

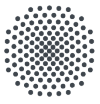
Universität Stuttgart

Institut für Raumfahrtssysteme (IRS)  
Deutsches SOFIA Institut (DSI)

# Upgrade of the SOFIA Focal Plane Imager FPI+

## VIS & NIR Guider and Science Camera FPI++

Enrico Pfueller, DSI



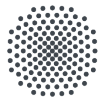
## Expand the tracking capability of the SOFIA telescope into the NIR up to about 1.6 $\mu\text{m}$

- Far-infrared observations in regions of large extinction would profit from NIR tracking, as the number of available tracking stars increases towards longer wavelengths
- Increase sky coverage to 100% in dark clouds (99% all sky)

## FPI+ NIR spectral coverage enables science applications requiring Y-, J- and H-band photometry

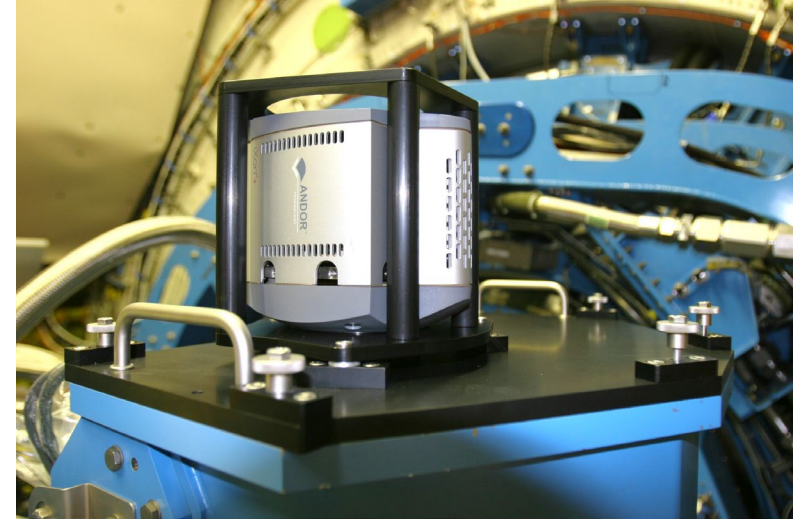
- Multi-channel capability necessary to study particle sizes and aerosol compositions during stellar occultations by atmospheres, rings or cometary coma

Motivation

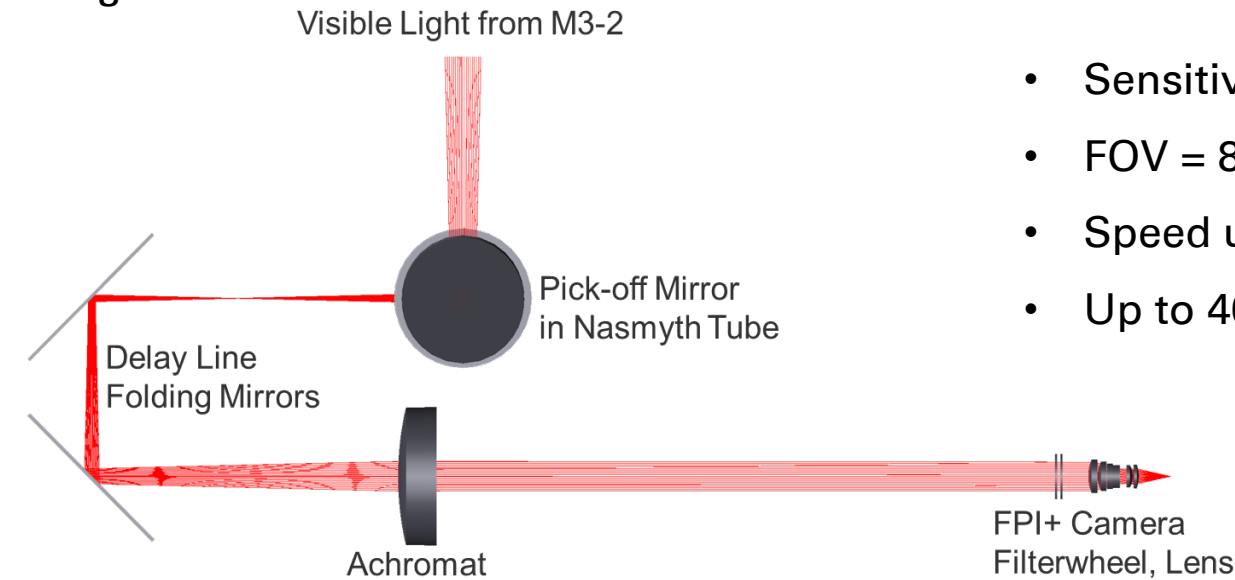


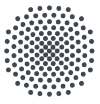
## The Focal Plane Imager (FPI) is SOFIA's primary tracking camera

