Preparing your SOFIA proposal

Science case / SOFIA motivation

Gather support documents/tools

Evaluate feasibility

Define observations in USPOT

Build / upload justification pdf

th

Submit! Deadline Sept 6 , 21h PDT





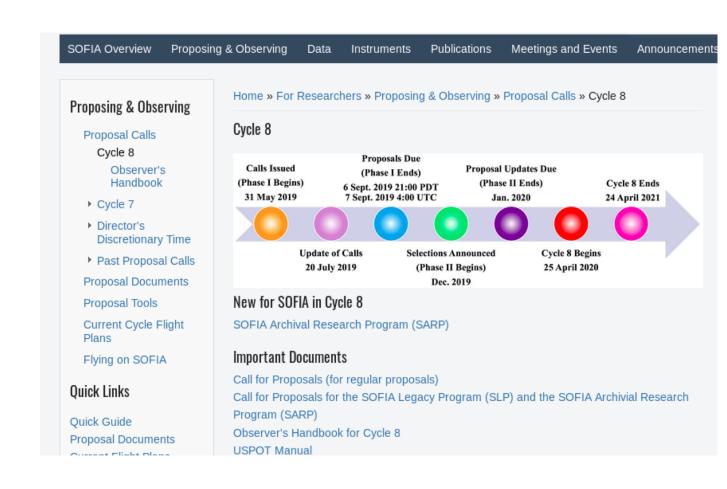
Gather relevant documentation and tools: Call for Proposal

What's offered - in short:

- All instruments
- 2 instruments deployed to NZ
- 'suitcase' deployments can be considered

New:

- improved mapping modes on FORCAST (slit-scan) and HAWC+ (OTF polarimetry)
- new filter for FIFI-LS improving sensitivity at the [OIII] 52μm line
- new HAWC+ band B (63μm) available







Gather relevant documentation and tools: Call for Proposal

3 different types of proposals

- **Regular**: up to 300h tot, includes surveys, ToO
- **Legacy:** 1-4 large proposals, up to 200h observations each, over 2 cycles. Up to 400 h total in C8. No proprietary period, team contributes enhanced products
- Archival Research: new (funding available, for US institutions only)





Gather relevant documentation and tools: Call for Proposal

Funding opportunities (for US-institutions only)

- Up to \$3M for Regular Proposals (~\$10k/ h)
- Up to \$2M / year for Legacy Proposals
- ~ \$300k for archival research proposals

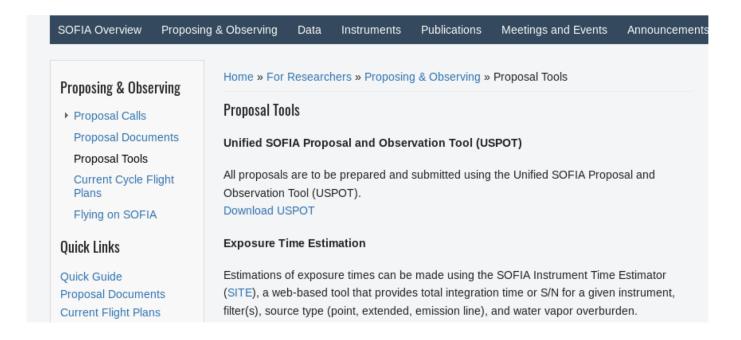
	Regular Proposals	Legacy (SLP)	Archival (SARP)
Exclusive Use Period	yes	no	n/a
Observing Period	1 cycle	2 cycles	n/a
Total Program Budget	\$3M	\$2M/yr	\$300k
Proposals Selected	varies	1-4	10

- For proposals which are central to a PhD thesis, additional funding can be requested through the <u>Thesis-enabling Program</u> (up to two years of graduate student funding)
- Timing of funding depends on program priority ('rank'): must do (\sim 25% of available time), should do (50%), do if time





Gather relevant documentation and tools



SOFIA Instrument Time Estimator (SITE)

Please Check 'Notes and Known Issues' Before Proceeding

Spectroscopic Time Estimators and Tools

FIFI-LS FORCAST GRISM FLITECAM GRISM GREAT EXES

Imaging Time Estimators

FORCAST FLITECAM FLITECAM_HIPO HAWC_Plus FPI_Plus

Observers' Handbook

USPOT manual

Download USPOT

Open SITE (website)

