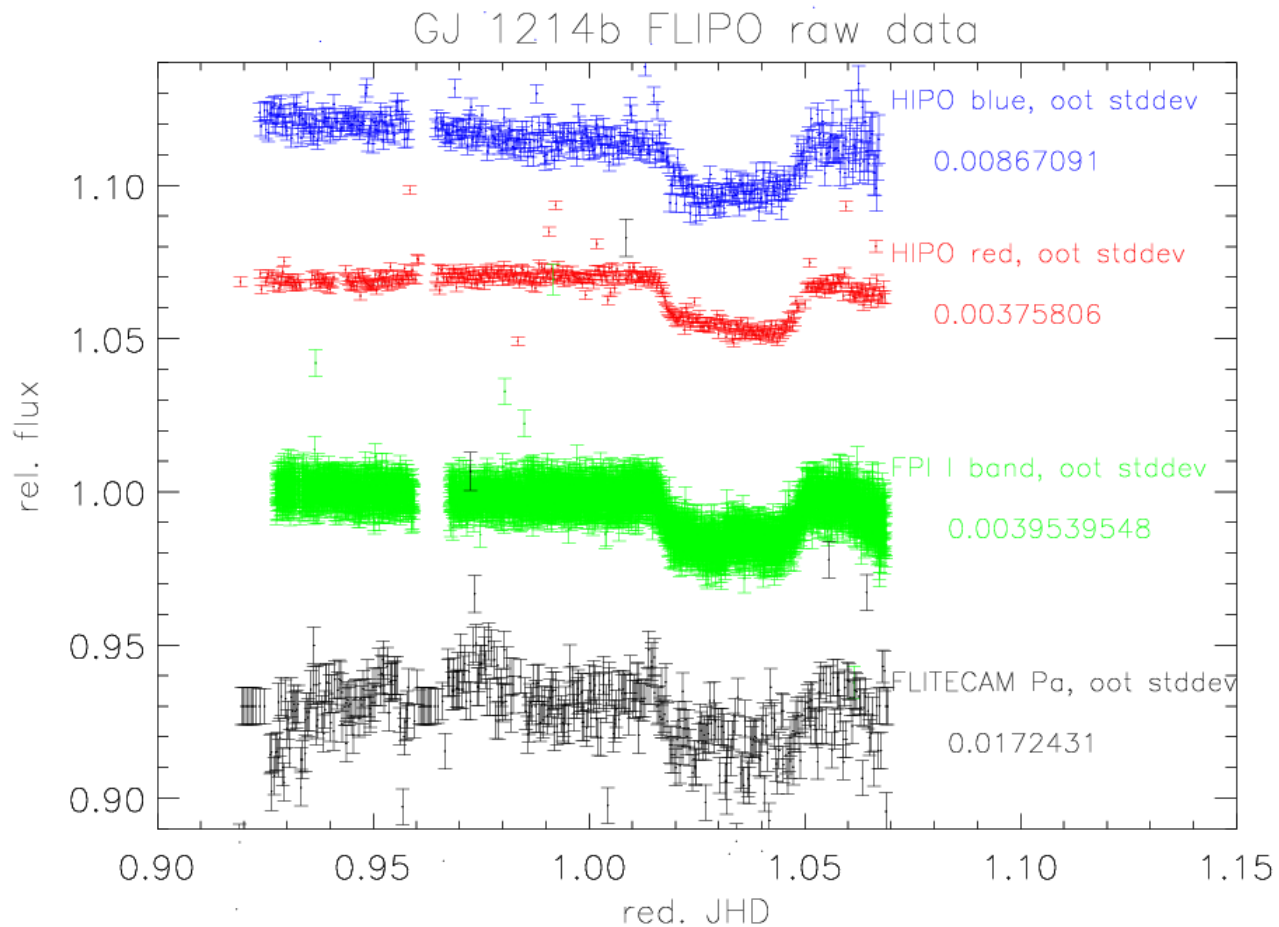


# GJ 1214 b raw data



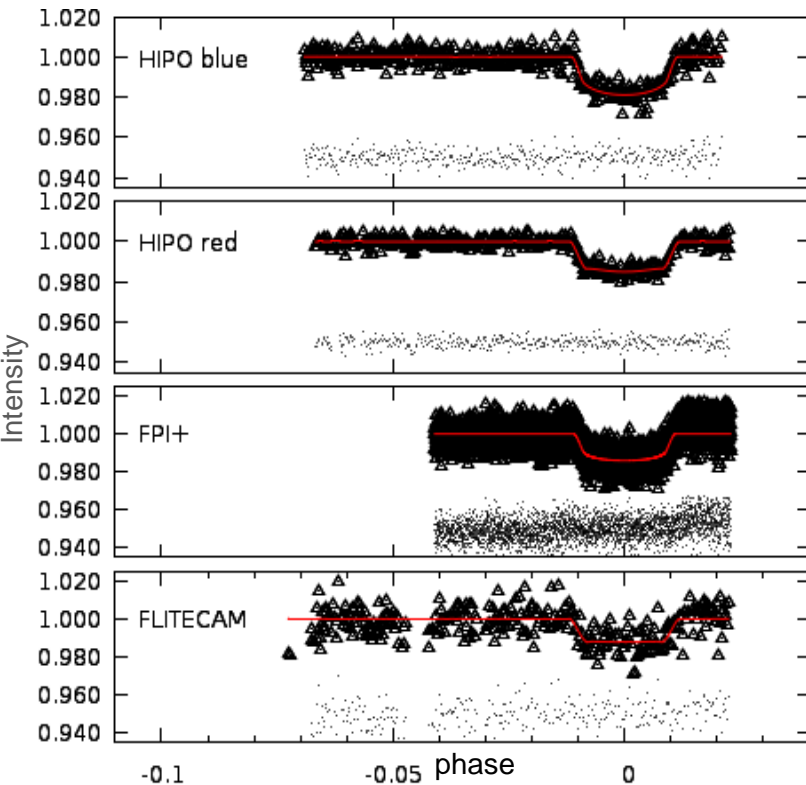
FLT 149 (2/26/14)

Final data set consists of about ~4500 measurements of 25 sec integrations in HIPO blue, HIPO red, FLITECAM and 2.5 sec integrations in FPI+



# GJ 1214b - Super-Earth or Mini-Neptune?

## Preliminary results

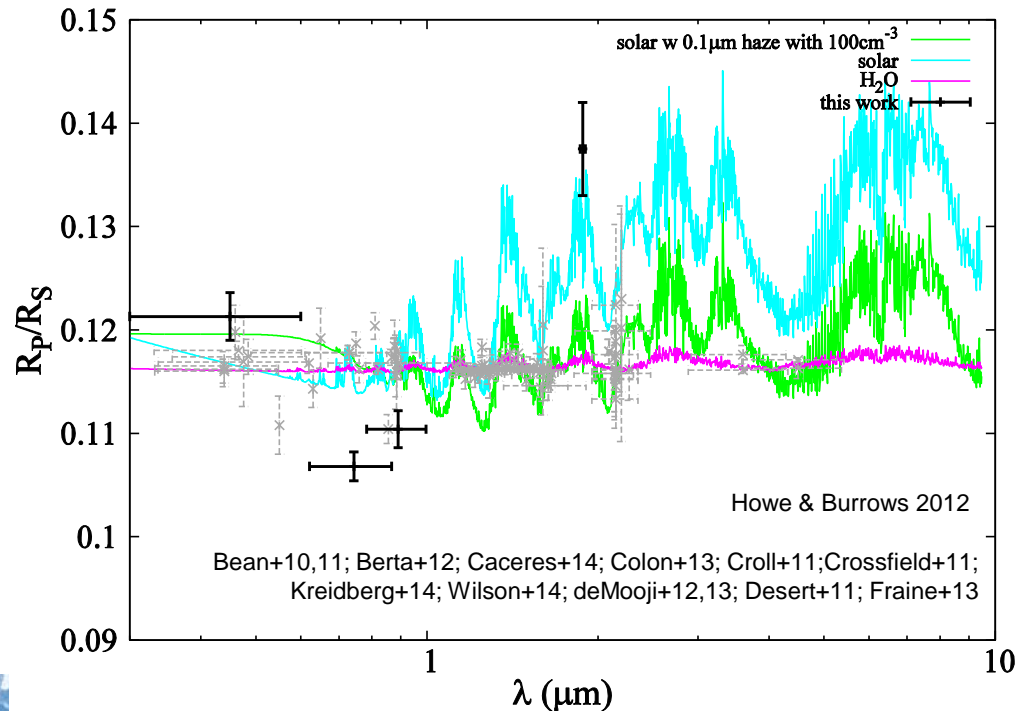


Light curves

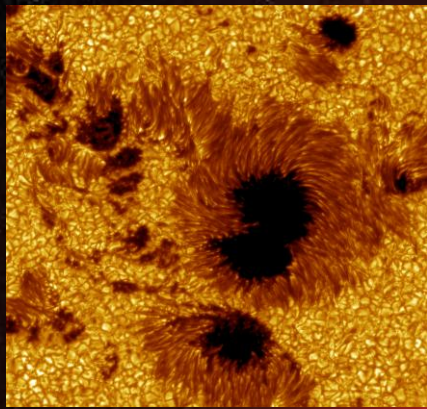
Transmission spectrum

### Fit results

Parameter	HIPO Blue	HIP Red	FPI	FLITECam
<b>P (d)</b>	<b>1.58040482</b>	<b>1.58040482</b>	<b>1.58040482</b>	<b>1.58040482</b>
$T_{14}$ (P)	$0.02287 \pm 0.0006$	$0.02317 \pm 0.00045$	$0.02188 \pm 0.00038$	$0.0235 \pm 0.0011$
$R_p/R^*$	$0.1234 \pm 0.0042$	$0.1199 \pm 0.0023$	$0.1083 \pm 0.0025$	$0.1122 \pm 0.0037$
$b$ ( $R^*$ )	$0.05 \pm 0.20$	$0.05 \pm 0.17$	$0.08 \pm 0.17$	$0.62 \pm 0.12$
u+	$0.63 \pm 0.35$	$0.54 \pm 0.31$	$0.64 \pm 0.24$	$-0.14 \pm 0.48$
u-	$0.49 \pm 0.85$	$0.58 \pm 0.65$	$0.04 \pm 0.63$	$1.42 \pm 0.91$
a/ $R^*$	16.62	15.42	16.06	12.48