- Original camera upgraded in 2013 with commercial EM-CCD camera → FPI+
- SOFIA Facility Science Instrument since cycle 4
- Main application: stellar occultations by small solar system bodies (Pluto!)
- Permanently mounted at the SOFIA telescope;  
available in each mission in addition to any science instrument on the main flange



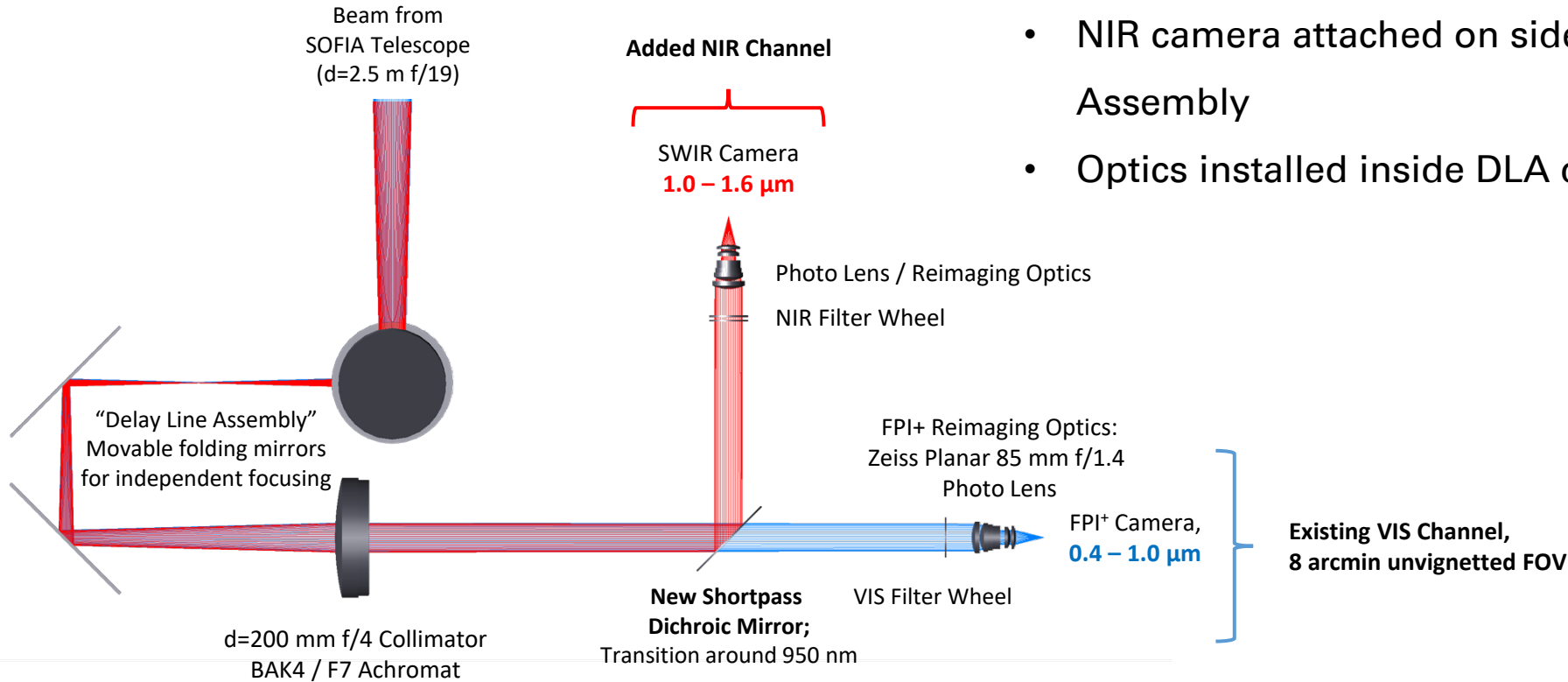
- Sensitivity: 16<sup>mag</sup> with  $S/N = 35$  in 1 sec
- FOV = 8 arcmin diameter
- Speed up to 10 fps full frame
- Up to 400 fps with subframes / binning





## FPI+ Near-Infrared Upgrade Approach

- NIR capability for tracking  $1\ \mu\text{m}$  to  $1.6\ \mu\text{m}$
- Retain current VIS FPI+ sensitivity with moveable dichroic mirror



- NIR camera attached on side of Delay Line Assembly
- Optics installed inside DLA distance tube