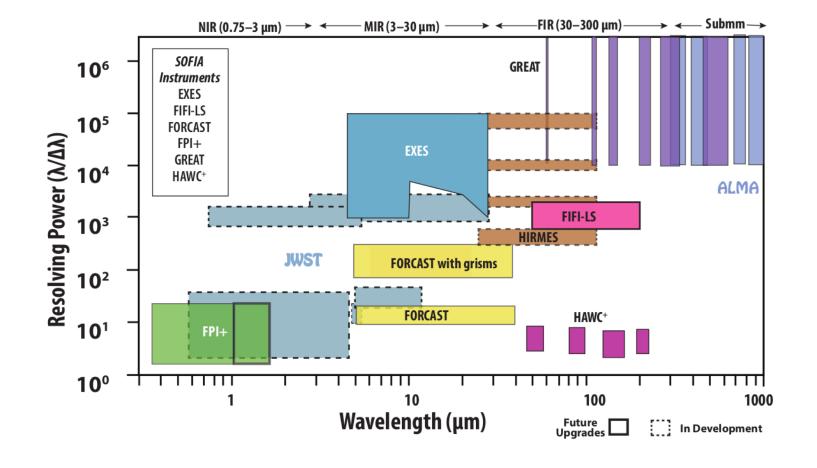
Any missing information: Help Desk





Feasability

SOFIA currently covers most of the Mid and Far-IR spectrum (5-600 μ m), at a variety of spectral resolutions







Feasability: Archival search and reserved observations catalogues

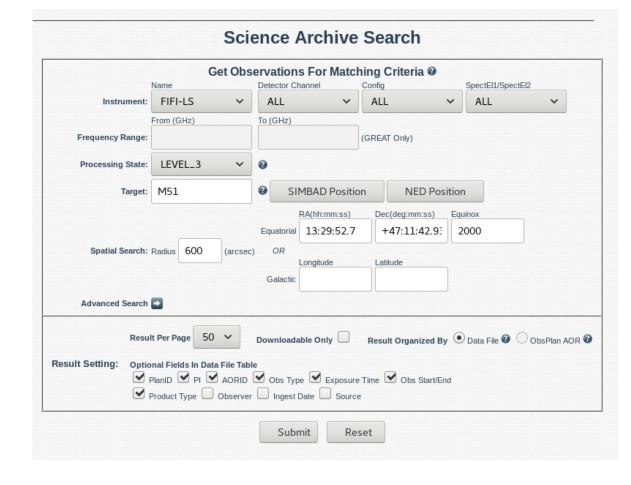
Reserved observations (GREAT and FIFI-LS): duplicates from Reserved Observations Catalog (ROC) lists not allowed, unless explicit permission from the instrument's PI (SMO should be notified prior to proposal submission)

Duplication of observations:

generally not allowed, and if proposed for must be identified and explicitly justified.

Check the SOFIA archive (registration required)

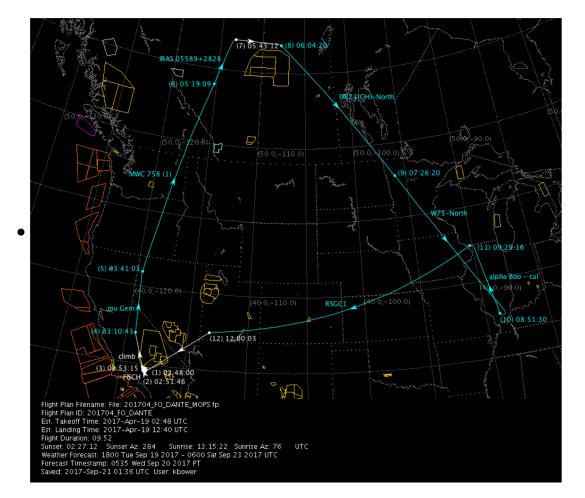
https://dcs.arc.nasa.gov/







Timing limitations considerations



Observing legs are limited (~ 2 hour max: hard to get long continuous observations)