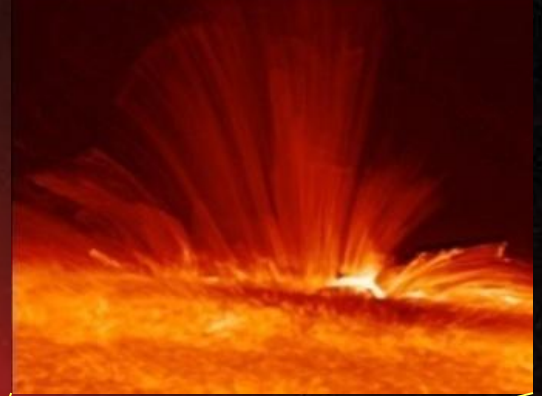
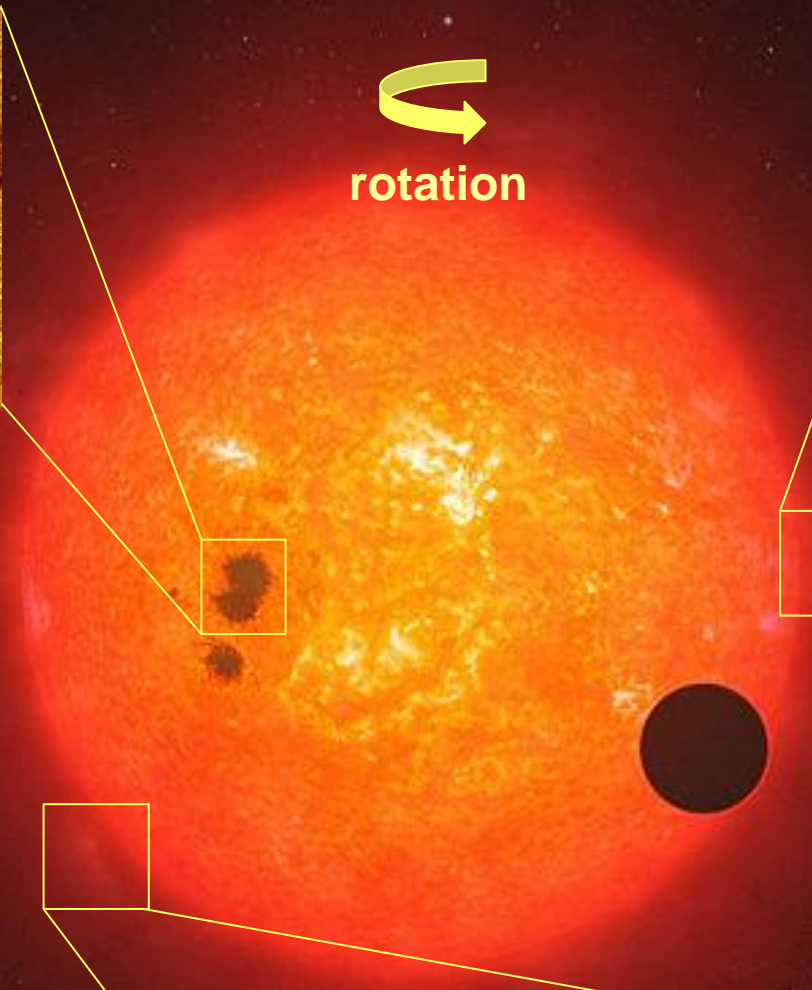