## Why tracking in the NIR?

- Difficulty to find suitable guide stars in dark clouds; Interesting IR objects „hide“ in VIS dark clouds
- There have been observations (HAWC+, FIFI-LS) that could not be scheduled due to missing guide stars



**Object: L 1688**

RA: 16:28:51.5, Dec: -24:17:03

Left: VIS no stars brighter than  
R = 16 mag

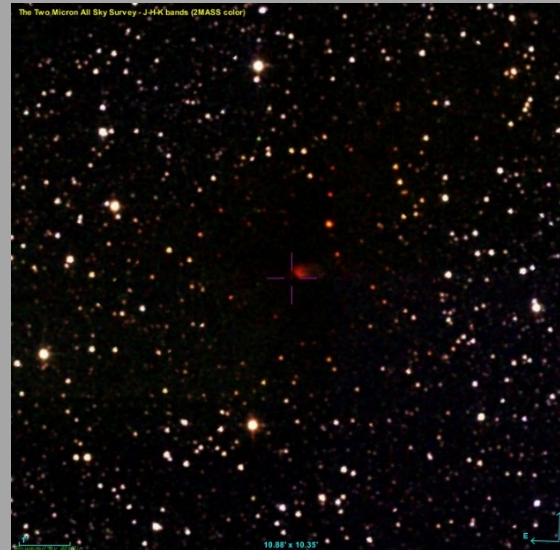
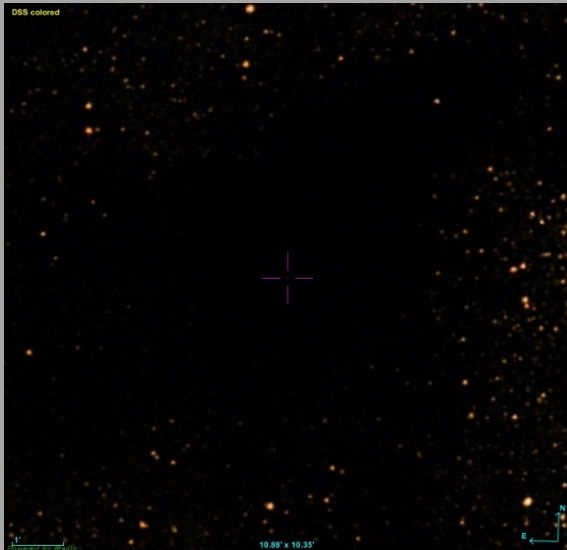
Right: NIR star with H = 11.4 mag  
and about 10 stars brighter than  
H = 14 mag.

NIR Tracking



## Dark Cloud and All-Sky Coverage Analysis

- 2892 dark clouds (Lynds, Hartley) analyzed: 96% have V=14mag guide star within FPI+ FOV (UCAC4)  
100% have H=14mag guide star (2MASS)
- Number of NIR guide stars in all dark clouds one magnitude larger than VIS guide stars
- All-sky coverage VIS 90%, NIR 99%



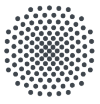
**Object: L 483**

RA: 18:17:29.9, Dec: -04:39:41

Left: no suitable VIS stars for tracking

Right: NIR star with H = 10.4 mag and about 50 stars brighter than H = 14 mag.