Observations with an instrument limited to 1 / 2 obs. series per cycle

GC center observations in Northern early fall

Observations from NZ in Northern Summer





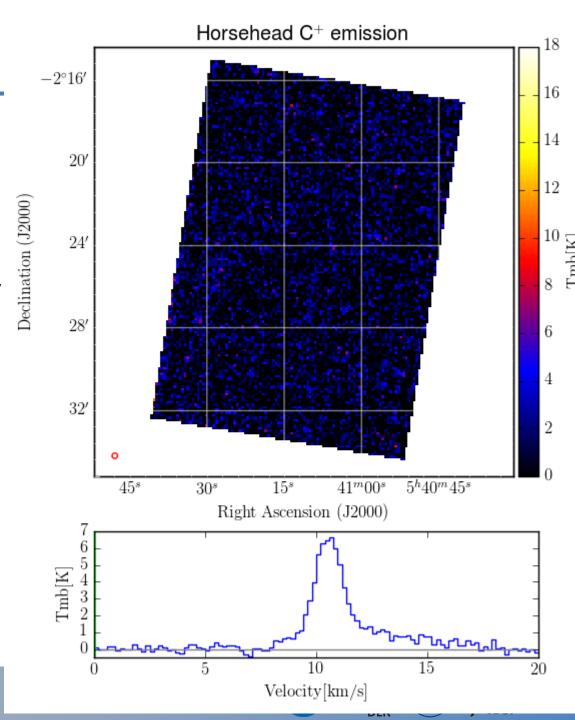
Estimation of Signal and Signal to Noise

Expected source signal: needs to be soundly justified!

- archival data (possibly w. SED extrapolation)
- your own radiative transfer model (describe)
- classic models/ correlations: Hyperion (dust), Planetary Spectrum Calculator, Meudon PDR ...

Signal must be defined by element size / beam FWHM or surface brightness Units (depend on instruments)

- Jy or W.m μm (flux density)
- Jy /arcsec (surface brightness)
 - W.m (flux density integrated on resolution unit)
 - T (brightness temperature)





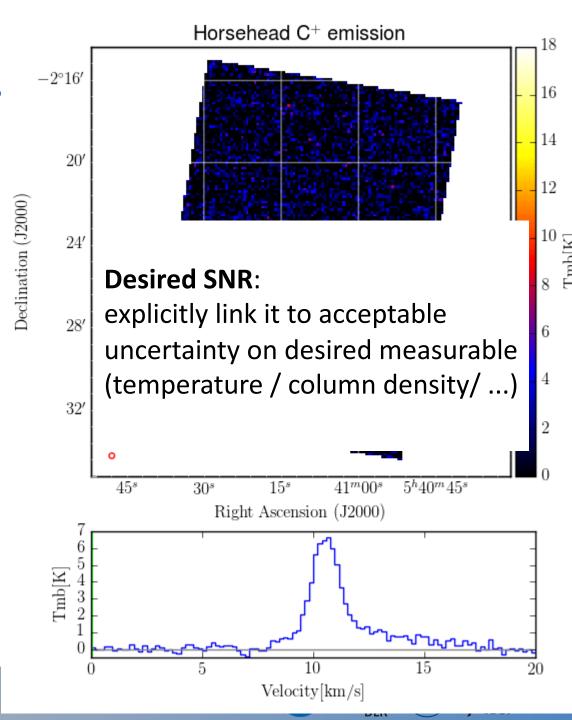
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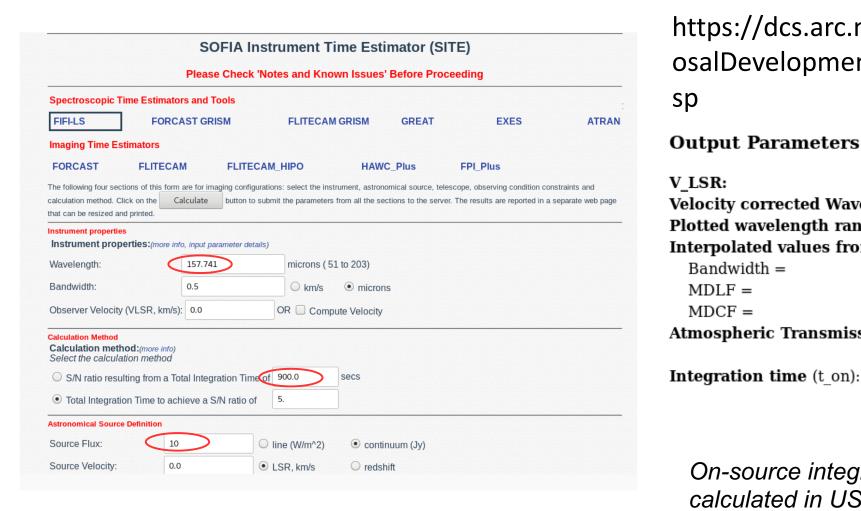
Signal must be defined by element size / beam FWHM or surface brightness Units (depend on instruments)