# Long-term monitoring of GJ1214s stellar variability

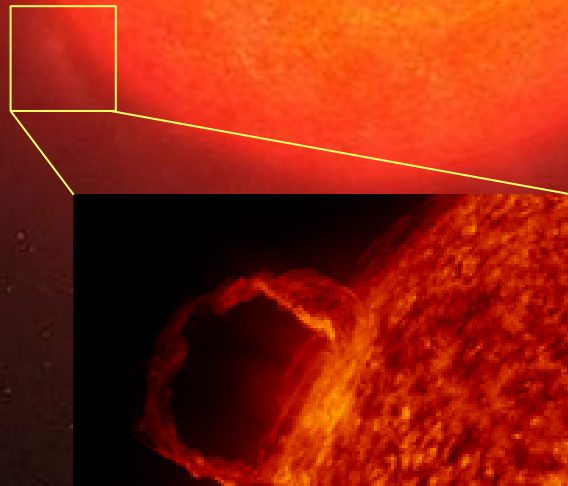


spots

rotation



flares

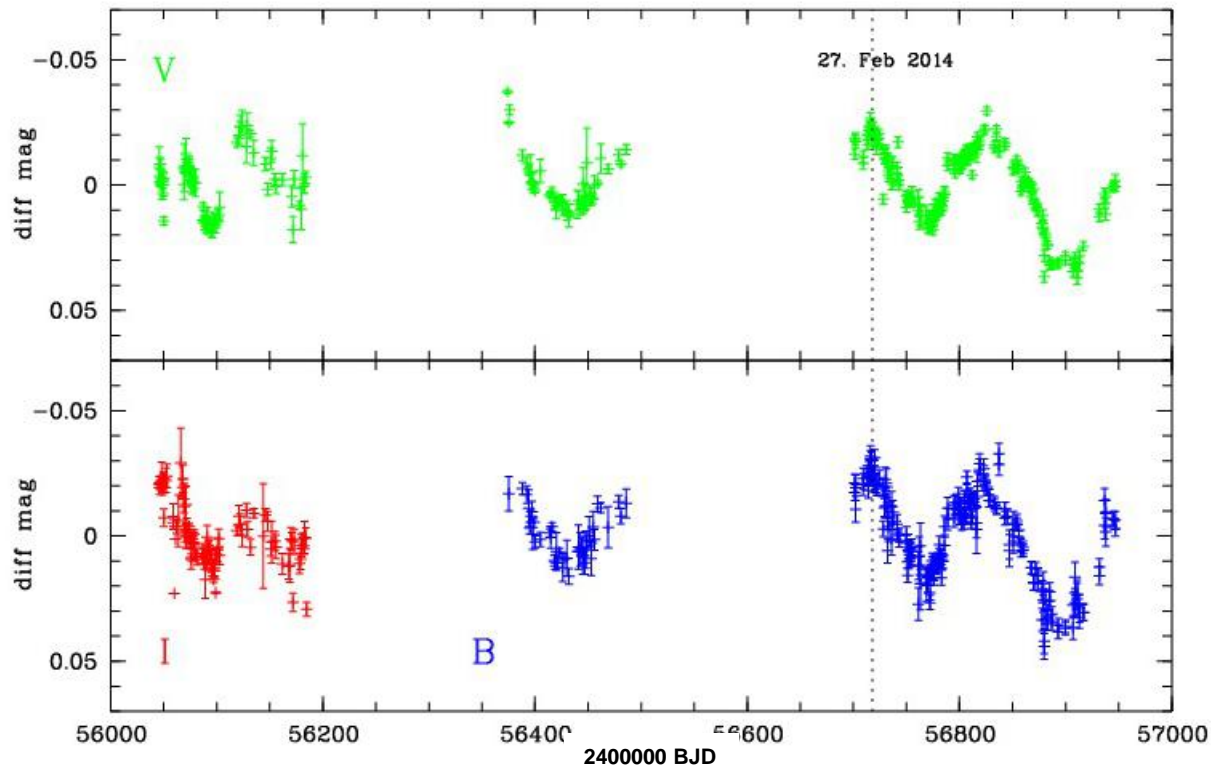


loops

# Long term monitoring of GJ 1214: stellar variability



- active star with period of  $\sim 80$  days (Charbonneau+ '09),  
 $\sim 44$  days (Narita+ '13);  $\sim 120$  days tbc (this work)
- stellar spots may mimic Rayleigh slope (Oshagh+ '14)



M. Mallonn



# Warm Uranus GJ 3470 b

## Host star GJ 3470

Right ascension ( $\alpha$ ):	07 <sup>h</sup> 59 <sup>m</sup> 06.0 <sup>s</sup>
Declination ( $\delta$ ):	+15° 23' 30"
Magnitude ( $m_V$ ):	12.27
Distance (d):	30.7 ( <sub>-2.1</sub> <sup>+1.7</sup> ) pc
Spectral type:	dM 1.5
Mass ( $m$ ):	0.539 (± 0.047) M <sub>☉</sub>
Radius ( $r$ ):	0.568 (± 0.037) R <sub>☉</sub>
Temperature ( $T$ ):	3600.0 (± 100.0) K
Age:	1.0 ( <sub>-0.7</sub> <sup>+0.7</sup> ) Gyr

