



# Data Processing for Observing Cycle 1

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**SOFIA Data Products** 

#### Defined in the Data Processing Plan for SOFIA SIs :

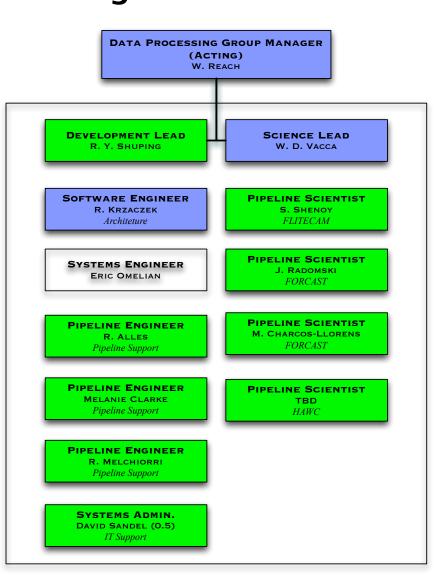
- Level 1: raw SI data in standardized format (FITS)
- Level 2: corrected for instrument artifacts (e.g., flats, darks, bad pixels)
- Level 3: flux calibrated (e.g. BSCALE/BUNIT keywords, MJy/sr)
- Level 4: high-order products possibly combining multiple observations

(e.g. mosaics, spectral cubes)





- Provide science-grade data products to GIs and science **community**;
  - process/analyze data;
  - quality assurance;
- Curate processed data in the archive and reprocess as needed.
- Develop pipelines and other needed software tools.
- Support SI teams on pipeline development/ delivery.