NIR Tracking

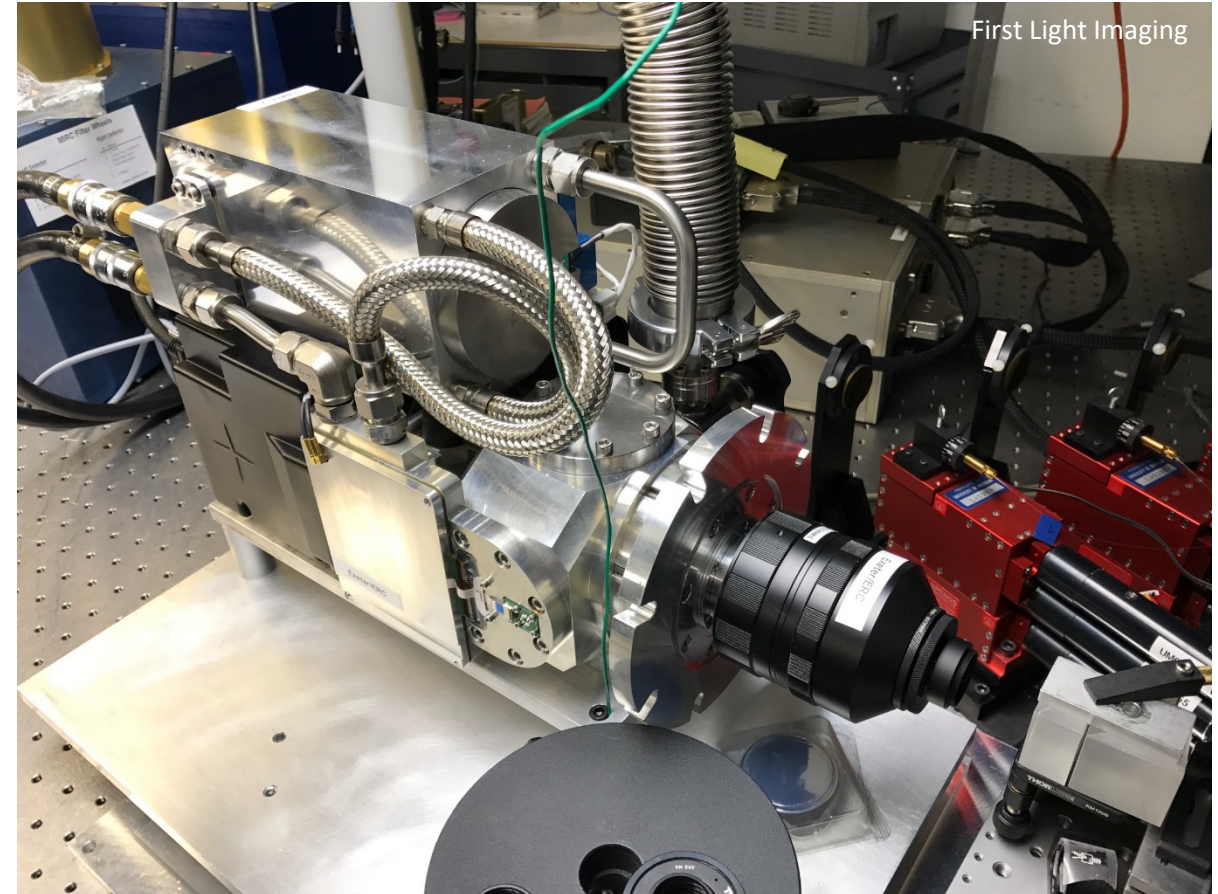


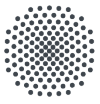
## Promising commercial solution for NIR upgrade

### First Light Imaging „C-RED One“ Camera

- Leonardo Saphira e-APD HgCdTe Array
- 320 x 256 pixel, 24  $\mu\text{m}$  pitch – sweet spot for existing FPI+ optics and shear layer seeing
- Pulse tube cooled to 80 K = no consumables
- Very low noise figures
- Optimized towards high frame rates