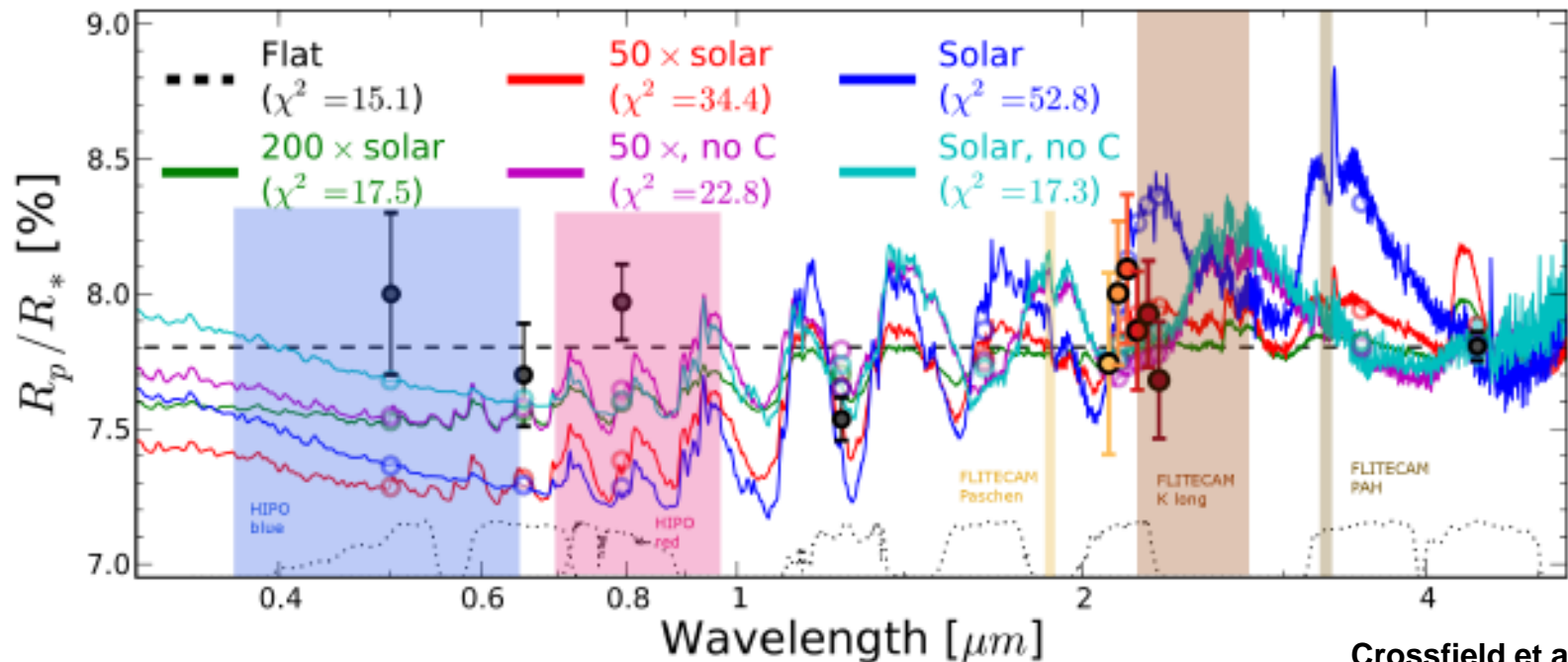
## Planet GJ 3470 b

Discovery:	2012 (Bonfils et al.)
Mass ( $m$ ):	0.044 (± 0.005) M <sub>J</sub>
Radius ( $r$ ):	0.37 (± 0.05) R <sub>J</sub>
Dichte ( $\rho$ ):	1.07 (± 0.43) g cm <sup>-3</sup>
Period (p):	3.33671 (± 5e-05) day
Transit duration (T):	~ 113 min
Semi major axis (a):	0.03557 (± 0.001) AU
Semi major axis (a):	0.03557 (± 0.001) AU
Transit duration (T):	~ 113 min

# Transmission spectrum of GJ 3470b

No detection of atmospheric features in the atmospheres (cf. GJ 1214b)

→ Determine its atmospheric structure and composition



- 3 optical channels (HIPO's red, blue filters and additional channel from FPI+)
- 1 infrared channel (FLITECAM 'Paschen alpha cont' filter)