- Jy or W.m, μm (flux density)
- Jy /arcsec (surface brightness)
- W.m (flux density integrated on resolution unit)
- T (brightness temperature)





SITE: estimating corresponding observing time



https://dcs.arc.nasa.gov/prop osalDevelopment/SITE/index.j sp

Output Parameters

V LSR: 0.000 km/sVelocity corrected Wavelength: 157.741 microns Plotted wavelength range: 156.938 - 158.544 microns Interpolated values from data table: Bandwidth = 0.803 microns MDLF =2.087e-17 W/m^2 MDCF =0.570 Jy Atmospheric Transmission: 0.843 0.862 (smoothed) (unsmoothed)

On-source integration time only; overheads are calculated in USPOT



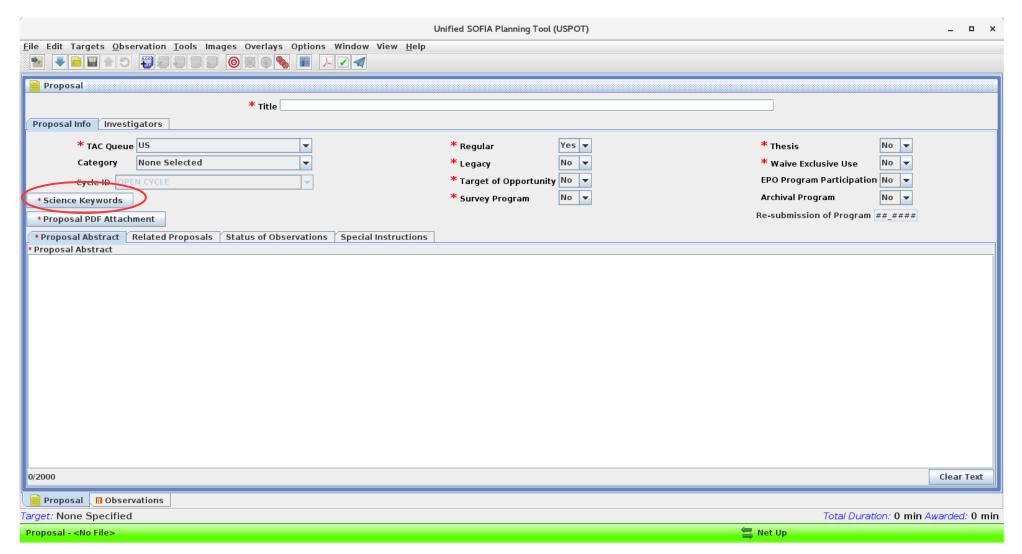


0.107

(smoothed) (unsmoothed)

