



Solar System Studies with SOFIA

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Quick tour of possibilities









Terrestrial Planets



Mercury: not visible

Venus: atmosphere composition, dynamics (winds, temporal ation) [EXES]

Mars: atmosphere composition, dynamics









Outer Planets





- Moons with subsurface oceans
 - Radiometry for energy balance [FORCAST, FIFI LS, HAWC+]
 - temporal and phase-angle changes due to surface features ["]









Small Bodies



Dwarf Planets

- Occultations [HIPO/FLIPO] for diameters, rings, atmosphere search, haze
- Radiometry [FORCAST]: diameter, thermal properties
- Surface composition [FORCAST & FLITECAM grism]



Comets: like asteroids, plus dust and gas composition from outgassing [GREAT, FORCAST grism]



mineralogy (Fe/Mg silicates) origin of terrestrial water (D/H, ortho/para)









More details on a few potential projects









Venus

- Visibility conditions on mid-Feb2017:
 - 40 degree solar elongation
 - 40" angular size
 - but only visible in early evening for <1 hr
- Molecules already observed with TEXES
 - CO₂, HDO and SO2
 - Detectability of lines depends critically on Doppler shift
- Other molecules remain uncharacterized
 - no access to infrared outside Earth's atmosphere, while Venus has many of same gases as Earth
 - Isotopic ratios trace chemical history; why were Venus' oceans lost?
- Dynamics
 - Vertical distribution of SO, SO₂ related to cycling
 - Winds
 - Factor of 5-10 temporal variation in SO₂ (Encrenaz et al 2012)









Jupiter's Stratosphere and Troposphere



SOFIA spectroscopic slits were stepped across the disk of Jupiter to make spectral maps.

With EXES at high spectral resolution: lib-brightened, narrow, stratospheric H₂ line.

With FORCAST at moderate resolution, measure pressurebroadened H₂

Para/Ortho state ratio, paired with CH₄ temperature, measures mixing rate in the atmosphere, because the rate of para/ortho conversion is known.









Image of Titan, which is covered in hydrocarbon haze

Hydrocarbon lakes found by Huygens spaceprobe

- Central flash only present for longest wavelength
 - Extinction and refraction can be separated because of their different wavelength dependence
 - Aerosol scattering is much stronger at shorter wavelengths, depends on $(a/\lambda)^4$
- Multi-wavelength observations are critical for understanding atmosphere USRA
 Zalucha et al. (2007): from Eliot Young

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Solar System Observation Notes









Tracking

- Rate should be slower than 1"/s
- "easiest" tracking for visibly-bright objects (V<14), better than 1"
- Invisible targets possible using offset guiding on celestial source, but this is a not-well-tested mode





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- The telescope stays in the elevation range 22-58 degrees
- Absolute minimum solar elongation is therefore 22 degrees
 - Practical limit of 25 degrees to allow time to acquire & observe
- We have flight rules to prevent sunlight on the telescope
- Observations right after sunset are possible
 - No time for initial calibration leg, so not best for 1st flight of series
 - Takeoff before sunset, door open at 10,000 feet, get on target
- Observations up to sunrise are possible but have risk
 - Need to close door and prepare alternate landing scenario such that a potentially "stuck open" door is pointed away from the Sun





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Moons

- Keeping main planet off the array (or slit) is highly desirable
- Horizons and XEphem are good for visualizing orientations









SOFIA Focal Plane



- View from Spot
- EXES Medium
 - X-dispersed slits are short
- FLITECAM
 - Horizontal slit
- FORCAST
 - Vertical slit
- FIFI LS
 - Blue channel ½ area