# Observation of the primary transit of GJ 3470b

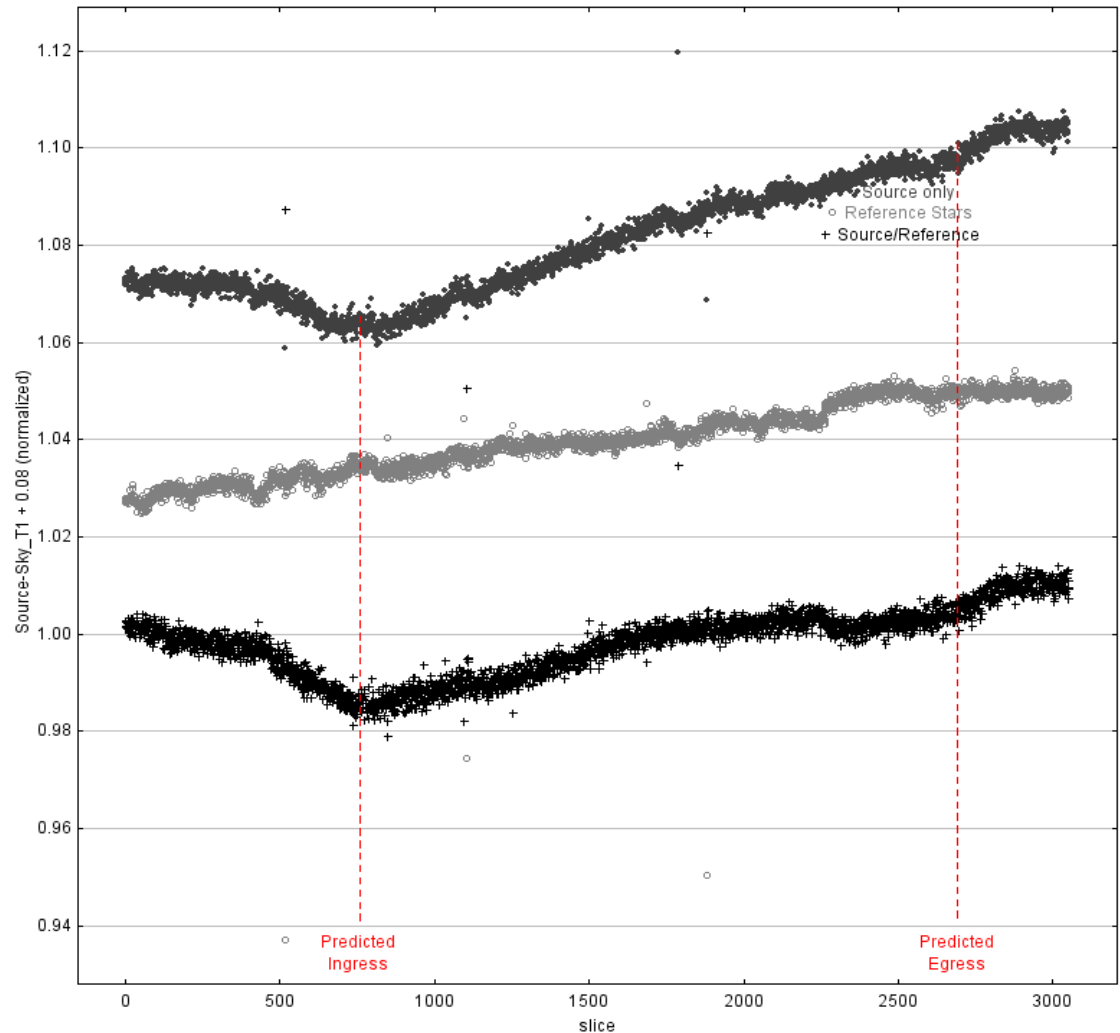
Take off date: 2015/09/29

FPI+ GJ3470b light curve

Top:  
signal of the target only

Middle:  
signal of the reference star

Bottom:  
differential between target  
and reference star





# Do starspots inflate the exoplanet CoRoT-2b ?

## Host star CoRoT 2

Right ascension ( $\alpha$ ):	19 <sup>h</sup> 27 <sup>m</sup> 07.0 <sup>s</sup>
Declination ( $\delta$ ):	+01° 23' 03"
Magnitude ( $m_V$ ):	12.57
Distance (d):	300.0 ( $\pm$ 100.0) pc
Spectral type:	G7V
Mass ( $m$ ):	0.97 ( $\pm$ 0.06) $M_\odot$
Radius ( $r$ ):	0.902 ( $\pm$ 0.018) $R_\odot$
Temperature ( $T$ ):	5575.0 ( $\pm$ 66.0) K
Age:	--

$\alpha$ :	--
Temperature ( $T$ ):	5575.0 ( $\pm$ 66.0) K
Radius ( $r$ ):	0.902 ( $\pm$ 0.018) $R_\odot$
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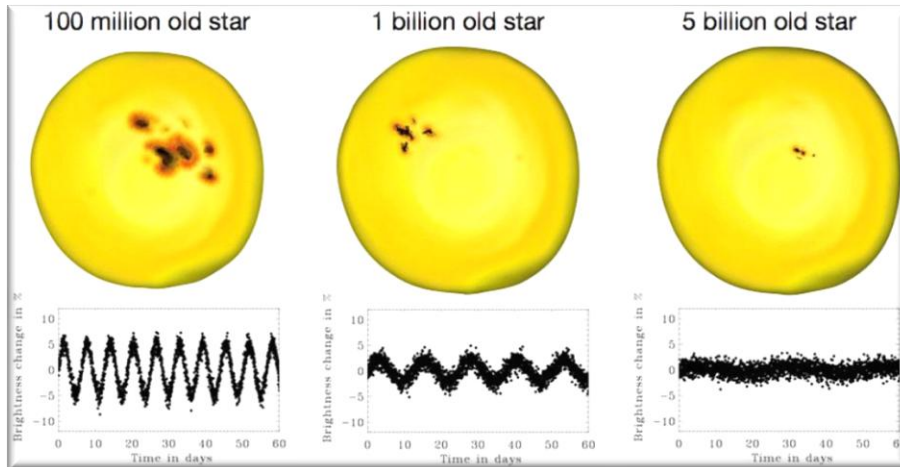
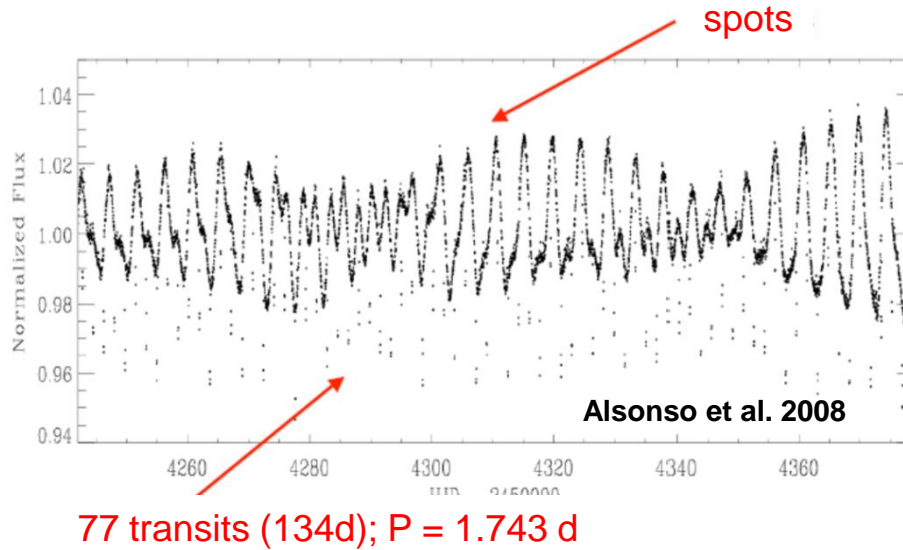
## Planet CoRoT 2b