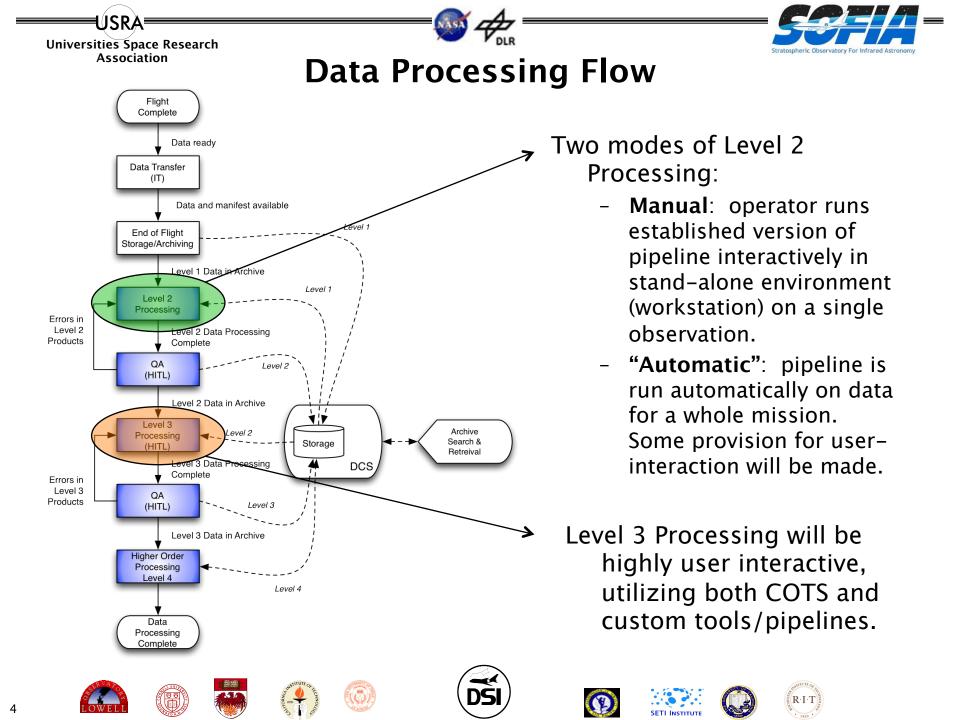


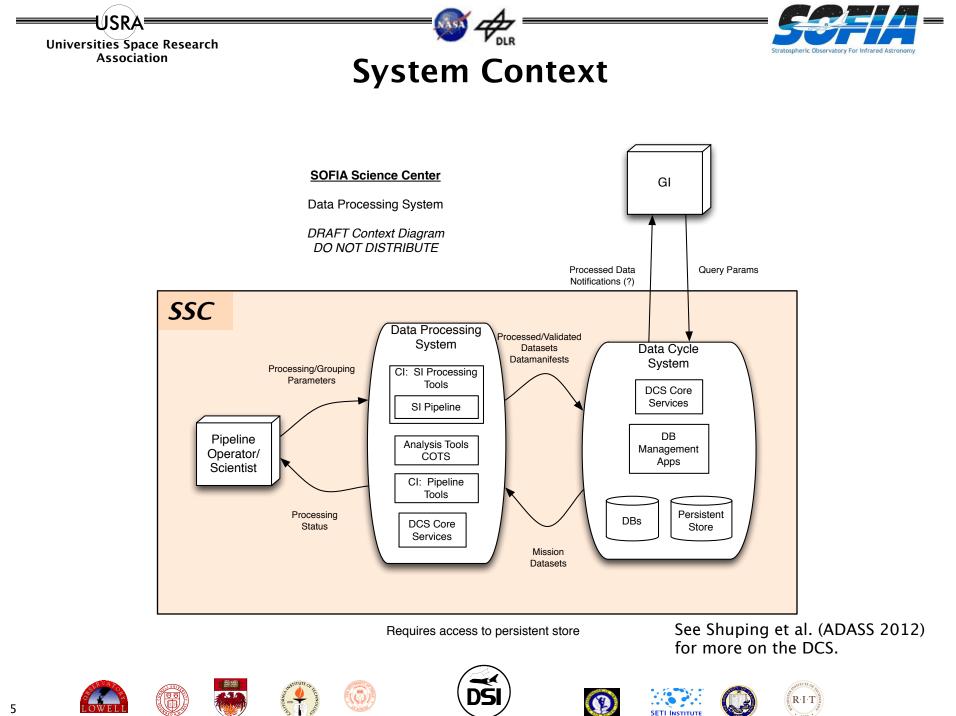












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## Data Processing System by instrument

- FORCAST imaging: Ready for operations
- FORCAST grisms: Operational; commissioning updates inwork
- FLITECAM imaging: Ready for operations
- FLITECAM grisms: Operational; waiting for additional wave cals.
- **GREAT**: Learning from Basic Science Level 3 deliveries (2013 Mar)
- FIFI-LS: Preliminary pipeline received; agreement with SI team to revisit
- **EXES**: Negotiating with SI team to obtain source code in summer
- **HAWC+**: HAWC pipeline installed and tested; tabled until HAWC+ ready
- HIPO: No need for reduced or calibrated products



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# **OC1 Level 2 Pipeline Processing**

- FORCAST (Imaging/Grism): DRIP/FG
  - Chop/Nod subtraction
  - Non-linear response correction
  - Bad-pixel removal
  - Droop & "jailbar" (cross-talk) correction
  - Optical distortion correction

     (I)
  - Field rotation/alignment (I)
  - Flatfield correction (G)
  - Optimal spectral extraction (G)
  - Wavelength calibration (G)

 FLITECAM (Imaging/Grism): FDRP/FSpextool

- Nod subtraction
- Flatfield correction
- Bad-pixel removal
- Optimal spectral extraction (G)
- Wavelength Calibration (G)(TBD)











#### Universities Space Research Associat Flux Calibration for FORCAST and FLITECAM

• On each flight, observations of standard stars will be obtained at each altitude (as in Basic Science).

- See backup slides for standard star selection

- For imaging, fluxes and wavelengths have been derived for each standard star for each filter using a comprehensive model of the instrument throughput and atmospheric transmission
- Corrections for differences in airmass, altitude, and pwv, between targets and standards have been derived from ATRAN models for each passband and incorporated into the calibration software
- Calibration parameters applied to Level 2 data to produce Level 3 products; calibration params also stored in archive for reference.