Plate scales 1.3 ... 1.6 arcsec/pixel appear feasible

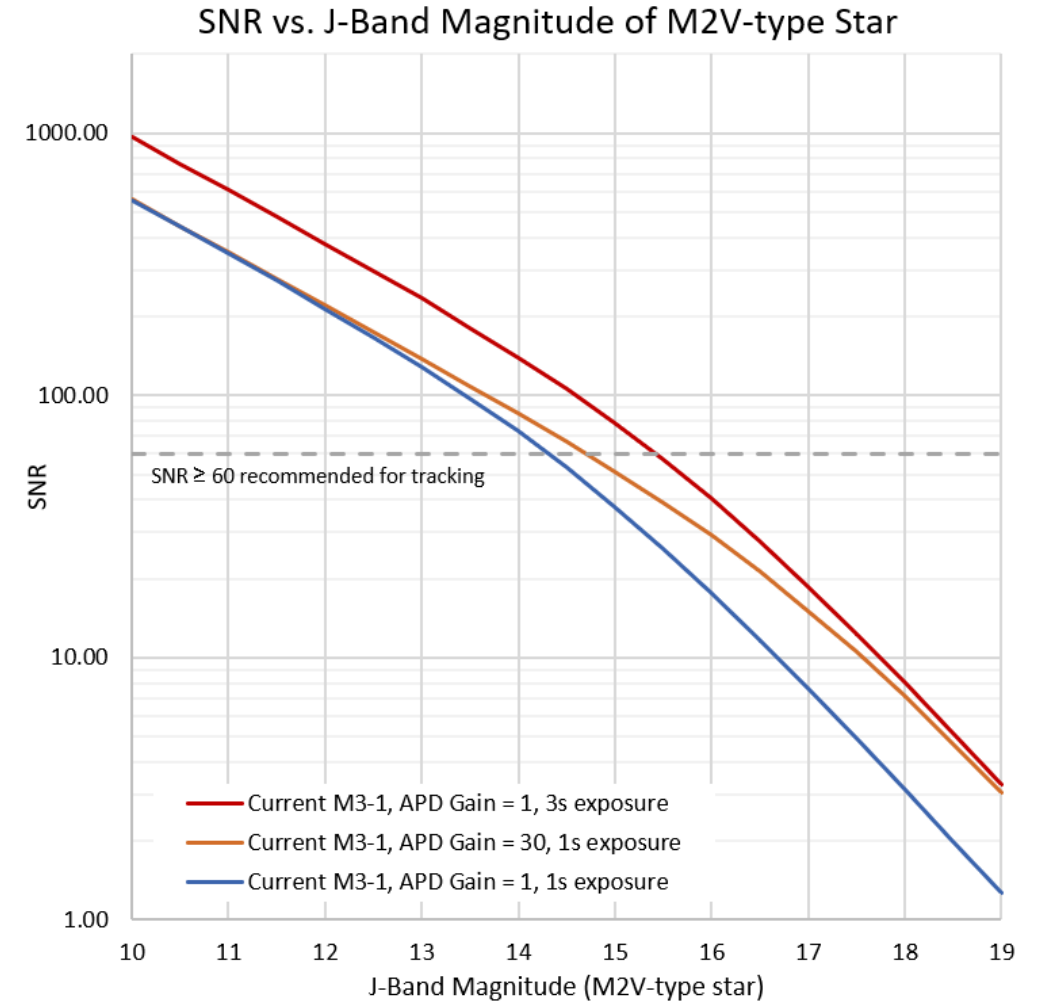
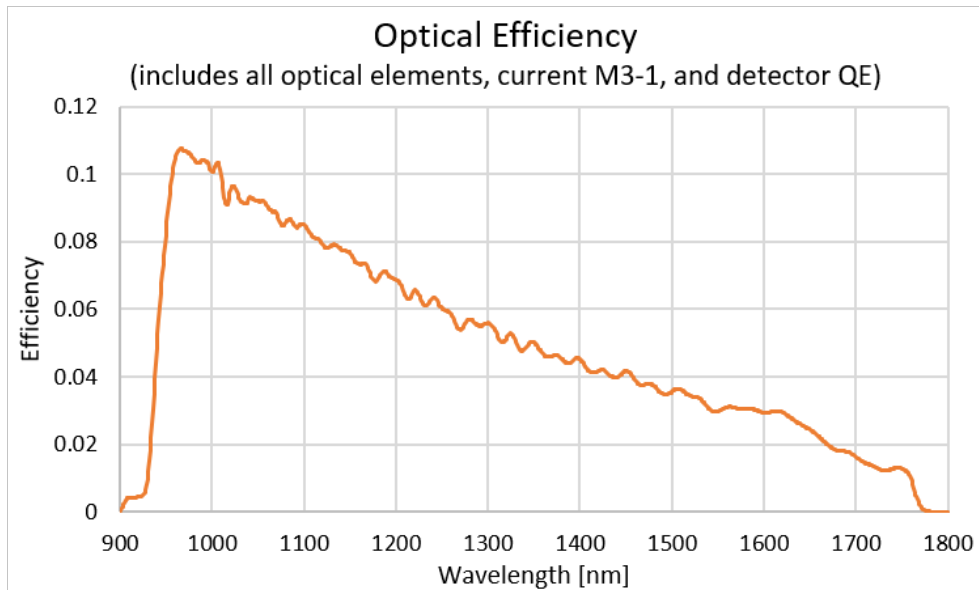




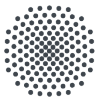
## Sensitivity analysis of proposed NIR channel

	SNR = 60 (tracking limit)	SNR = 10 (detection limit)
1 s exposure	J=14.0 mag	J=16.3 mag
3 s exposure	J=15.1 mag	J=17.4 mag