Discovery:	2007 (Alonso et al.)
Mass ( $m$ ):	3.31 ( $\pm$ 0.16) $M_J$
Radius ( $r$ ):	1.465 ( $\pm$ 0.029) $R_J$
Dichte ( $\rho$ ):	1,310 ( $\pm$ 0,040) $g\ cm^{-3}$
Period (p):	1.7429964 ( $\pm$ 1.7e-06) day
Transit duration (T):	~ 125 min
Semi major axis (a):	0.0281 ( $\pm$ 0.0009) au

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Transit duration (T):	~ 125 min

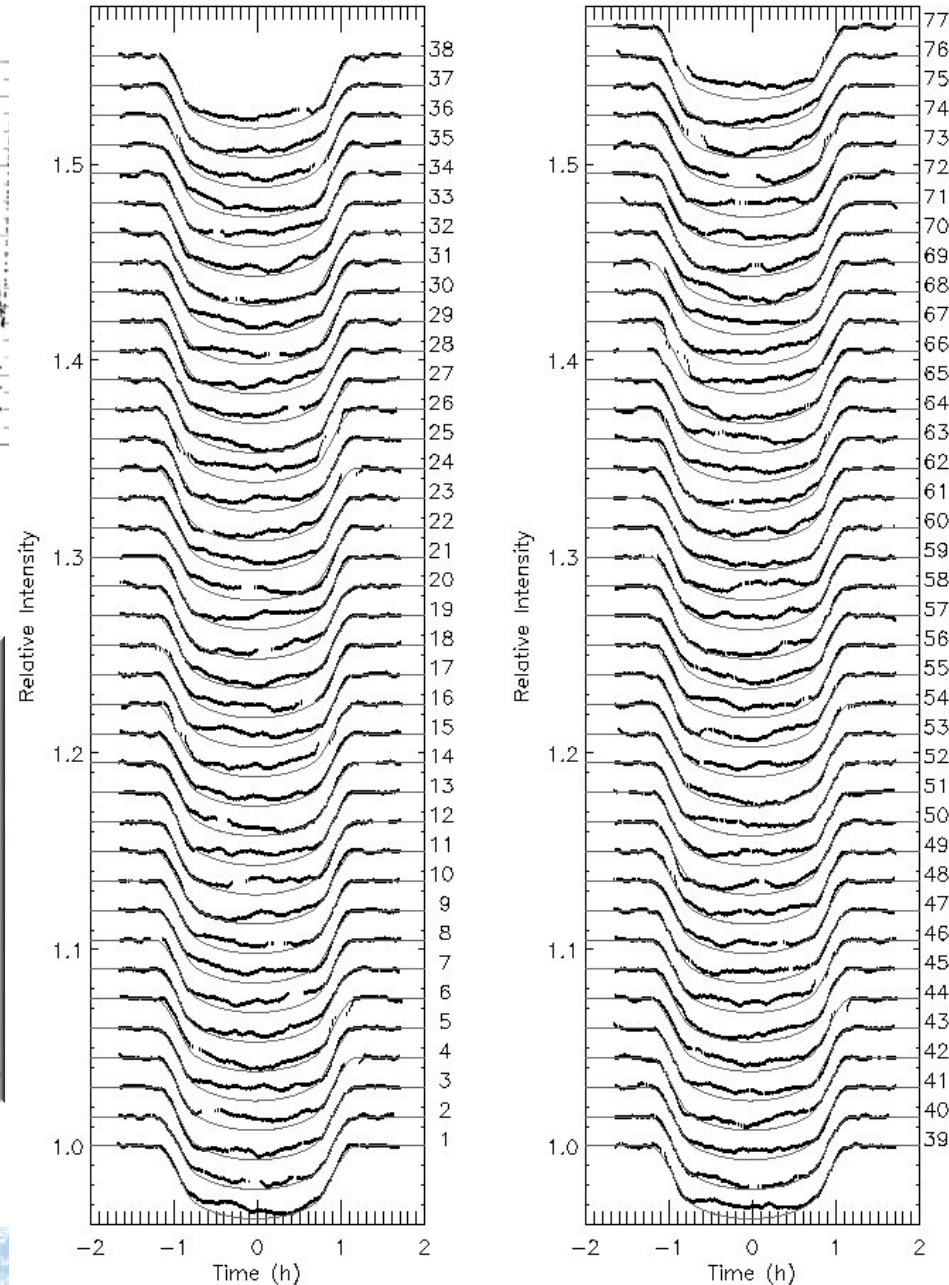
# Do starspots inflate the exoplanet CoRoT-2b ?