0.102 minutes

USPOT: cover page







×

Close

Reset

ASTROMETRY

BAL QUASARS

BLACK HOLES

CALIBRATION

COMETS

DUST

DYNAMICS

EVOLUTION

FAST IMAGING

ABSORPTION LINES ACCRETION DISKS AGN PHYSICS ASTEROIDS

ATMOSPHERES AND CHI

BL LAC OBJECTS AND BL

CENTRAL STARS OF PLA

CHEMICAL ABUNDANCES

CLUSTER BINARY STARS

CLUSTERS OF GALAXIES

COSMOLOGICAL PARAM DAMPED LYMAN-ALPHA

COOLING FLOWS

DARK MATTER DETACHED BINARIES

DWARF GALAXIES

ECLIPSING BINARIES ELLIPTICAL GALAXIES EMISSION LINES

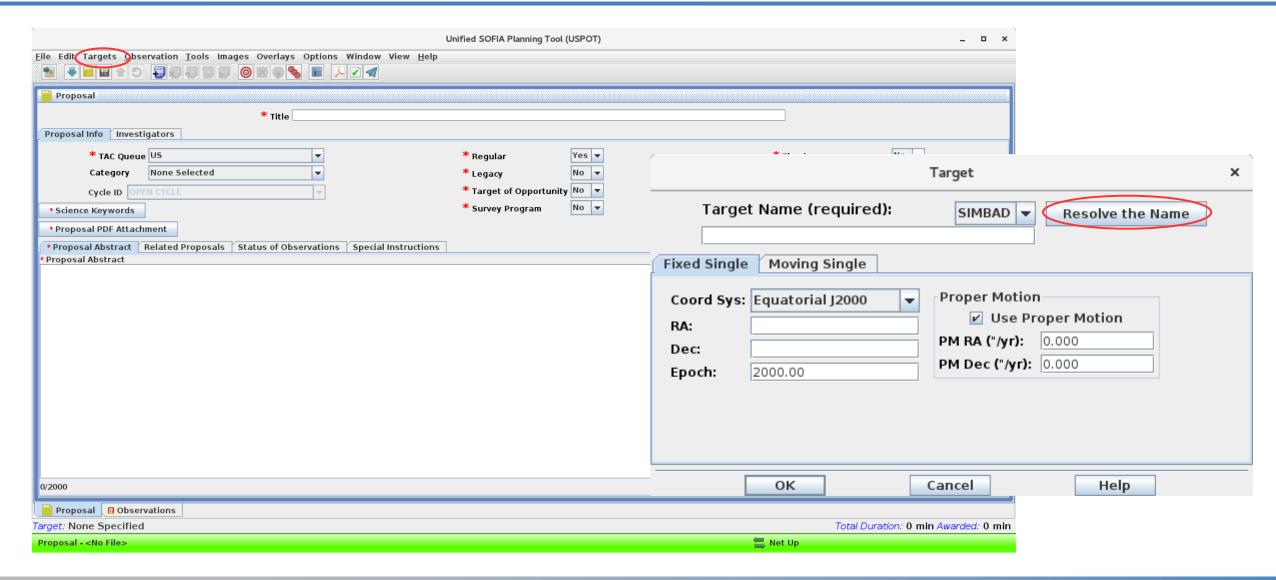
■ ERUPTIVE BINARY STARS

EXOSPHERIC ATMOSPHI

EXTRA-SOLAR PLANETS



USPOT: define targets







USPOT: define AORs FIFI-LS [AOR ID: N/A] Unique AOR Label: FIFI_LS-0000 Target: None Specified Modify Target Target List... New Target File Edit Target Observation Jools Images Overlays Options Window Observing Condition & Acquisition / Tracking Proposal On-source exp. time (sec) * Instrument Mode Symmetric Chop * Title Observation Order 30.000 * On source exp. time per cycle (sec) Proposal Info Investigators * Rest Wavelength Blue (micron) 63.184 Chop Type * Cycles * TAC Queue US ▾ * Width of Spectrum Blue (km/s OR micron) 0.000 Min Contiguous Exp Time (sec) 0.000 Total Chop Throw (arcsec) 120.000 Category None Selected Width of Spectral Feature Blue (km/s OR micron) 0.000 Grid * MapType Chop Angle Coordinate J2000 Cycle ID 157.741 * Rest Wavelength Red (micron) 90.000 * Science Keywords Chop Pos Angle (deg) ullet Number of Points Along Lat (Grid Only) $\, 1$ * Width of Spectrum Red (km/s OR micron) 0.000 * Proposal PDF Attachment ullet Number of Points Along Lon (Grid Only) 1Set Chop Angle Ranges Width of Spectral Feature Red (km/s OR micron) 0.000 *Proposal Abstract Related Proposals Status of Observations Specia Step Size Along Lat (arcsec) 30.000 Reference Position Width Unit Proposal Abstract Step Size Along Lon (arcsec) 30.000 Ref Type Source Velocity (km/s) 0.00000 Map Offset RA (arcsec) 0.000 By Offset Dichroic 105 micron ▼ Map Offset Dec (arcsec) 0.000 O By Position **Pointing Array** Blue ▼ Map Priority Map order ▼ Spectral 1 FIF BLUE Map Ref. Pos. FOV Angle (deg) 0.000 Spectral 2 FIF RED Reference Name Custom Map Area (arcsec^2) RA Offset (arcsec) Offset East/Row/Perpendicular (") Offset North/Column/Parrell (") Number Dec Offset (arcsec) 0.0 0.0 RA (deg) Dec (deg) Position: 0h00m00.0000s,+0d00m00.000s

Export Map Offsets

Apply

Export Map Positions

Cancel

Observation Est... | Comments... | Proposal Info...