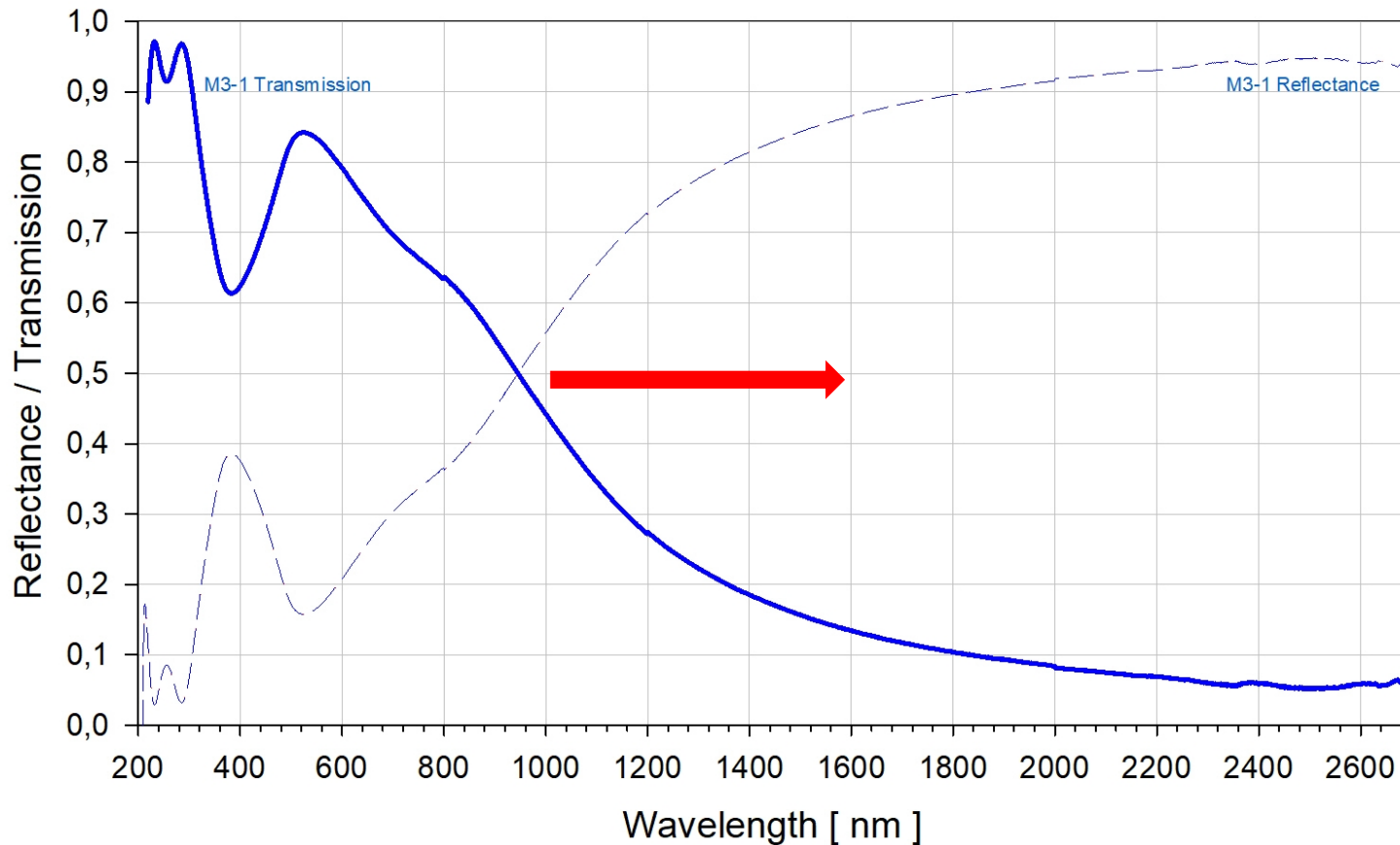
Assumptions: EM off (APD gain = 1), M2V type star







## Limiting NIR transmission of dichroic tertiary mirror M3-1



Improved M3-1 requirements:

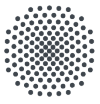
- Shift 50% transmission to  $\lambda > 1.6 \mu\text{m}$
- Keep FIR reflectance  $> 95\%$

Very promising new coating is under investigation

Benefits:

- 2 stellar magnitudes improvement of FPI++ NIR sensitivity
- M3-1 spare part

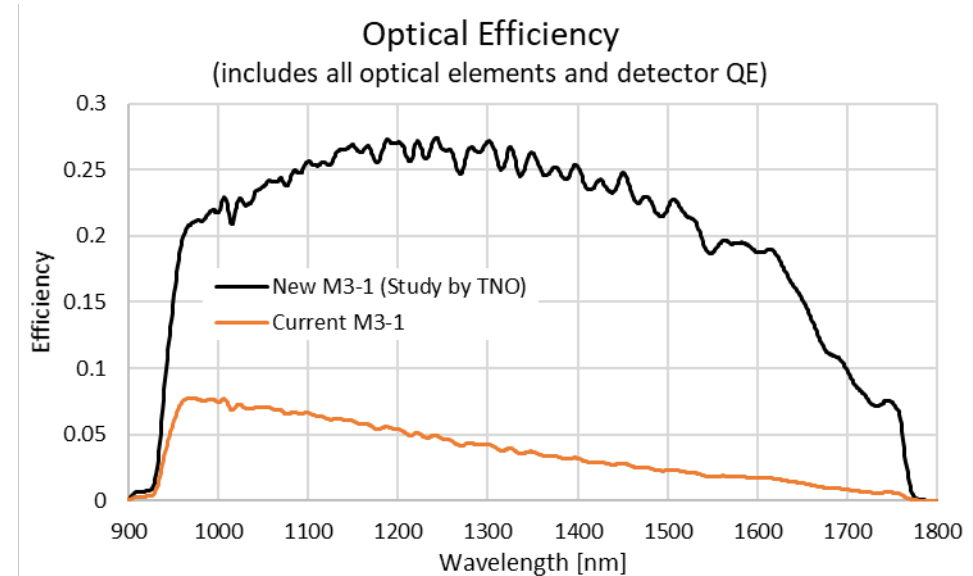
Improved  
Tertiary M3-1  
Mirror

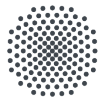


## Sensitivity analysis of proposed NIR channel

Current M3-1	SNR = 60 (tracking limit)	SNR = 10 (detection limit)
1 s exposure	J=14.0 mag	J=16.3 mag
3 s exposure	J=15.1 mag	J=17.4 mag
Upgraded M3-1	SNR = 60 (tracking limit)	SNR = 10 (detection limit)
1 s exposure	J=15.8 mag	J=18.1 mag
3 s exposure	J=16.9 mag	J=19.2 mag

- matches or exceeds Andor iXon 888 at comparable V-mag
- does not even consider further sensitivity improvements using avalanche multiplication



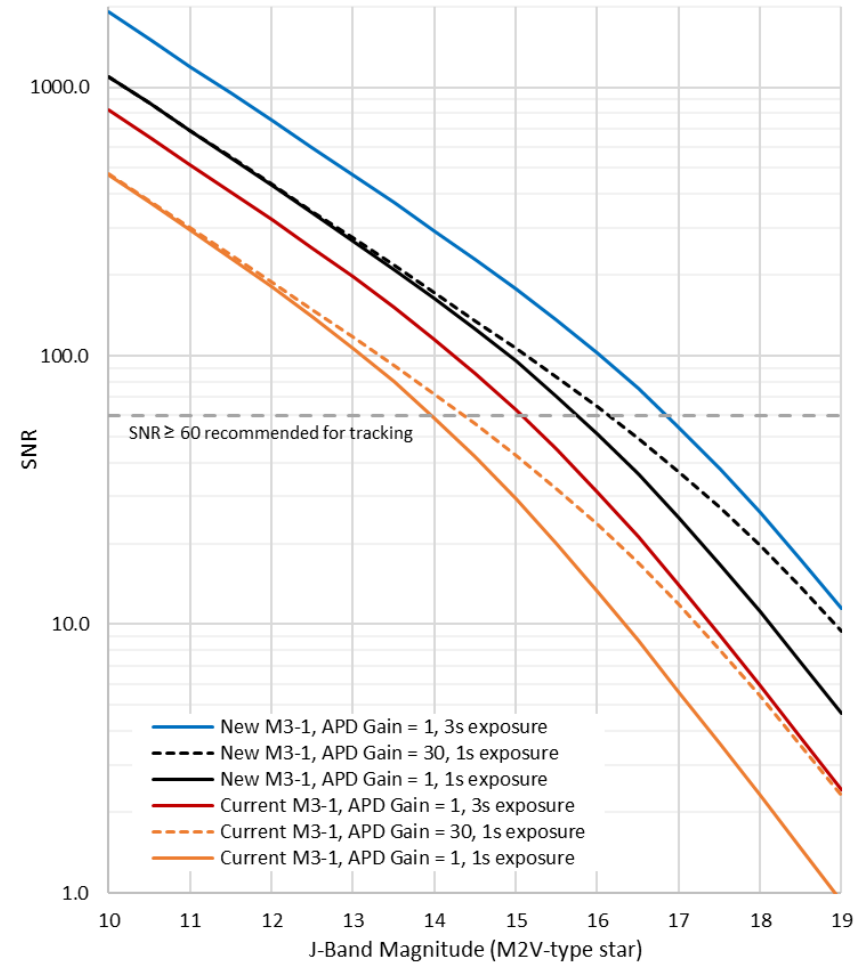


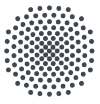
## Sensitivity analysis of proposed NIR channel

Current M3-1	SNR = 60 (tracking limit)	SNR = 10 (detection limit)
1 s exposure	J=14.0 mag	J=16.3 mag
3 s exposure	J=15.1 mag	J=17.4 mag
Upgraded M3-1	SNR = 60 (tracking limit)	SNR = 10 (detection limit)
1 s exposure	J=15.8 mag	J=18.1 mag
3 s exposure	J=16.9 mag	J=19.2 mag