Silva-Valio et al. 2010

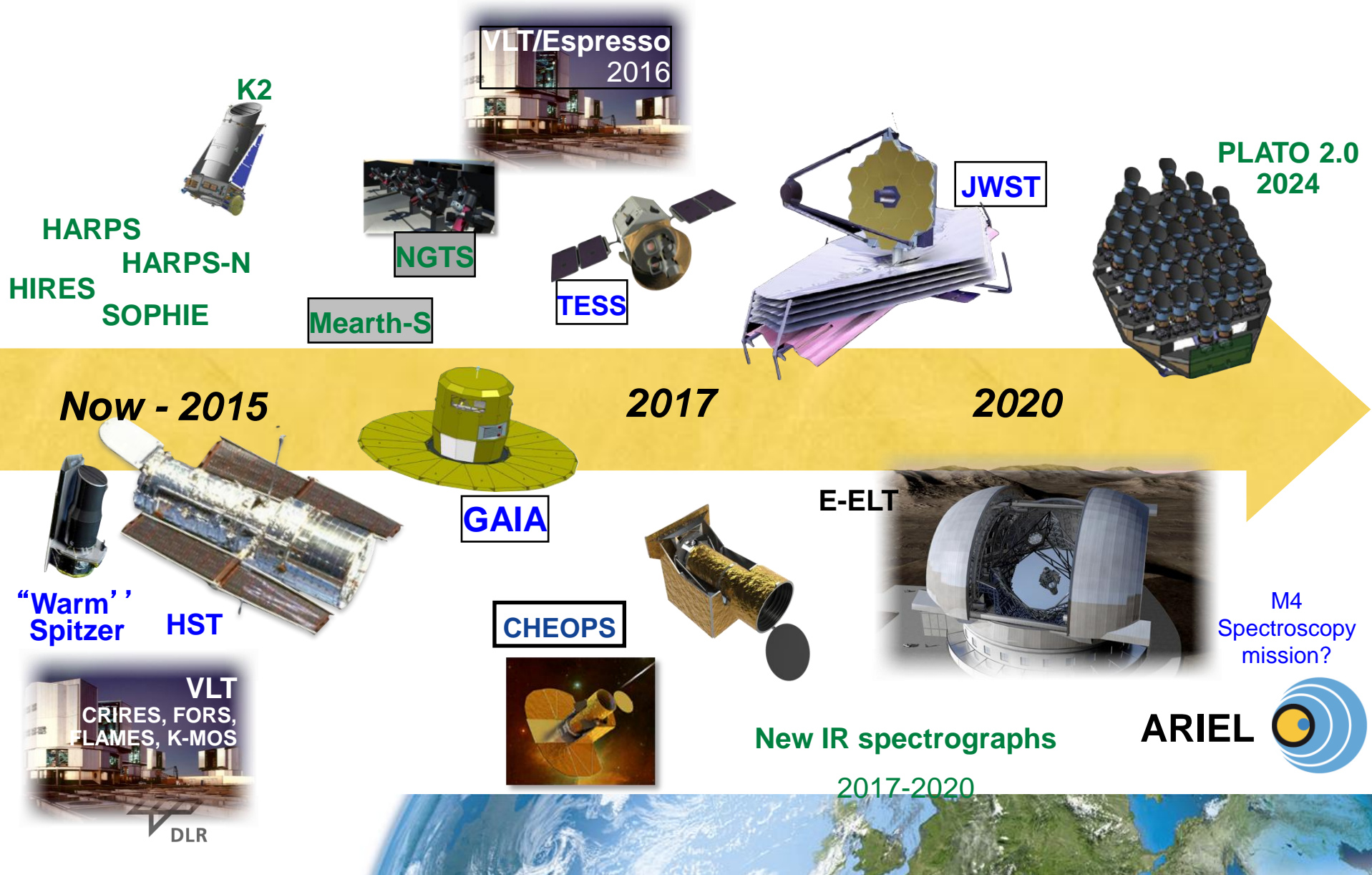


older stars rotate slowly & have fewer and smaller spots

DLR



# A time line of exoplanet characterization



**Thank you for your  
attention !**



**Have a safe flight**