

Stratospheric Observatory for Infrared Astronomy

HAWC+ Upgrade Roadmap

SOFIA Instrumentation Workshop

June 22, 2020

Michael Gordon

michael.s.gordon@nasa.gov

HAWC+ Upgrade Paths

- Minor upgrades (~12-18 months)
 - Narrow-band filters
 - Tuning/optimization of cryo parameters
 - New observing modes
 - 4th array (?)

- Major upgrades (2-3 years)
 - 4 new arrays
 - BLAST-TNG kinetic inductance detector arrays









Narrow-Band Filters

Enabling new science

- HAWC+ filter wheel currently has one open aperture
 - two if we remove Band B (Cycle 9 usage dependent)













• Potential options include [O I], [O II], [C II], and OH













- Synergies with other SOFIA instruments
 - GREAT, FIFI-LS
 - emission kinematics in [C II] or [O I]















- Synergies with other SOFIA instruments
 - GREAT
 - emission kinematics in [C II] or [O I]
 - Magnetic fields are aligned and compressed in PDRs in high gas and radiation pressure environments
 - Combine narrow-band polarimetry with continuum emission (~40 K) to trace both gas and dust









- Synergies with other SOFIA instruments
 - GREAT, FIFI-LS

[O I] 145 µm













ADR Tuning

Better sensitivity

HAWC+ ADR Tuning

Detector Temp (K)	ADR initial current (A)	Expected initial current in flight (A)	Expected final current of a 10h flight (A)*	Expected hold time in flight (h)	Comments
0.20	2.02	1.72	0.47	13.8	Nominal ADR w/old HS.
0.17	2.02	1.34	0.09	10.7	Nominal ADR w/old HS.
0.15	2.03	1.19	-0.06	9.5	Nominal ADR w/new HS. Runout of ADR current by the end of flight
	2.46	1.54	0.29	12.3	Potential new temperature
0.14	2.46	1.41	0.16	11.3	Potential new temperature
	2.36	1.31	0.06	10.5	Potential new temperature
	2.41	1.36	0.11	10.9	Potential new temperature
0.13	2.00	0.94	-0.31	7.5	Runout of ADR current by the end of flight
	2.46	1.26	0.01	10.1	Runout of ADR current by the end of flight



DLR



New Observing Modes

Improved efficiencies















Issues:

• ~50% observing efficiency





ssues:

 ~50% observing efficiency • For polarimetry: Chop-nod @ 4 HWP positions







Issues:

- ~50% observing efficiency
- Requires clean "off" position over the entire FOV











HAWC+ Scanning Modes

- SOFIA/HAWC+ scan mapping has two modes:
 - <u>Lissajous</u>: for small fields comparable to the FOV
 - <u>Raster</u>: for large mapping areas

• Scan speeds up to 200 arcsec/s



ScanPol Mode

- Scan-mapping:
 - extremely efficient for total intensity imaging of extended and point sources
- Scan-polarimetry:
 - now offered as shared risk observing mode for SOFIA
 Cycles 8 & 9 (2019-2021)
 - Great for point sources. Accurate pol angle for all spatial scales
- Ongoing Work
 - Re-characterize instrumental polarization in scan mode
 - Correct polarization bias in low surface brightness regions







Imaging Polarimetry of 30 Dor

Dec (J2000)



New Arrays

Improved mapping speed, sensitivities, FOV

Current Instrument Parameters

Band / Wavelength	Δλ/λ	FWHM (arcsec)	Total Intensity FOV (arcmin)	Polarization FOV (arcmin)	
A / 53 µm	0.17	4.7	2.7 × 1.7	1.3 × 1.7	
B / 63 μm	0.15	5.8	4.2 × 2.6	2.1 × 2.6	
C / 89 µm	0.19	7.8	4.2 × 2.6	2.1 × 2.6	
D / 154 µm	0.22	14	7.3 × 4.5	3.6 × 4.5	
E / 214 µm	0.20	19	8.0 × 6.1	4.0 × 6.1	



DLR

USR



Current Instrument Parameters

Band / Wavelength	Δλ/λ	FWHM (arcsec)	Total Intensity FOV (arcmin)	Polarization FOV (arcmin)	
A / 53 µm	0.17	4.7	2.7 × 1.7	1.3 × 1.7	
B / 63 μm	0.15	5.8	4.2 × 2.6	2.1 × 2.6	
C / 89 µm	0.19	7.8	4.2 × 2.6	2.1 × 2.6	4
D / 154 µm	0.22	14	7.3 × 4.5	3.6 × 4.5	
E / 214 µm	0.20	19	8.0 × 6.1	4.0 × 6.1	









PACS / 60-200 μm ~5–12 3.5 × 1.8

HAWC+ Imaging Polarimetry





















• **~54%** of the pixels on the R0, R1, T0 arrays are *unusable*























Additional Arrays

- Four arrays (two in R, two in T) allows for doubling our FOV in polarimetry
- For scanpol, this effectively doubles our scan-speed

- Caveats:
 - Requires fabrication
 - Uncertain if Goddard has maintained that capability
 - Unknown expense/timeline
 for 4 high-quality 32x32 arrays



DIR

USRA

HAWC+ Upgrade Paths

- Minor upgrades (~12-18 months)
 - Narrow-band filters
 - Enables new science
 - Cryo tuning/new observing modes
 - Sensitivity and overhead improvements
 - 4th array
 - Increase mapping speed and FOV

- Major upgrades (2-3 years)
 - 4 new arrays
 - BLAST-TNG kinetic inductance detector arrays