- matches or exceeds Andor iXon 888 at comparable V-mag
- does not even consider further sensitivity improvements using avalanche multiplication

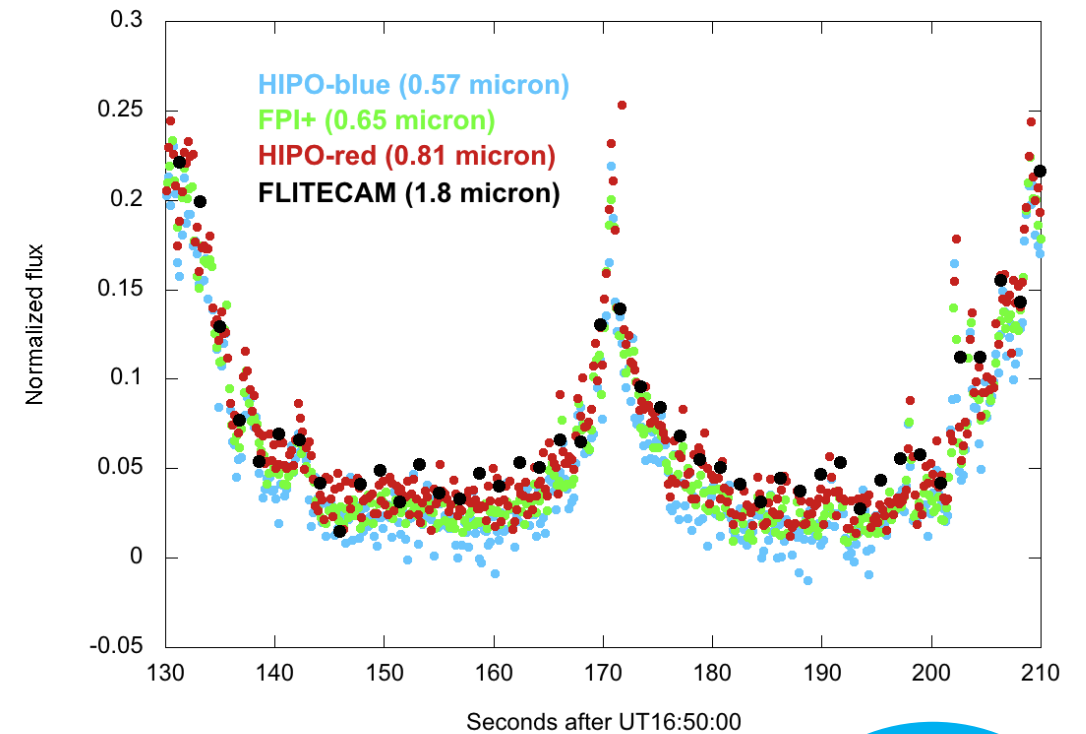
SNR vs. J-Band Magnitude of M2V-type Star



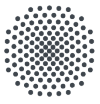


## Science Applications of NIR-upgraded FPI++

- The upgraded FPI++ will partially restore the capabilities of the former instrument combination FLIPO (HIPO + FLITECAM)
- Allows for high-speed observations at two wavelengths simultaneously (VIS & NIR)
- Critical for stellar occultations by bodies with an atmosphere, e.g. Pluto, Triton, Titan
  - Atmospheric extinction is wavelength dependent
  - VIS and NIR light is therefore scattered differently in atmospheres
    - Existence/Depth of hazes can be shown
    - Aerosol particle sizes can be determined
    - Temperature gradients can be verified



Multi-Wavelength Occultations



- The upgraded Focal Plane Imager FPI++ will enable tracking in the NIR, up to  $\sim 1.6 \mu\text{m}$
- This results in practically 100% sky coverage with usable guide stars even in dark clouds
- The upgraded Focal Plane Imager FPI++ will enable dual wavelengths observations (VIS, NIR) particularly useful for stellar occultations by solar system bodies with atmospheres

A two-step approach is possible:

### Step 1

Mechanical and optical design for a second VIS camera

Possible benefit for tracking with deep depletion CCD ( $\sim 1 \mu\text{m}$ )

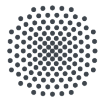
Would fully prepare the installation of the NIR camera and enable dual wavelength observations (VIS red & VIS blue)

Development time  $< 2$  years

### Step 2

Replace the second VIS camera with NIR camera when available ( $\sim 1$  year lead time) and characterized in the lab





Universität Stuttgart

Institut für Raumfahrtssysteme (IRS)  
Deutsches SOFIA Institut (DSI)