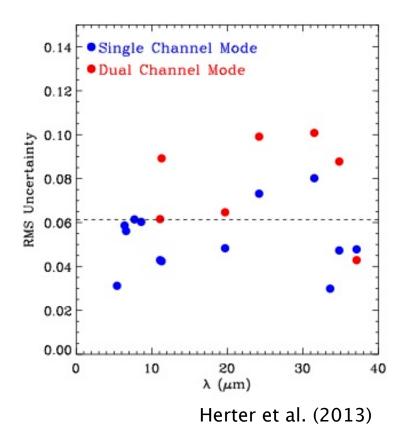


## **FORCAST Basic Science Flux Calibration Results**

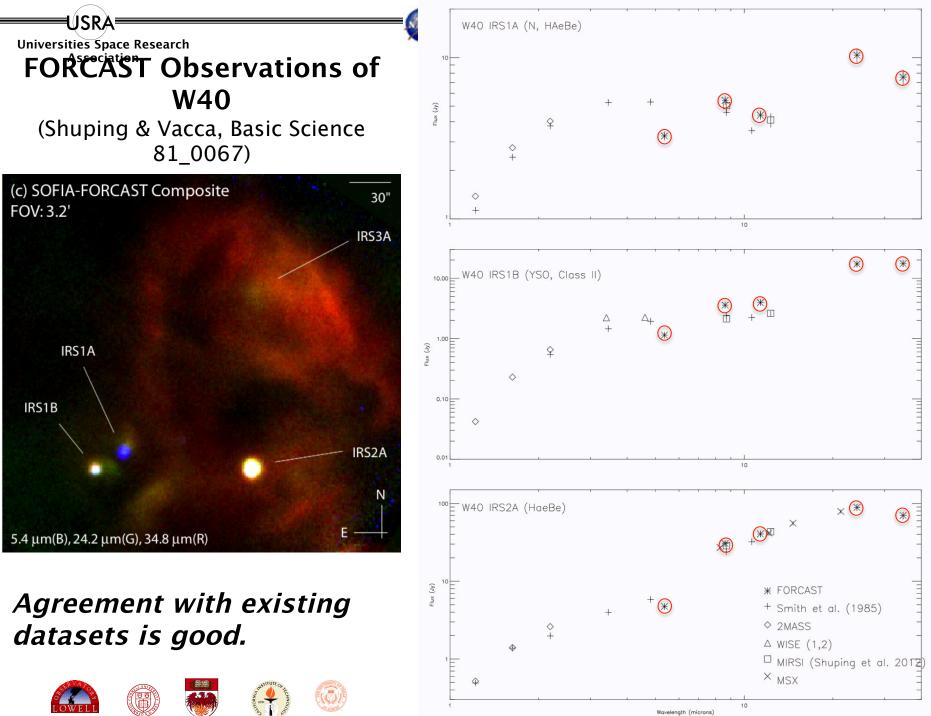
 Method worked well for calibrating FORCAST Basic Science data (Herter et al. 2013, in press)

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- RMS Response Uncertainty: ~6.1% (flight-by-flight)
- Anecdotal results for a few Basic Science targets indicate that the calibration is good to <20% (3 sigma)</li>
- Currently analyzing standards from Basic Science to evaluate overall accuracy.









# Data Distribution Timeline

- Level 2 data available to GI within 2 weeks of flight series completion
  - OC1B Exception: delivery within 4 weeks
- Level 3 (Flux Calibrated) data available to GI within 4 *weeks* of flight series completion
  - OC1B Exception: delivery within 6 weeks
- Data distribution via SOFIA Archive
  - GI data released to public after 1-year (from archive insertion).
  - Standards and calibration parameters released immediately.











## System Development Status

- Requirements and design reviews complete
- Development hardware/network ready and in-use
- Production hardware/network almost ready, awaiting formal test
  - Dev environment serves as backup
- V&V with commissioning data underway
- Development of standard operating procedures underway
- Formal system testing coming up in prep for Cycle 1