*Import Map Offsets (Custom Only)

OK



Target: None Specified

Proposal - <No File>

Proposal B Observations

0/2000



Help

Choose Position

(** = Advanced) (* = required for Phase I)

USPOT: building justification pdf

- Context, aim, expected results (0.5p / 1p)
- Scientific justification (3p + references / 5 p)
- Feasibility + path to publication (3p) (instrument and modes, exposure time, time constraints)
- Budget (for Legacy only, 2p)
- Implementation (for Legacy only, 2p)
- Bio sketches
- Thesis enabling program (1 p)

Please note that for proposals longer than 9 (16) pages total, DCS will return a warning about a too long proposal document. If all individual section limits have been adhered to, this warning may be ignored.





USPOT: building justification pdf

Tips from Call from proposal:

- Preference given to substantial investigations that demonstrate significant scientific impact from SOFIA
- Programs using multi-wavelength data from major facilities (ALMA, HST, Spitzer, etc.) in conjunction with SOFIA are highly encouraged
- Programs that will inform future JWST observations are highly encouraged
- Criterion: degree to which the investigation uses SOFIA's Unique capabilities.
- Criterion: competence and relevant experience of the PI and any collaborators to carry the investigation to a successful conclusion.

General tips:

- do not forget instrument and data justification (as opposed to i.e., TEXES, archival data, ...)

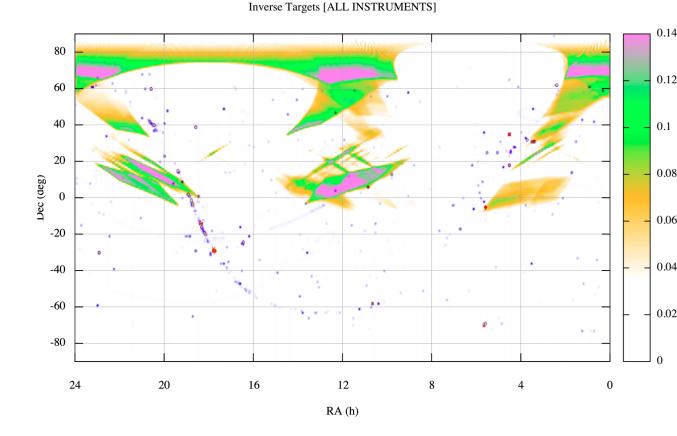




Other strategic considerations

Even though bulk of targets are concentrated in the GP and ecliptic, you can only spend 1/2 of the time flying to view those directions and the other 1/2 on the return

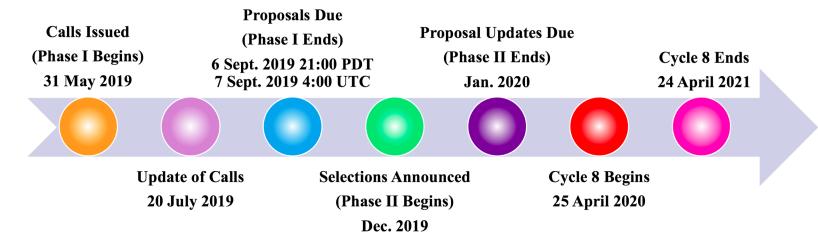
Whenever possible (surveys), choose targets with distribution at all parts of the sky. More chances to win! High latitudes are desirable.







Post deadline – what to expect



December 2019: results announced

Budget submission (US GOs)

Phase 2 – further definition of AORs, supported by instrument scientists

Observations: all GOs invited to their flights, receive flight summaries postflight series





Post deadline – what to expect

Processed data delivered and staged in archive ~ 1 month after observations

- GOs notified by email

Note that the SOFIA archive is moving to **IRSA** at the end of 2019! https://irsa.ipac.caltech.edu/Missions/sofia.html

Post-delivery: assistance available from science center / instrument scientists

Additional resources are available at the SOFIA website data section:

https://www.sofia.usra.edu/science/data - HelpDesk always open for questions:

sofia_help@sofia.usra.edu



